

Recent Results on Direct Reactions with Stored Radioactive Beams and with Active Targets



#### Peter Egelhof GSI Darmstadt, Germany for the EXL Collaboration

9<sup>th</sup> Int. Conference on Direct Reactions with Exotic Beams DREB 2016

> Halifax, Canada July 11 – 15, 2016



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- I. Introduction
- II. Direct Reactions with RIB's at Storage Rings A new Approach for Low Momentum Transfer Measurements
- III. First Experiments and Feasibility Studies at the ESR Storage Ring
  - a) Elastic Proton Scattering on  ${}^{56}Ni \Rightarrow$  Nuclear Matter Distribution
  - b) Inelastic Alpha Scattering on  ${\rm ^{58}Ni} \Rightarrow$  Giant Monopole Resonance
- **IV.** Future Perspectives
- V. Conclusions

#### I. Introduction: Direct Reactions with Radioactive Beams in Inverse Kinematics

#### classical method of nuclear spectroscopy:

- $\Rightarrow$  light ion induced direct reactions: (p,p), (p,p'), (d,p), ...
- $\Rightarrow$  to investigate exotic nuclei: inverse kinematics
- $\Rightarrow$  important information at low momentum transfer!

#### of particular interest:

- $\Rightarrow$  radial shape of nuclei: skin, halo structures
- $\Rightarrow$  doubly magic nuclei: <sup>56</sup>Ni, <sup>132</sup>Ni
- $\Rightarrow$  giant resonances: nuclear compressibility

#### future perspectives at FAIR:

- $\Rightarrow$  profit from intensity upgrade (up to 10<sup>4</sup> !!)
- $\Rightarrow$  explore new regions of the chart of nuclides and new phenomena
- $\Rightarrow$  use new and powerful methods:
- EXL: direct reactions at internal storage ring target
  - ⇒ high luminosity even for very low momentum transfer measurements

First Experiments at the ESR



Nuclear Physics with Radioactive Beams at FAIR: NUSTAR: NUclear STructure, Astrophysics and Reactions

#### **I** High intensity primary beams from SIS 100 (e.g. $10^{12} \, {}^{238}\text{U}$ / sec at 1 GeV/u)





### The EXL Project: EXotic Nuclei Studied in Light-Ion Induced Reactions at the NESR Storage Ring



# Light-Ion Induced Direct Reactions at Low Momentum Transfer

- elastic scattering (p,p), (α,α), ...
  nuclear matter distribution ρ (r), skins, halo structures
- inelastic scattering (p,p'), (α,α'), ...
  deformation parameters, B(E2) values, transition densities, giant resonances
- transfer reactions (p,d), (p,t), (p, <sup>3</sup>He), (d,p), ... single particle structure, spectroscopic factors, spectroscopy beyond the driplines, neutron pair correlations, neutron (proton) capture cross sections
- charge exchange reactions (p,n), (<sup>3</sup>He,t), (d, <sup>2</sup>He), ...
  Gamow-Teller strength
- knock-out reactions (p,2p), (p,pn), (p,p <sup>4</sup>He)...
  ground state configurations, nucleon momentum distributions

#### for almost all cases:

region of low momentum transfer contains most important information

#### Speciality of EXL:

measurements at very low momentum transfer

 $\Rightarrow$  complementary to R<sup>3</sup>B !!!

Experiments to be Performed at Very Low Momentum Transfer – Some Selected Examples

- Investigation of Nuclear Matter Distributions:
  - $\Rightarrow$  halo, skin structure
  - $\Rightarrow$  probe in-medium interactions at extreme isospin (almost pure neutron matter)
  - $\Rightarrow$  in combination with electron scattering (ELISe project @ FAIR ):

separate neutron/proton content of nuclear matter (deduce neutron skins )

method: elastic proton scattering  $\Rightarrow$  <u>at low q</u>; high sensitivity to nuclear periphery

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- Investigation of the Giant Monopole Resonance:
  - $\Rightarrow$  gives access to nuclear compressibility  $\Rightarrow$  key parameters of the EOS
  - $\Rightarrow$  new collective modes (breathing mode of neutron skin)

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- Investigation of Gamow-Teller Transitions:
  - $\Rightarrow$  weak interaction rates for N = Z waiting point nuclei in the rp-process

 $\Rightarrow$  electron capture rates in the presupernova evolution (core collaps) method: (<sup>3</sup>He,t), (d,<sup>2</sup>He) charge exchange reactions <u>at low q</u>

#### Kinematical Conditions for Light-Ion Induced Direct Reactions in Inverse Kinematics



- required beam energies: E ≈ 200 ... 740 MeV/u (except for transfer reactions)
- required targets: <sup>1,2</sup>H, <sup>3,4</sup>He
- most important information in region of low momentum transfer
  - $\Rightarrow$  low recoil energies of recoil particles
  - $\Rightarrow$  need thin targets for sufficient angular and energy resolution

#### Advantage of Storage Rings for Direct Reactions in Inverse Kinematics

- low threshold and high resolution due to: beam cooling, thin target (10<sup>14</sup>-10<sup>15</sup> cm<sup>-2</sup>)
- gain of luminosity due to: continuous beam accumulation and recirculation
- low background due to: pure, windowless <sup>1,2</sup>H<sub>2</sub>, <sup>3,4</sup>He, etc. targets
- experiments with isomeric beams

Experiments at very low momentum transfer can only be performed at EXL (except with active targets, but with substantial lower luminosity)

# III. First Experiments with RIB's and Feasibility Studies at the ESR Storage Ring

# specially designed scattering chamber for the ESR:





#### reactions with <sup>58</sup>Ni:

#### proof of principles and feasibility studies:

- UHV capability of detector setup
- background conditions in ESR environment at the internal target
- Iow energy threshold
- beam and target performance

#### reactions with <sup>56</sup>Ni:

#### <sup>56</sup>Ni: doubly magic nucleus!!

- (p,p) reactions: nuclear matter distribution
- (α,α`) reactions: giant resonances (GMR) EOS parameters (nucl. compressibility)
- (<sup>3</sup>He,t) reactions: Gamow-Teller matrix elements, important for astrophys.

#### **Theorectical Predictions**



needed: large solid angle detectors with low threshold and large dynamic range

#### Setup at the ESR Storage Ring





#### **Experimental Concept**



# Experimental Concept



## **Experimental Concept**

Auxilliary vacuum side



Ultra-high vacuum side



#### Experimental Setup at the ESR



#### Scattering Chamber mounted at the Internal Target of the ESR

challenge: UHV capable and bakeable DSSD and Si(Li) detectors





#### FRS: In-Flight Separator & High-Resolution Spectrometer



Preparation of the Stored Radioactive <sup>56</sup>Ni Beam

beam after

fragmentation of 600 MeV/u <sup>58</sup>Ni beam FRS:

injection to ESR: 7 x 10<sup>4</sup> <sup>56</sup> Ni per injection

stochastic cooling, bunching and stacking (60 injections):

**4.8 x 10<sup>6</sup>** <sup>56</sup>Ni in the ring rf deposition injected beam 50 vacuum chamber recorded DC current transformer data linear fit to the data 40 30 [1] electron cooled beam after 20 intense stack stochastic precooling 10 target profile injections not recorded 0 40 35 30 25 20 15 500 1000 1500 2000 2500 t [s]  $H_2$  target: 2 x 10<sup>13</sup> cm<sup>-2</sup> luminosity: 10  $\Rightarrow$  **L = 2 x 10<sup>26</sup> cm<sup>-2</sup> sec**<sup>-1</sup> F. Nolden et al., 5 (reduced by aperture) 0 Proc. IPAC2013 -2 -8 -6 0 2 4 Beam position [mm] **MOPEA013**  $\sigma = 3.78 \text{ mm}$   $x_0 = 0.58 \text{ mm}$ 

#### First Nuclear Reaction Experiment with Stored Radioactive Beam!!!!

M. von Schmid et al., Phys. Scr. T166 (2015) 014005

P. Egelhof et al., JPS Conf. Proc. 6 (2015) 020049



<sup>56</sup>Ni(p,p), E = 390 MeV/u Reconstructed Energy





<sup>56</sup>Ni(p,p), E = 390 MeV/u Benefit of the 1mm Aperture



<sup>56</sup>Ni(p,p`), E = 390 MeV/u Identification of Inelastic Scattering



<sup>56</sup>Ni(p,p), E = 390 MeV/u Angular Distribution



M. v. Schmid, PHD thesis 2015

#### Concept of the Data Analysis

- Glauber multiple-scattering theory for calculation of cross sections:
  - use measured free pp, pn-cross sections as input (in medium effects negligible)
  - fold with nucleon density distribution
  - take into account multiple scattering (all terms!) (small for nuclear periphery)
- variation of the nuclear matter density distribution:
  - a) phenomenological parametrizations (point matter densities):
    - SF: Symmetrized Fermi
  - b) "model independent" analysis:

SOG: Sum Of Gaussians (standard method for electron scattering data: I. Sick, Nucl. Phys. A 218 (1974) 509) Nuclear Matter Density Distribution of <sup>56</sup>Ni from Elastic Proton Scattering

## <sup>56</sup>Ni(p,p), E = 390 MeV/u Angular Distribution Cross Section fitted using the Glauber Theory



M. v. Schmid, PHD thesis 2015

Nuclear Matter Density Distribution of <sup>56</sup>Ni from Elastic Proton Scattering

# Nuclear Matter Distribution of <sup>56</sup>Ni Cross Section fitted using the Glauber Theory



Nuclear Matter Density Distribution of <sup>56</sup>Ni from Elastic Proton Scattering

## Nuclear Matter Distribution of <sup>56</sup>Ni Cross Section fitted using the Glauber Theory







 ${}^{58}Ni(\alpha, \alpha)$ , E = 100 MeV/u



challenge: detect and identify very low energy recoils



J. C. Zamora, PHD thesis 2015



#### **IV. Future Perspectives**

#### short term perspectives:

- $(\alpha, \alpha)$  on <sup>56</sup>Ni  $\Rightarrow$  investigate ISGMR, compressibility of nuclear matter
- (<sup>3</sup>He,t) and (d,<sup>2</sup>He) on <sup>56</sup>Ni  $\Rightarrow$  investigate Gamow Teller strength
- (p,p) on heavier Ni and Sn isotopes  $\Rightarrow$  nuclear matter distributions, skins
- transfer reactions at Cryring (GSI) and TSR@ISOLDE (CERN)

#### upgrade of detector setup and readout:



# Future Perspectives

#### long term perspectives (EXL @ FAIR):

- for first phase of FAIR:
  - $\Rightarrow$  install EXL at the HESR
  - $\Rightarrow$  and/or install transfer line from SUPER-FRS / CR to the ESR



#### The E105 Collaboration



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# V. Conclusions

- For the First Time (World Wide) a Nuclear Reaction Experiment with Stored Radioactive Beams was successfully performed.
- A "Proof of Principle" of the Experimental Concept with UHV compatible Detectors and Infrastructure around the Internal Target was successfull.
- The Nuclear Matter Density Distribution of <sup>56</sup>Ni was determined with High Accuracy.
- A Feasibility Study for Investigations of the ISGMR was performed, the Cross Section was measured down to  $\Theta_{cm} < 1$  deg.
- EXL@ESR and EXL@FAIR has a large Potential for Nuclear Structure and Nuclear Astrophysics.