# Particle Physics Parallel Session Summary

# **& TRIUMF** Science Week 2020

Patrick de Perio & David McKeen

Talks by <u>Daniel Stolarski</u>, <u>Joseph Formaggio</u>, & <u>Masha Baryakhtar</u>

### **Particle Physics: A Roadmap for Discoveries**

Daniel Stolarski **Discovery Opportunities** at Future Colliders

Masha Baryakhtar Discoveries at the **Precision Frontier:** Axions and New **Ultralight Particles** 

Joseph Formaggio Trying to weigh the lightest particles in the universe without losing your patience

Tim M.P. Tait

### The Future Gets Brighter?

- With the discovery of the Higgs boson, the Standard Model of Particle Physics has been established as a complete theory that could in principle describe physics up to very high energies.
- Still, many questions remain:
  - The nature of dark matter and dark energy
  - The origin of the baryon asymmetry of the Universe.
  - Flavor and neutrino masses.
- Future experiments, including the high luminosity Large Hadron Collider, future observatories, and searches for dark matter offer the opportunity to shed light on these mysteries.
- The next few years offer the opportunity for great discoveries!





SM says Higgs breaks electroweak symmetry with this potential.

We've discovered the Higgs boson but we do not know much about its potential

**Daniel Stolarski** 

# HGGS POTENTIAL

No direct experimental evidence of this.

Can measure derivatives of potential.



Taylor series:

 $V(h) \sim \frac{1}{2} m_h^2 h^2 + \frac{1}{3!} \lambda_3 h^3 + \frac{1}{4!} \lambda_4 h^4 + \dots$ 

![](_page_2_Picture_16.jpeg)

In the early universe, electroweak symmetry is restored.

We've discovered the Higgs boson but we do not know much about its potential

Shape of the Higgs potential can tell us about why we are here

**Daniel Stolarski** 

### **ELECTROWEAK PHASE TRANSITION (hh)**

![](_page_3_Figure_9.jpeg)

BSM theories (with new states) could have violent transition, possible baryogengesis mechanism.

Curtin, Meade, Yu, arXiv:1409.0005. See also talk by T. Tait on Wednesday.

![](_page_3_Picture_14.jpeg)

SM makes definite predictions for these coefficients:

### To map out the Higgs potential we need new colliders

Can directly measure these couplings with multi-Higgs production (very hard at LHC).

**Daniel Stolarski** 

## N-HIGGS PRODUCTION (hh)

![](_page_4_Figure_7.jpeg)

![](_page_4_Figure_8.jpeg)

![](_page_4_Picture_11.jpeg)

In lepton collider, can use knowledge of initial state to detect that a Higgs was created without seeing it.

Search for Higgs decays to new particles.

Future colliders can also tell us about new states (e.g. DM) that are impossible to see at the LHC **Daniel Stolarski** 

# NEW LIGHT PARTICLES (ee/he?)

![](_page_5_Picture_9.jpeg)

Can also look for new electroweakly charged particles with difficult decays.

![](_page_5_Picture_11.jpeg)

Could be connected to dark matter or SUSY.

![](_page_5_Picture_15.jpeg)

If muon g-2 remains discrepant from SM value, future muon collider has a "no lose theorem"

**Daniel Stolarski** 

## MUON COLLIDERS

### arXiv:2006.16277

### A Guaranteed Discovery at Future Muon Colliders

Rodolfo Capdevilla<sup>a,b</sup>,\* David Curtin<sup>a</sup>,† Yonatan Kahn<sup>c</sup>,‡ and Gordan Krnjaic<sup>d</sup>§ <sup>a</sup>Department of Physics, University of Toronto, Canada <sup>b</sup>Perimeter Institute for Theoretical Physics, Waterloo, Ontario, Canada <sup>c</sup> University of Illinois at Urbana-Champaign, Urbana, IL USA and <sup>d</sup>Fermi National Accelerator Laboratory, Batavia, IL USA (Dated: July 1, 2020)

The longstanding muon g-2 anomaly may indicate the existence of new particles that couple to muons, which could either be light ( $\leq \text{GeV}$ ) and weakly coupled, or heavy ( $\gg 100 \text{ GeV}$ ) with large couplings. If light new states are responsible, upcoming intensity frontier experiments will discover further evidence of new physics. However, if heavy particles are responsible, many candidates are beyond the reach of existing colliders. We show that, if the  $(g-2)_{\mu}$  anomaly is confirmed and no explanation is found at low-energy experiments, a high-energy muon collider program is guaranteed to make fundamental discoveries about our universe. New physics scenarios that account for the

![](_page_6_Figure_9.jpeg)

![](_page_6_Picture_12.jpeg)

A myriad of experiments demonstrated that neutrinos transmute flavor (oscillate).

Proof that neutrinos must have mass.

There are predictions that stem from alteration of the Standard Model.

![](_page_7_Picture_4.jpeg)

### Joseph Formaggio

![](_page_7_Picture_7.jpeg)

![](_page_8_Figure_1.jpeg)

Cosmology (@ 50 meV) and  $0v\beta\beta$  (@ 15 meV) has the potential to probe the deepest into the oscillation prediction for the mass scale over the next decade.

However, the method with the most strongly tested assumptions is direct kinematic searches through beta decay.

### Joseph Formaggio

unknown.

experiments.

![](_page_8_Picture_12.jpeg)

Necessary [Experimental] Conditions: **High Flux and High Precision** 

![](_page_9_Figure_3.jpeg)

![](_page_10_Figure_1.jpeg)

### Joseph Formaggio

Electron transfers all of its energy to the absorbing medium.

> Calorimetric (Cryogenic Bolometers)

### ion spectroscopy of T<sub>2</sub>

, Phys. Rev D 80:051301 (2009)

$$\underbrace{1}_{4\times 10^{-8} \text{ for } \Delta E_{\text{kin}}} \frac{1}{m_{\text{e}}} \underbrace{\frac{eB}{2m_{\text{e}}}}_{2} \left(1 - \frac{E_{\text{kin}}}{m_{\text{e}}c^{2}} + \underbrace{\left(\frac{E_{\text{kin}}}{m_{\text{e}}c^{2}}\right)^{2}}_{4\times 10^{-8} \text{ for } \Delta E_{\text{kin}} = 100 \text{ eV}}\right)$$

$$B^2 \left( E_{\mathrm{kin}}^2 + 2 E_{\mathrm{kin}} m c^2 \right) \sin^2 \theta$$

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![](_page_10_Picture_12.jpeg)

eous emission from field.

### icy-Based Emission Spectroscopy)

![](_page_10_Picture_15.jpeg)

![](_page_11_Figure_1.jpeg)

Calorimetric experiments such as ECHO and HOLMES are progressing well toward the eV scale.

> **KATRIN** is now taking data, finally pushing the mass scale limit below the eV scale for the first time.

The CRES technique through Project 8 is pusing forward, with the eventual target of using an atomic tritium source.

![](_page_11_Picture_5.jpeg)

![](_page_11_Picture_6.jpeg)

10<sup>-22</sup> eV

DM candidates can span many orders of magnitude in mass-lots of work moving beyond WIMP (10 GeV-TeV) regime

### Theorists Searching for New Physics Sikivie: Experimental Tests of the "Invisible" Axion (1983) Goodman, Witten Detectability of Certain Dark Krauss, Moody, Wilczek, Morris: Calculations for Cosmic Matter Candidates (1984) Axion Detection (1985) ADMX-G2 1 peV QCD axion 1 GeV WIMP 10<sup>19</sup> GeV 1 eV \* **s**ee T.T.Yu talk see A.Arvanitaki, 1. Directional detection of WIMPs M. Safronova talks dran, Zobrist, Sushkov, Walsworth, Lukin with NV center spectroscopy 2. Quantum sim/optimization for WIMP nuclear-recoil response 3. LHe and GaAs for calorimetry for Knapen, Lin, Pyle, Zurek hidden-sector DM **MB**, Huang, Lasenby 4. Nanowire single-photon detectors for hidden photons Arvanitaki, Dimopoulos, Van Tilburg **5**. Qubits and Rydberg atoms for QCD axions J 6. Qubits and ultrahigh-Q cavities for hidden photons Graham, Mardon, Rajendran, Zhao 7. Photon upconverters for QCD axions (needed for nuclear spins and lumped resonators) 8. Lumped resonators for QCD axions Kahn, Safdi, Thaler Chaudhuri, Graham, Irwin, Mardon, Rajendran, Zhao 9. Nuclear spins for QCD axions Budker, Graham, Ledbetter, Rajendran, Sushkov 10. Clocks / cold atoms for scalar dark matter Arvanitaki, Huang, Van Tilburg Arvanitaki, Graham, Hogan, Rajendran, Van Tilburg Adapted from K. Irwin

### See also talks by A. Arvanitaki, M. Safranova, T. T. Yu

![](_page_12_Picture_7.jpeg)

Excellent light (<eV) dark matter candidates include axion(-like particles) and dark photons

![](_page_13_Picture_2.jpeg)

Masha Baryakhtar

### Axions and New Ultralight Particles

- Axions are
  - Solutions to a theoretical puzzle of small numbers: the strong-CP problem: approximately massless particle with mass and couplings fixed by a high scale *f<sub>a</sub>*,

$$m_a = 5.70(6)(4) \,\mu \text{eV} \left(\frac{10^{12} \text{GeV}}{f_a}\right)$$

• Low-energy remnants of complex physics at high scales Arvanitaki, Dimopoulos, Dubovsky, Kaloper, March-Russell

• Automatically produced as dark matter in the early universe

Nelson, Scholtz Arias, Cadamuro, Goodsell, Jaeckel, Redondo, Ringwald Graham, Mardon, Rajendran

Preskill, Wise, Wilczek

![](_page_13_Picture_14.jpeg)

These bosons naturally interact with photons—a handle to detect them with!

![](_page_14_Figure_2.jpeg)

### Searches for Axions and Dark Photons

- Wide parameter space of weakly coupled, light particles
- Holdom (1986)

• Axions and dark photons generically couple to photons: opens new search strategies with recent technological advances

![](_page_14_Figure_8.jpeg)

Kim (1979); Shifman, Vainshtein, Zakharov (1980) Dine, Fischler, Srednicki (1981); Zhitnitsky, (1980)

![](_page_14_Figure_10.jpeg)

![](_page_14_Picture_12.jpeg)

Kinematics of converting light (but not massless) boson DM to photons means we need to change photon dispersion relation

![](_page_15_Figure_2.jpeg)

for an e particle: Masha Baryakhtar

### Seeing Dark Matter

For a given energy, photons have much more momentum than dark matter

 $\omega_A = \omega_{\sim} \Rightarrow k_{\sim} \sim 0$ 

Even when *interactions* in the dark matter model allow one-to-one conversion to photons, *kinematics* do not

photons  

$$\frac{10^{-15} \text{sec}}{10^{-15} \text{sec}}$$
 mm  
dark photon  
dark matter

Need systems which can efficiently absorb the momentum mismatch

![](_page_15_Picture_11.jpeg)

New idea: stack of different dielectrics. DM axions or dark photons convert to photons that are focused onto detector

![](_page_16_Picture_2.jpeg)

### Nanowire Detection of Photons from the Dark Side

• Small area single photon detector with ultra low noise MB, J. Huang, R. Lasenby PRD 2018

![](_page_16_Picture_7.jpeg)

- Jeff Chiles, NIST
- Signal photons perpendicular to stack: efficiently focused

- High index of refraction contrast, more layers increase conversion
- e.g. silicon (n<sub>2</sub>=3.4) and silica  $(n_1 = 1.46)$

DOE QuantiSED grant, DE-SC0019129 (\$300,000 for two years)

Bosonic Dark Matter Search Using Superconducting Nanowire Single-Photon Detectors.

(Exp) Berggren, Charaev; Chiles, Nam; (Th) Arvanitaki, **MB,** Huang, Lasenby, Van Tilburg.

![](_page_16_Picture_16.jpeg)

New idea: stack of different dielectrics. DM axions or dark photons convert to photons that are focused onto detector

![](_page_17_Figure_2.jpeg)

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### Searches for dark matter with light

Dielectric materials /crystal structures/ molecules can correct the dispersion mismatch in waves between a massless and massive particle of the same energy

First steps underway, use well-established optics and detector technology; possible to reach very small couplings with larger setups

• Improve on parameter space by orders of magnitude, and perhaps see dark matter

![](_page_17_Picture_13.jpeg)

Summary<sup>2</sup> Particle physics has made incredible progress Questions remain! Neutrino masses, DM, matterantimatter asymmetry of the Universe... Many avenues for progress across wide range of energy scales (<eV to 100 TeV) Discussed just a few possibilities—next 20+ years will be exciting