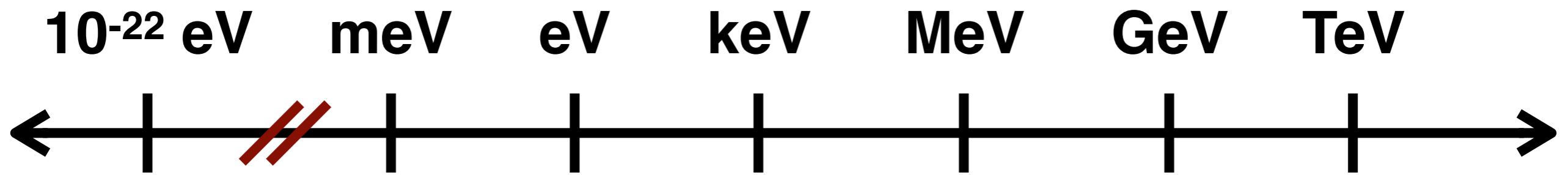


Finding Dark Matter in the Lab

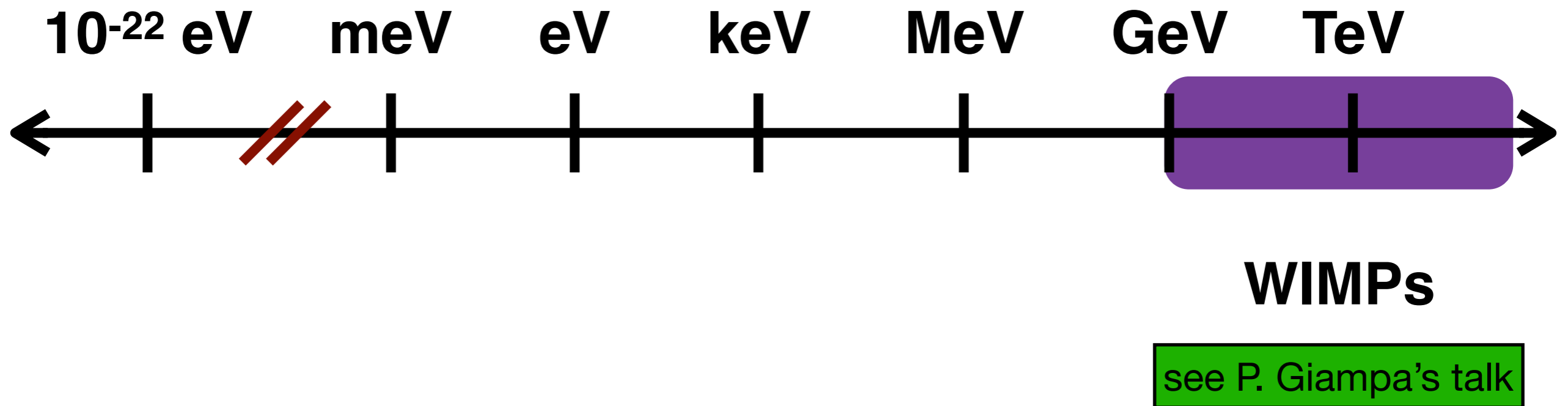
Tien-Tien Yu
(University of Oregon)

TRIUMF Science Week – August 20, 2020

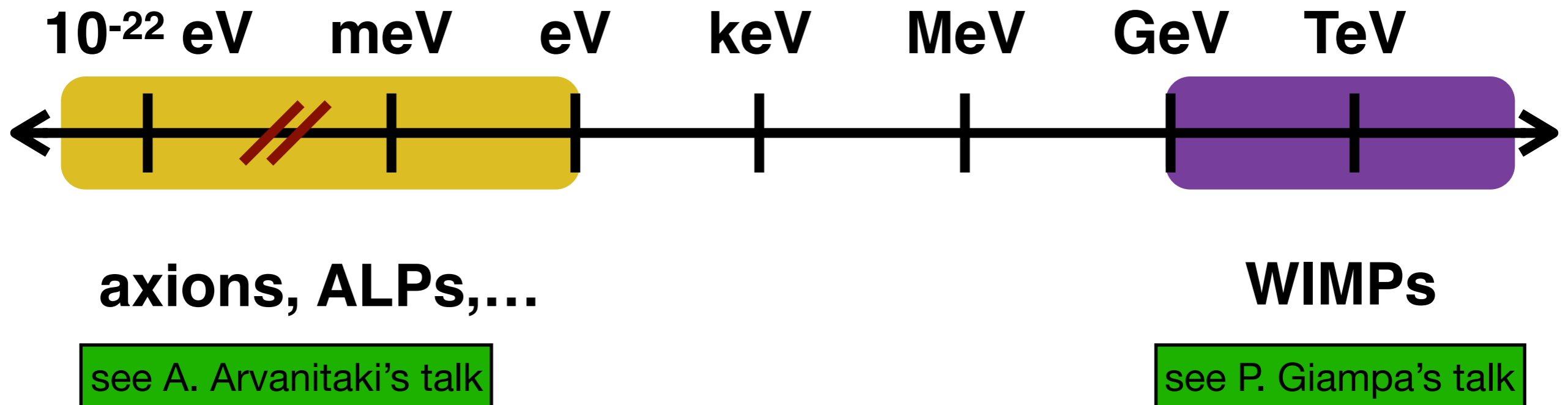
dark matter candidates



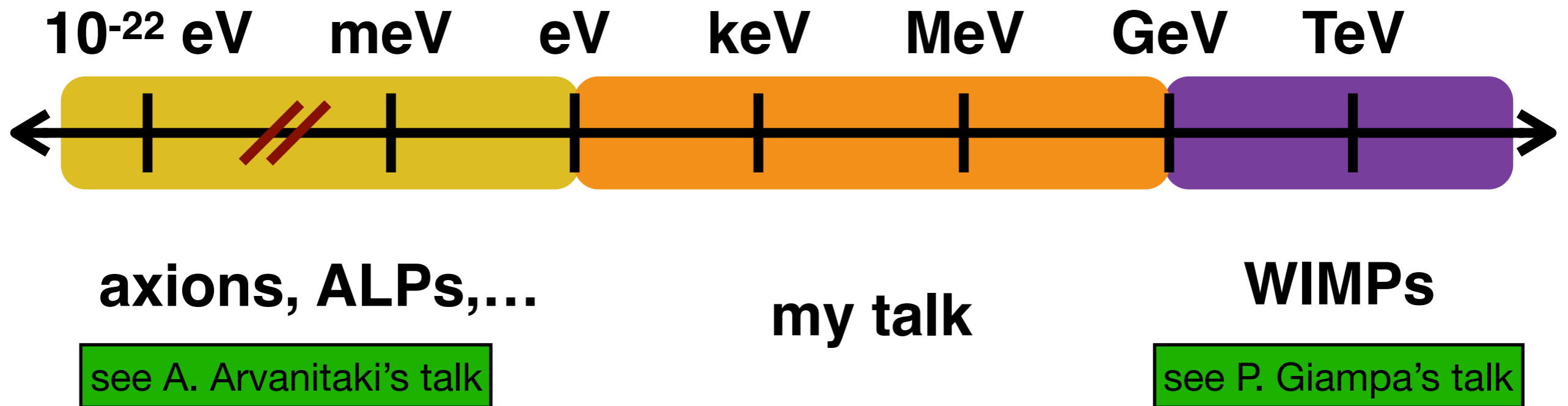
dark matter candidates



dark matter candidates



dark matter candidates



Ways to find DM in the lab

- **produce** DM at accelerators or colliders
- **detect** galactic DM with underground detectors
 - look for scattering of DM with target material
 - look for absorption of DM within target material

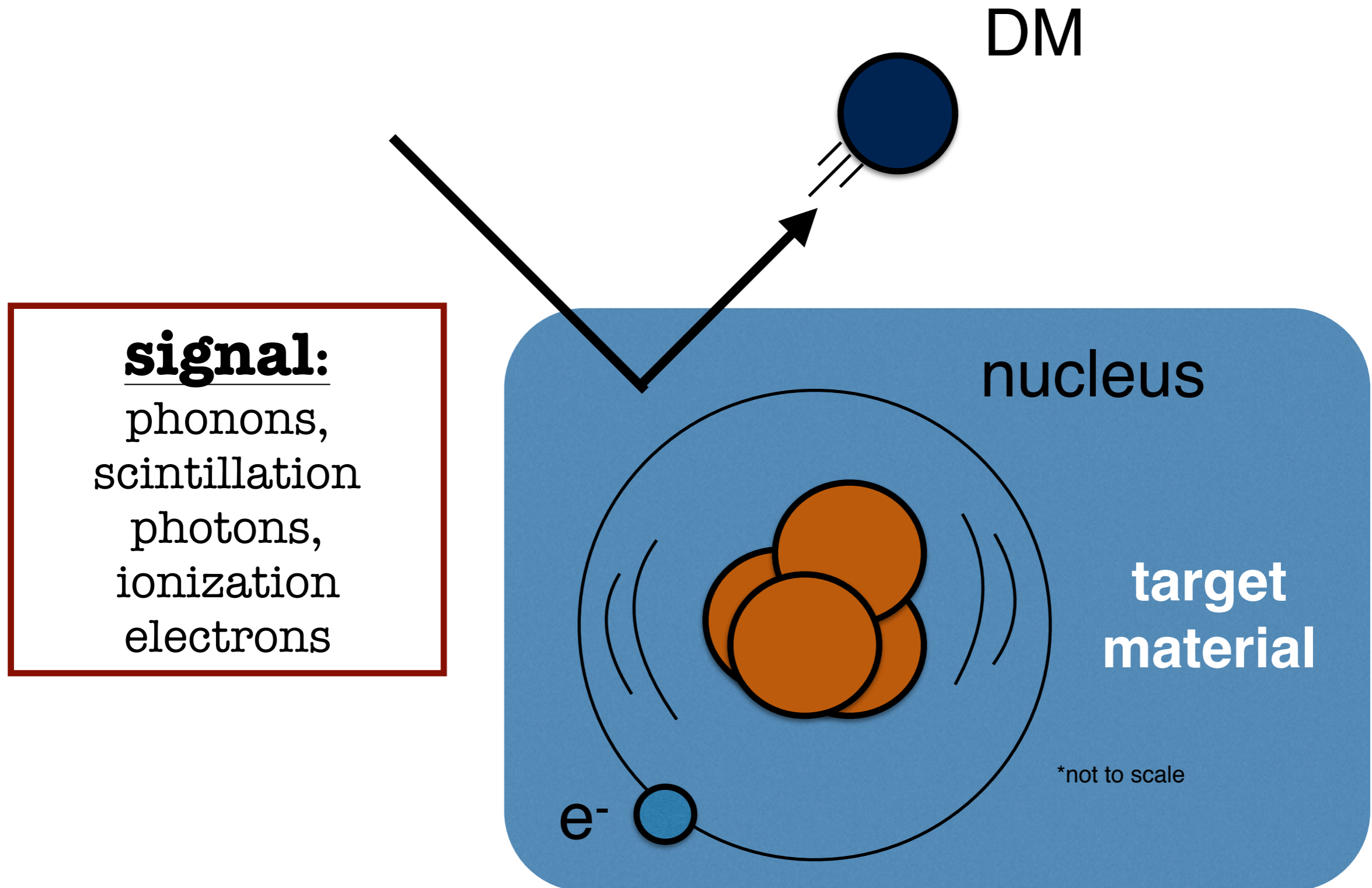
Ways to find DM in the lab

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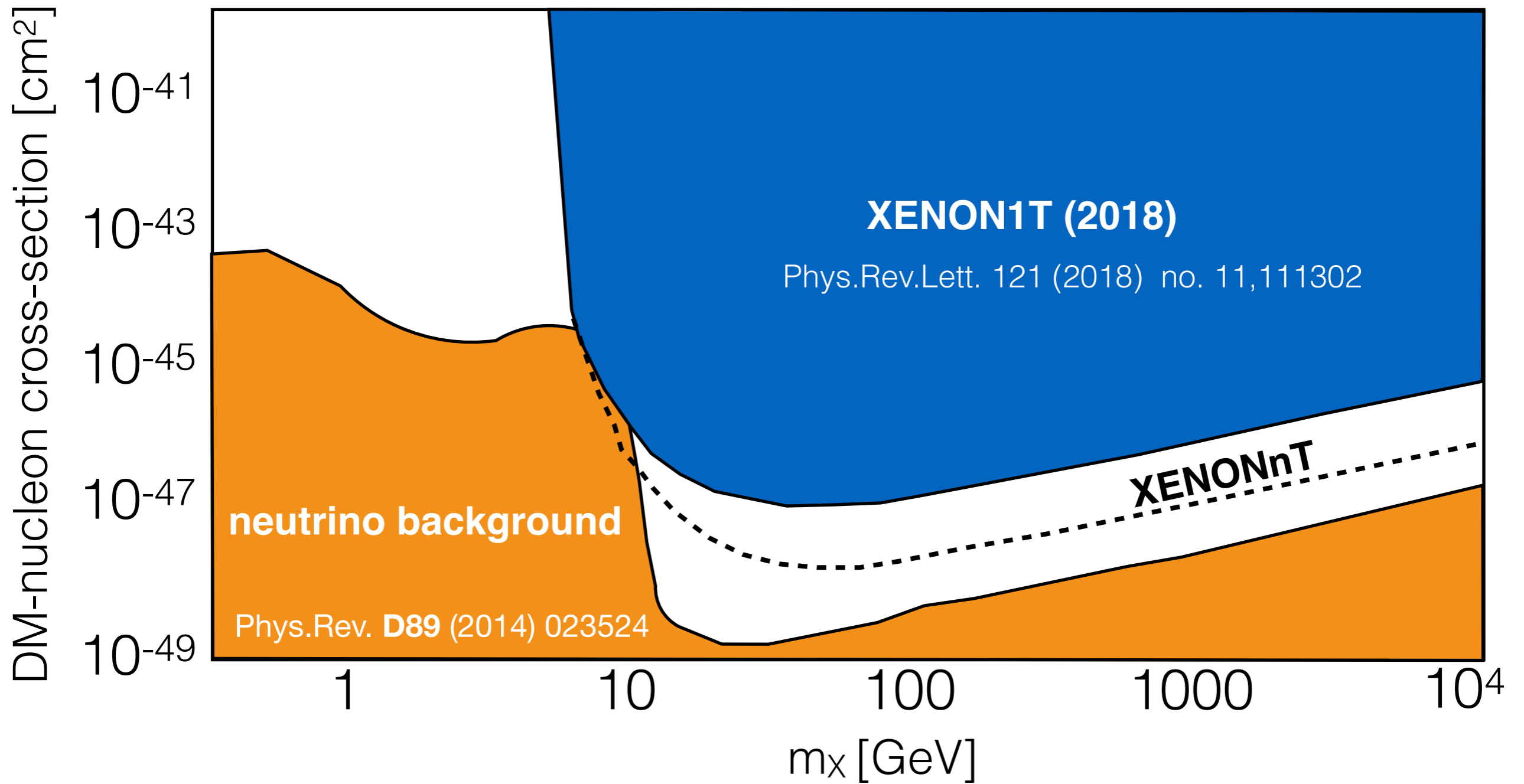
- **detect** galactic DM with underground detectors
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I will focus on this

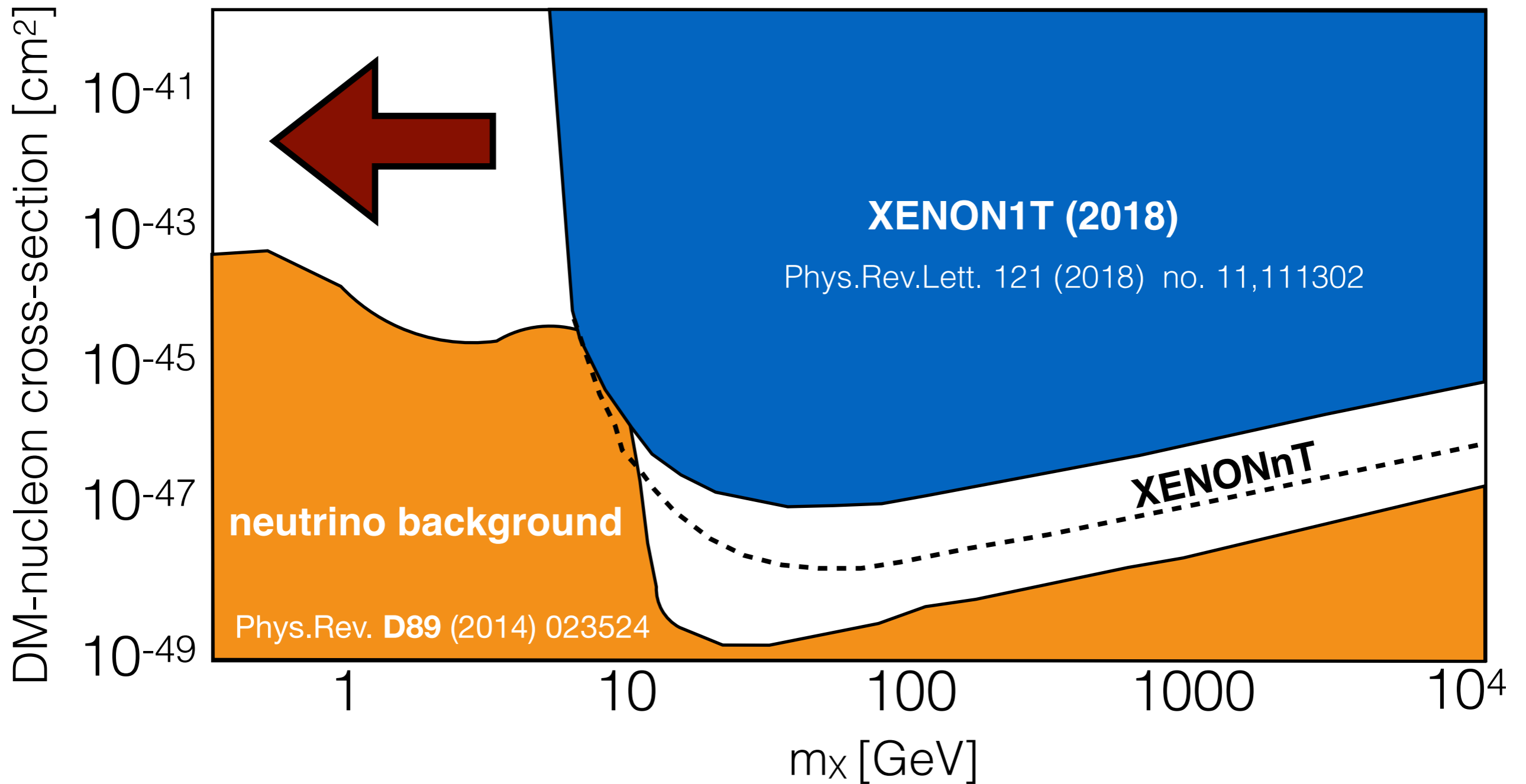
DM direct detection



direct detection



direct detection

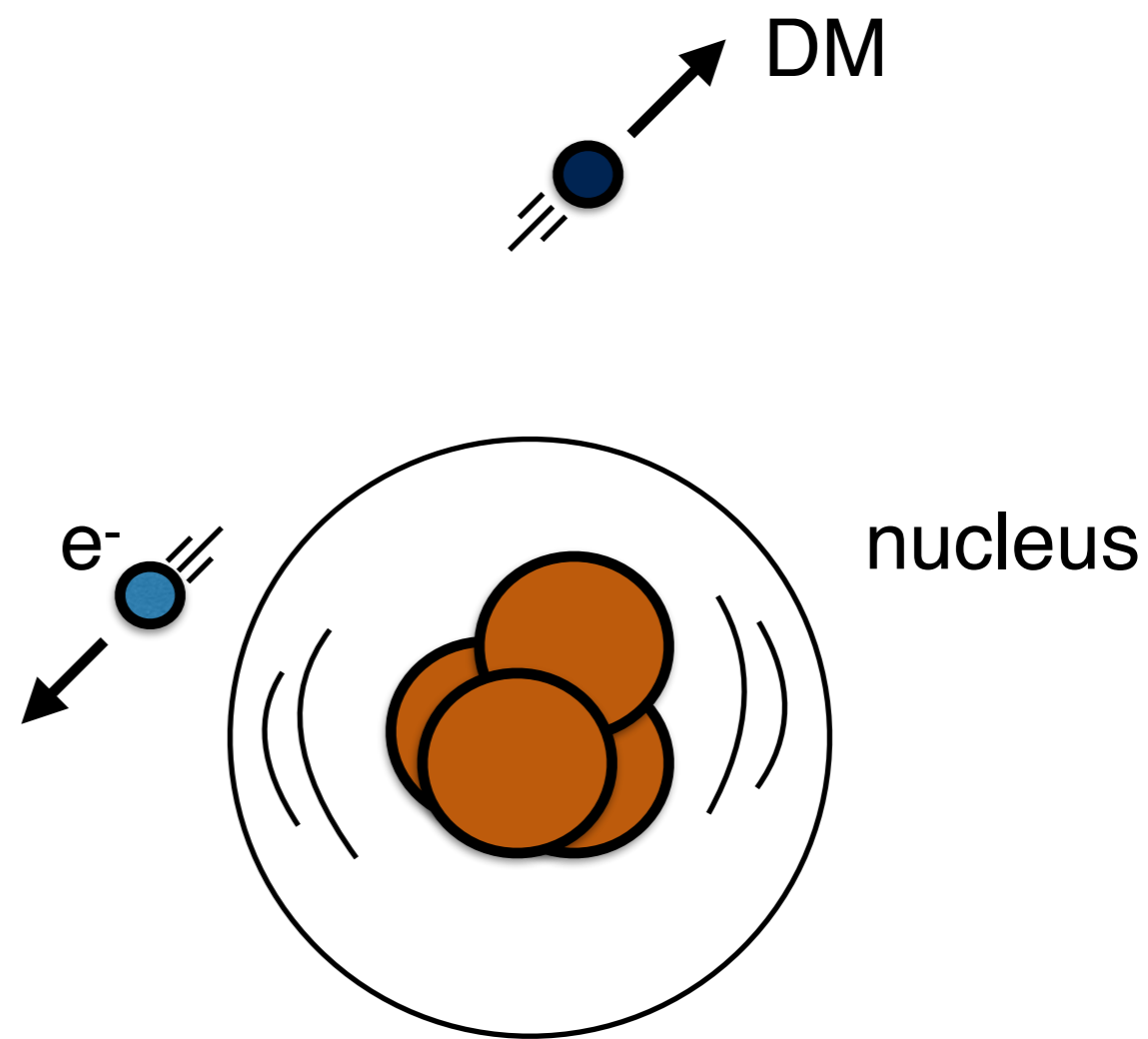
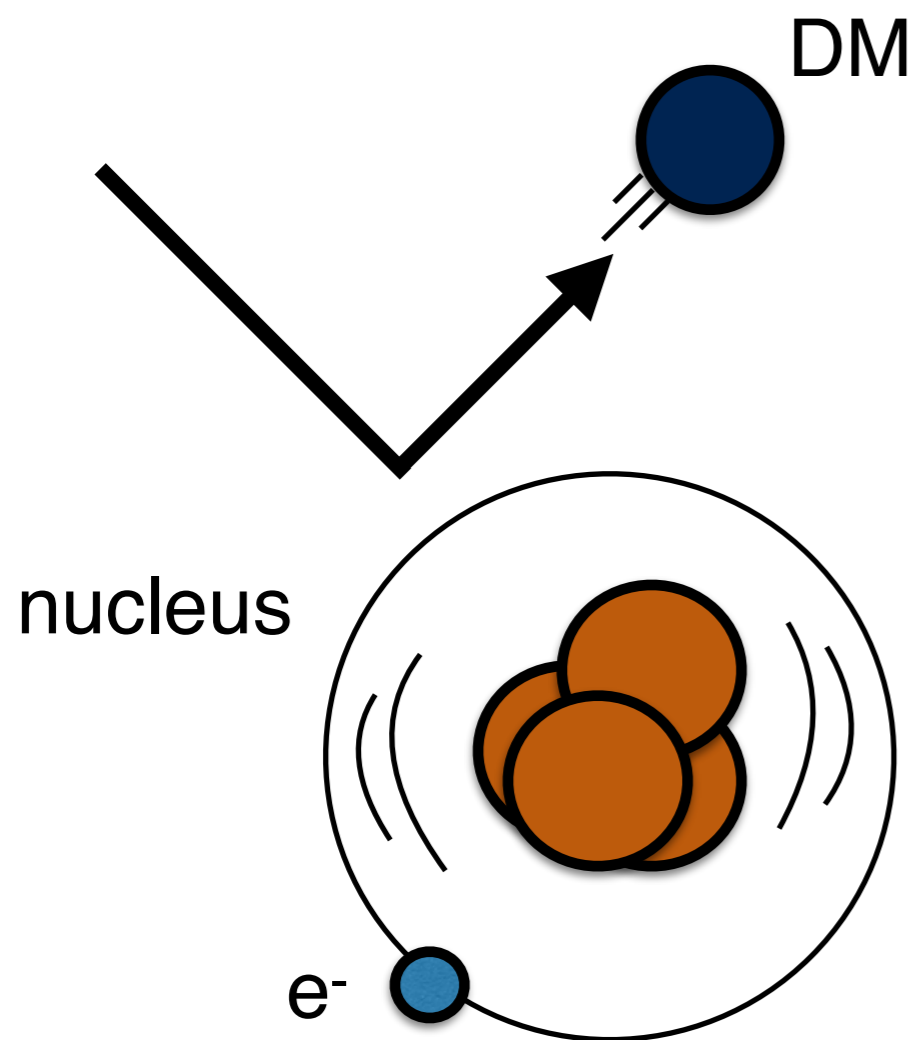


sub-GeV DM direct detection

- **Dark matter-electron scattering** in noble liquids, semiconductors, and organic molecules
- **Dark matter-nuclear scattering** through the Migdal scattering and bremsstrahlung
- **Absorption** of light dark matter, including axion-like particles and dark photons.
- **Dark matter scattering off collective modes** in molecules and in crystals (including phonons, plasmons and magnons)

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$$E_R = \frac{q^2}{2m_N}$$

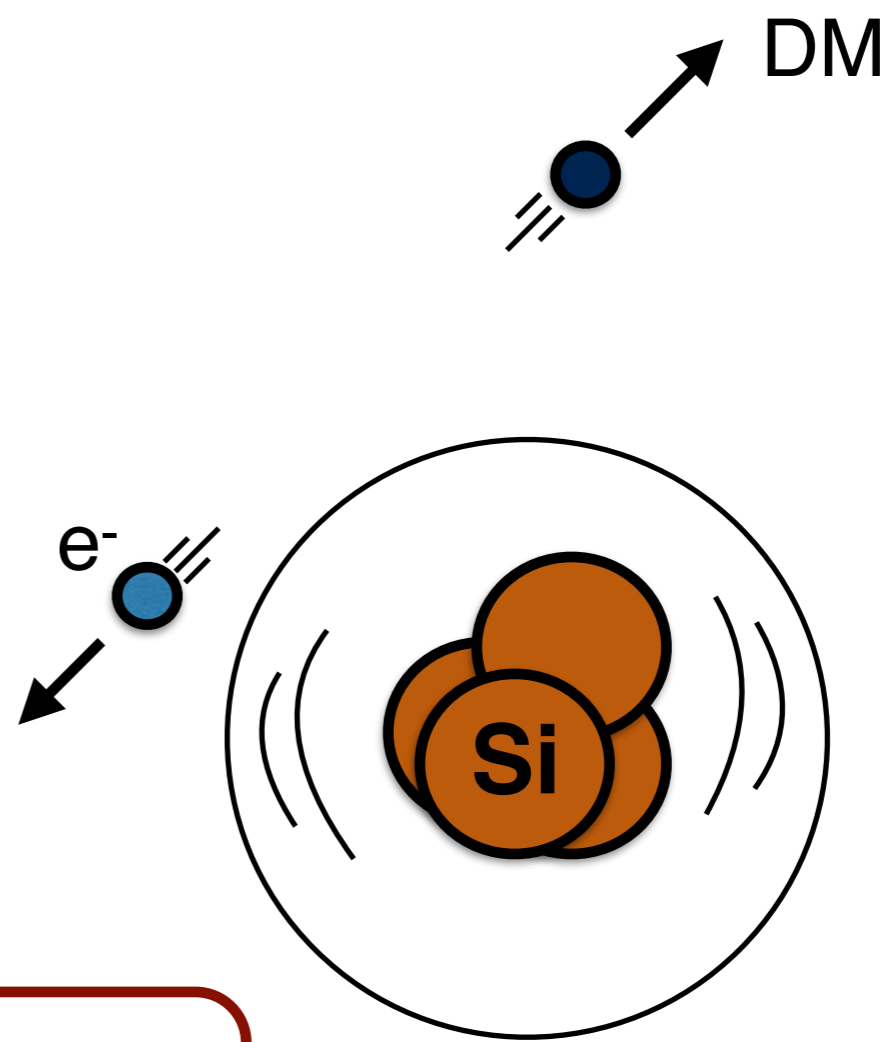
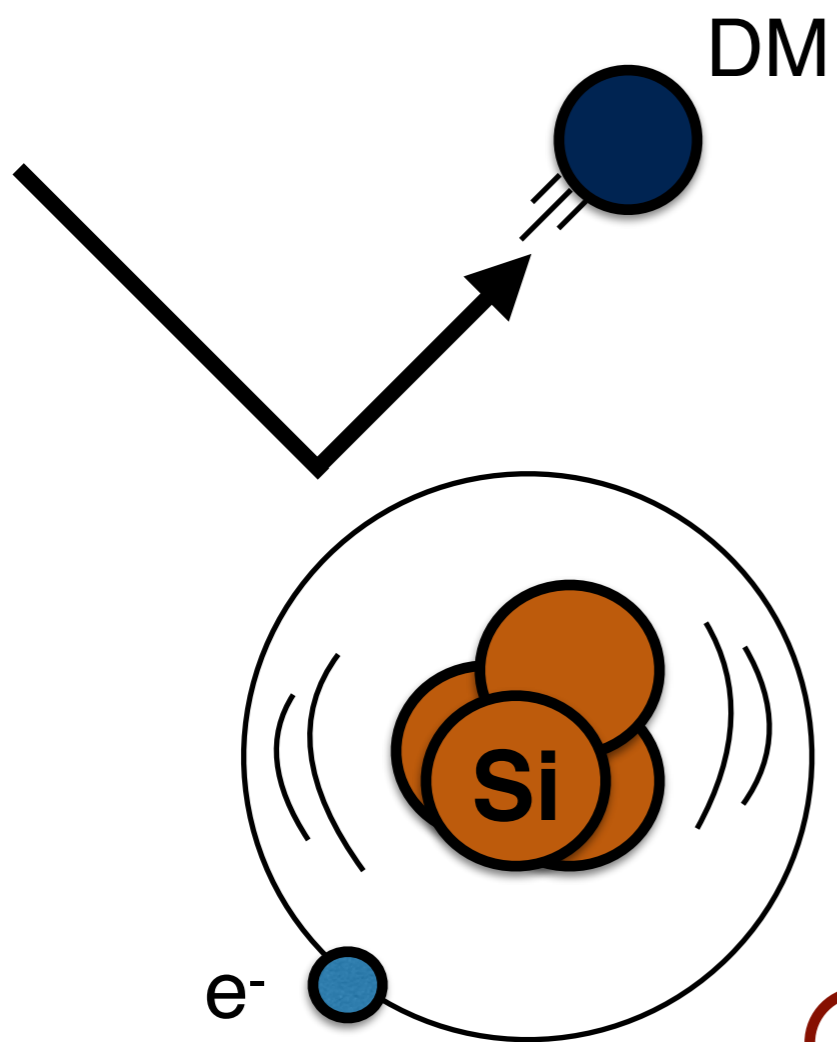
$$\simeq 50 \text{ keV} \left(\frac{m_\chi}{100 \text{ GeV}} \right)^2 \left(\frac{100 \text{ GeV}}{m_N} \right)$$

$$E_R = \vec{q} \cdot \vec{v} - \frac{q^2}{2\mu_{\chi N}}$$

$$\sim \frac{1}{2} \text{ eV} \times \left(\frac{m_\chi}{\text{MeV}} \right)$$

DM-nuclear scattering

DM-electron scattering



$$m_N = 28 \text{ GeV}$$
$$m_\chi = 100 \text{ MeV}$$

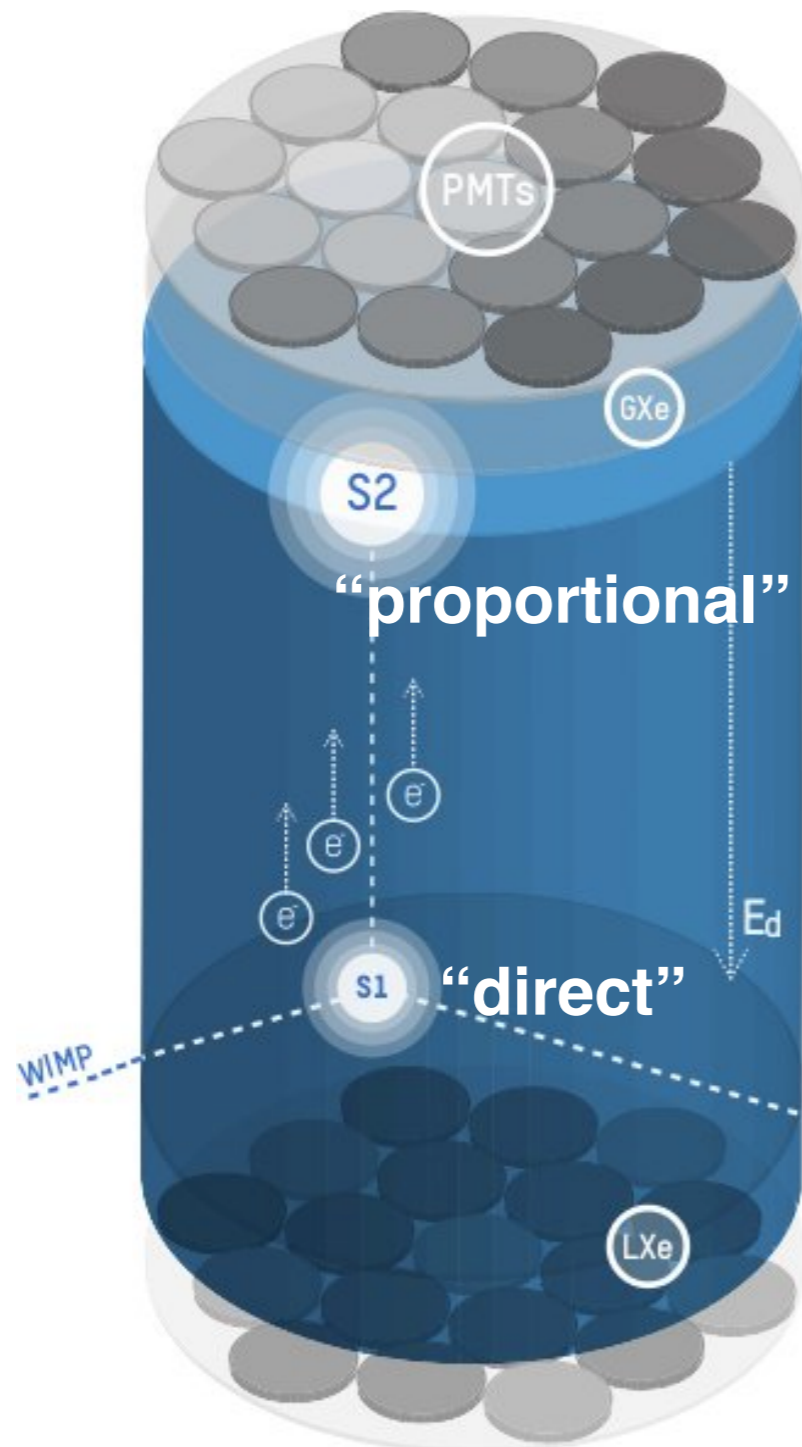
$$E_R \sim 0.1 \text{ eV}$$

$$E_R \sim 50 \text{ eV}$$

DM-nuclear scattering

DM-electron scattering

Liquid Xenon/Argon



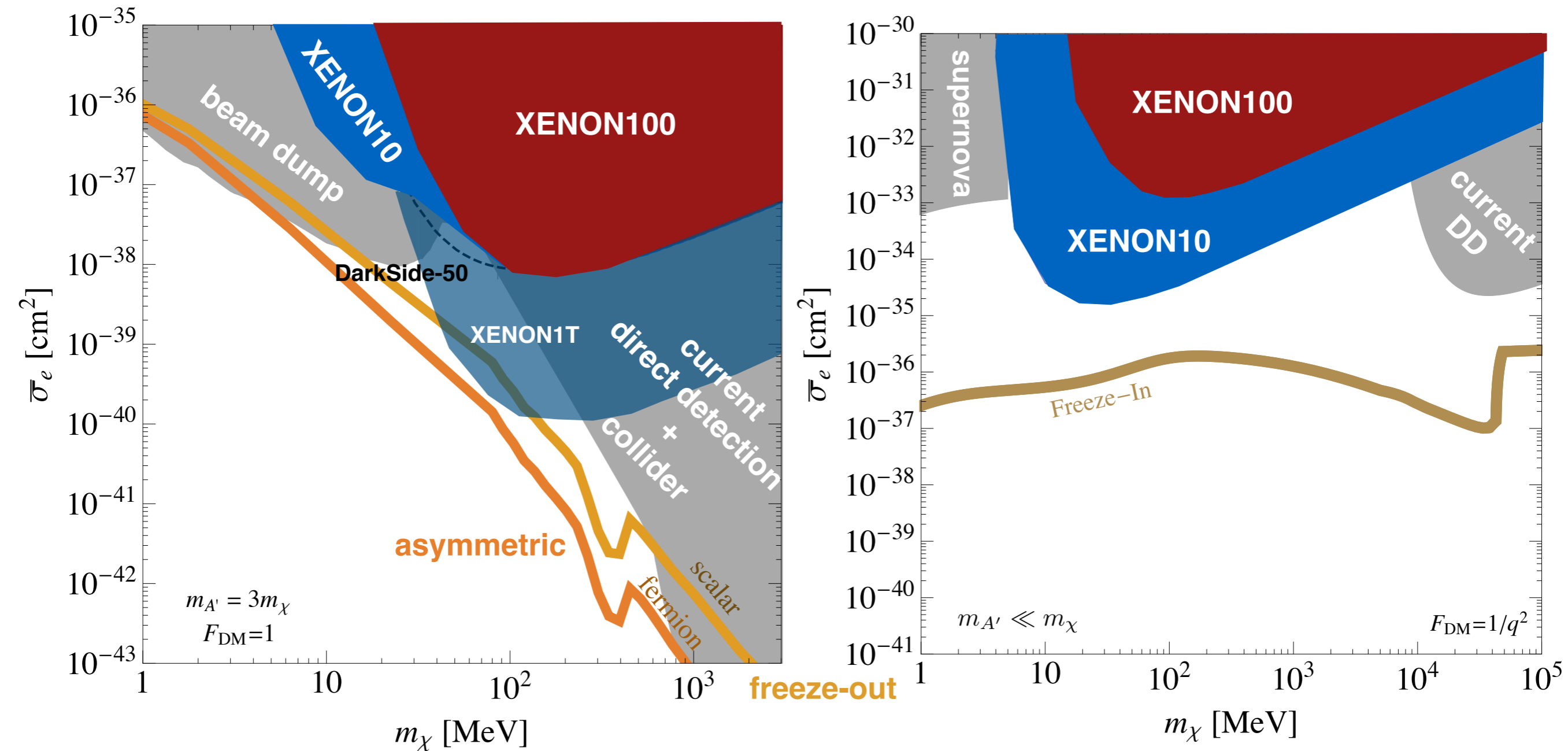
i.e. XENON10, XENON100, XENON1T, LUX,
DarkSide...

DM-electron scattering
=
S2 only signal

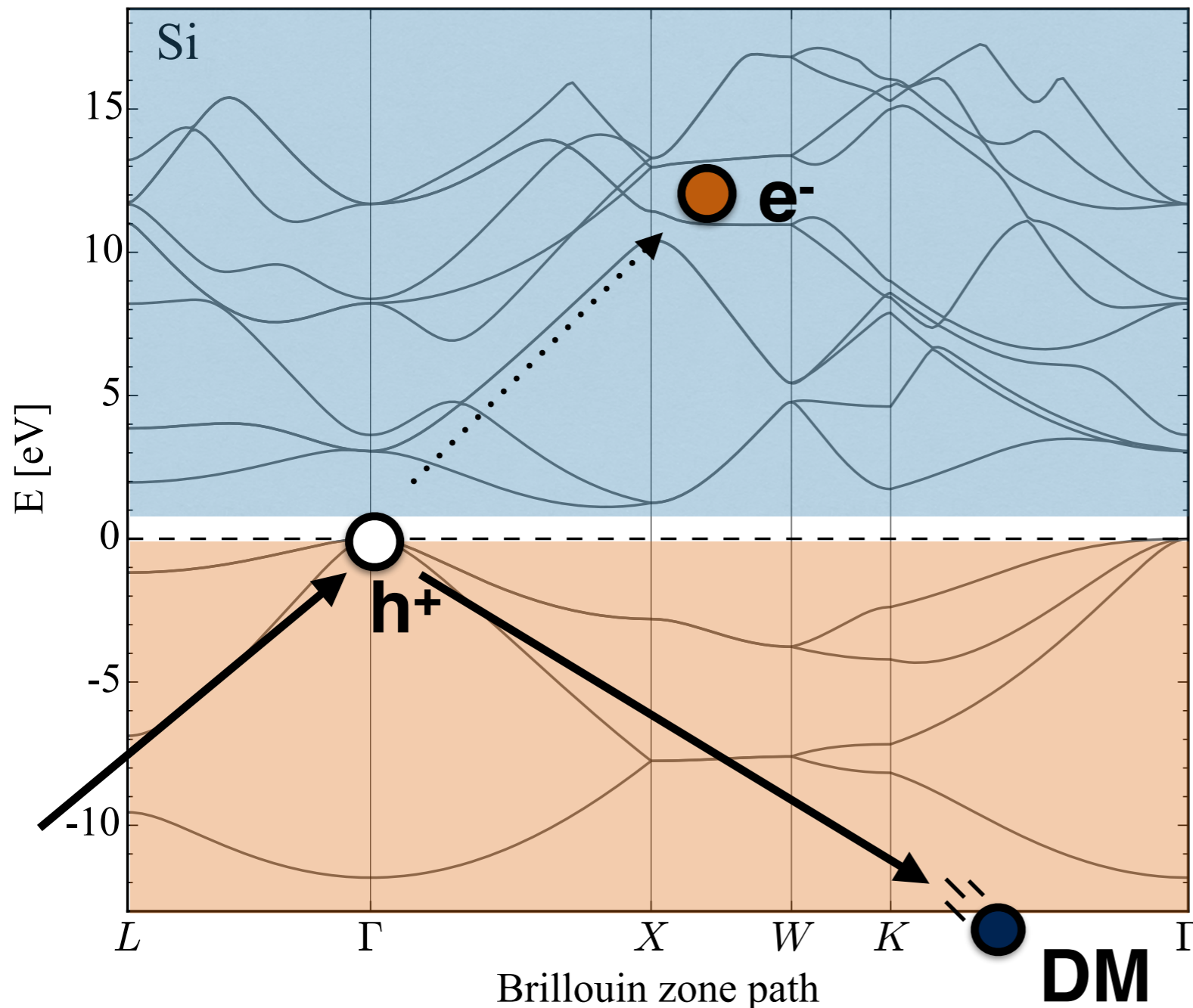
sensitive to ~10 eV energy depositions

measures **PhotoElectrons**

dark photon



semiconductor targets



detect the electron(s)

sensitive to $\sim eV$ energy depositions

i.e. silicon, germanium

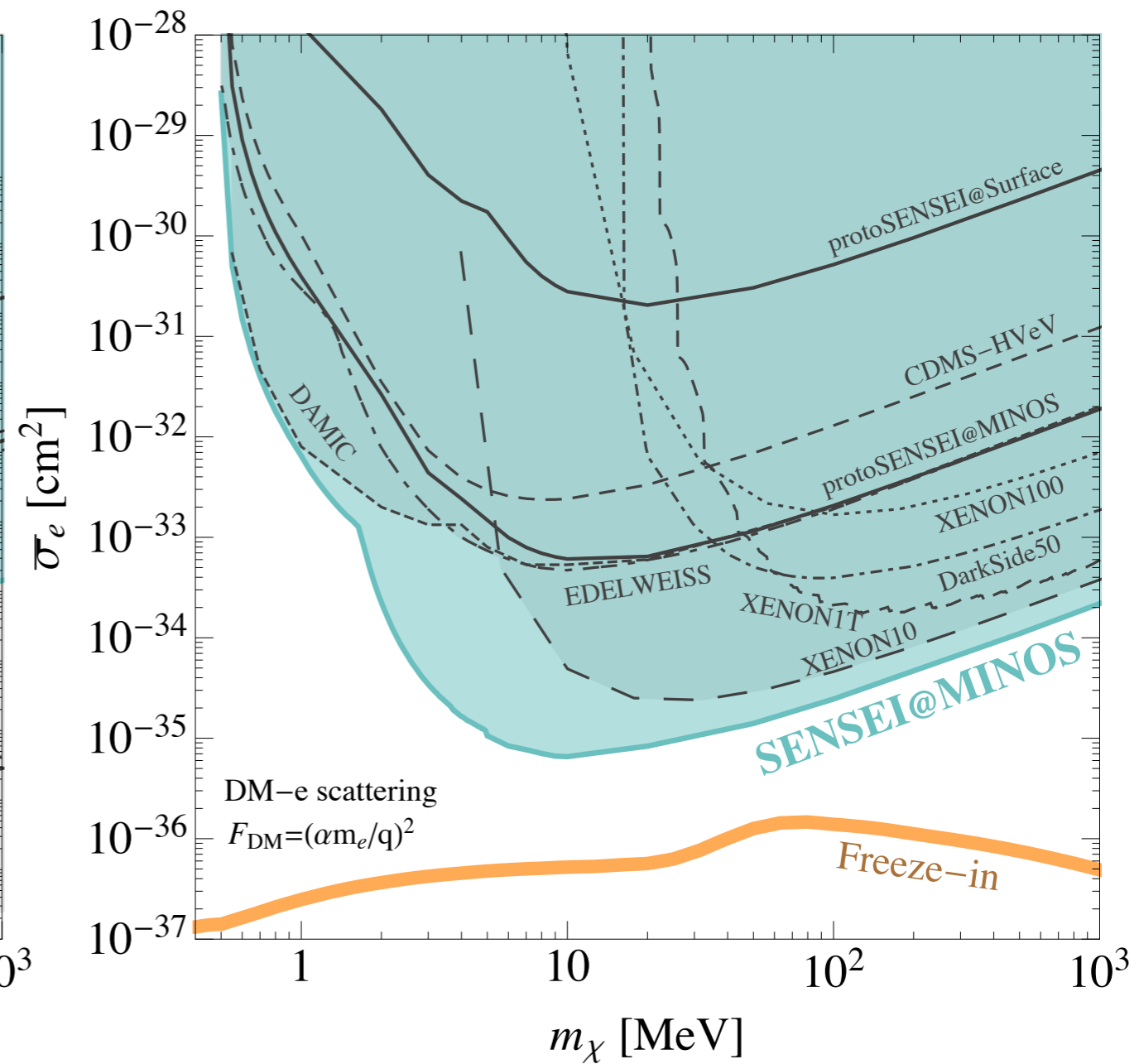
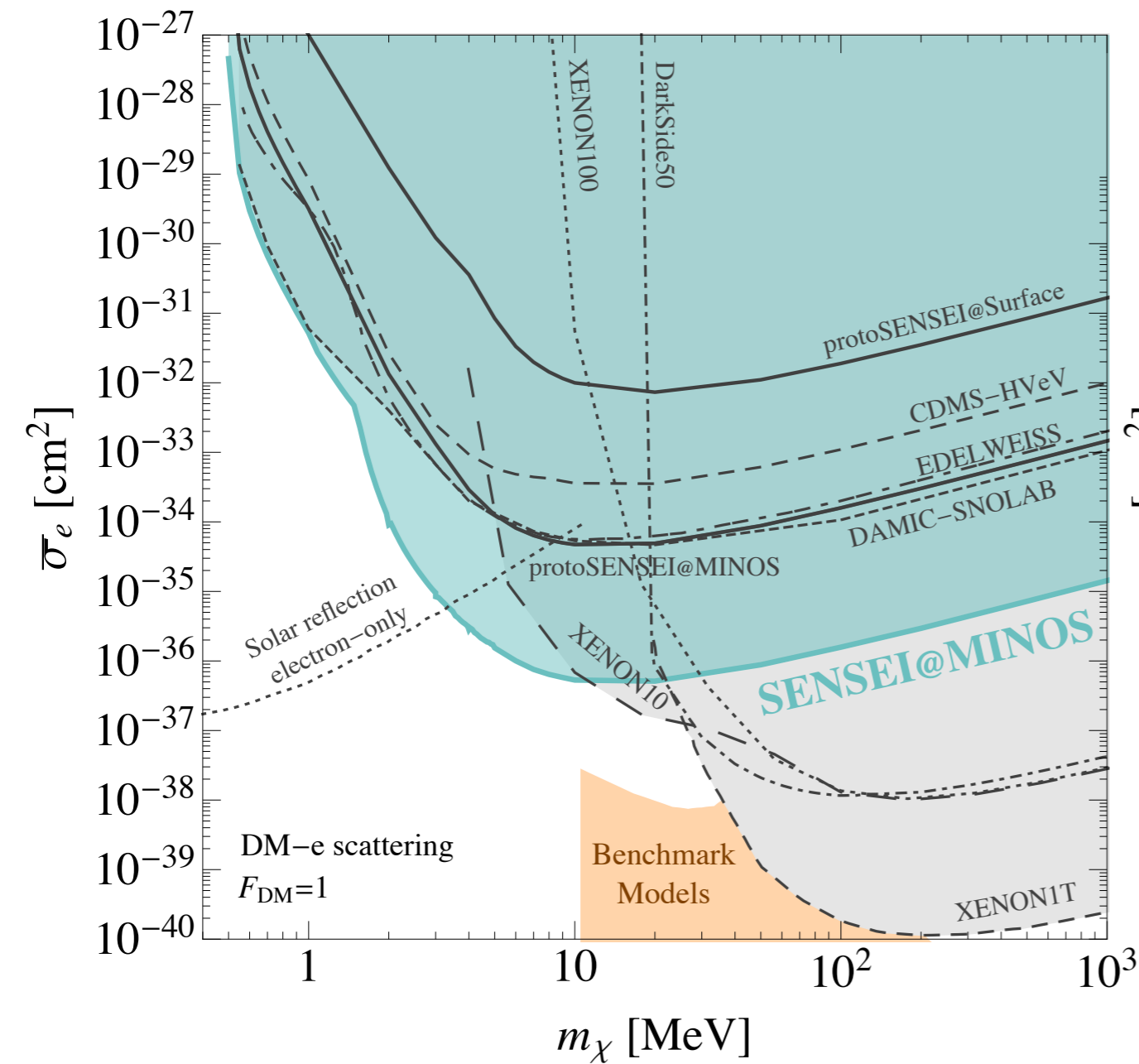
Essig, Mardon, Volansky [1108.5383]

Graham, Kaplan, Rajendran, Walters [1203.2531]

Lee, Lisanti, Mishra-Sharma, Safdi [1508.07361]

Essig, Fernandez-Serra, Mardon, Soto, Volansky, TTY [1509.01598]

DM-electron scattering



sub-GeV DM direct detection

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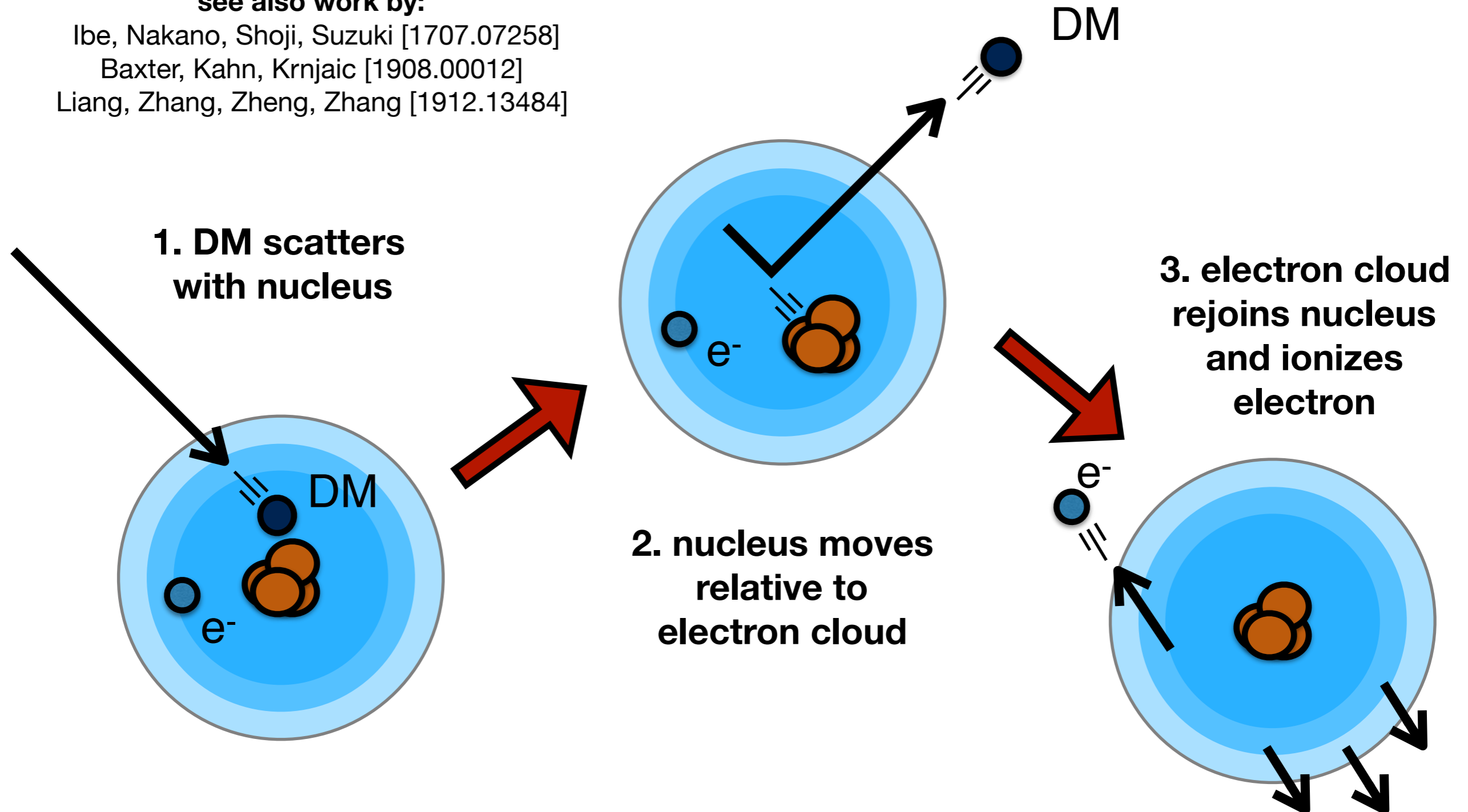
“Migdal” scattering

see also work by:

Ibe, Nakano, Shoji, Suzuki [1707.07258]

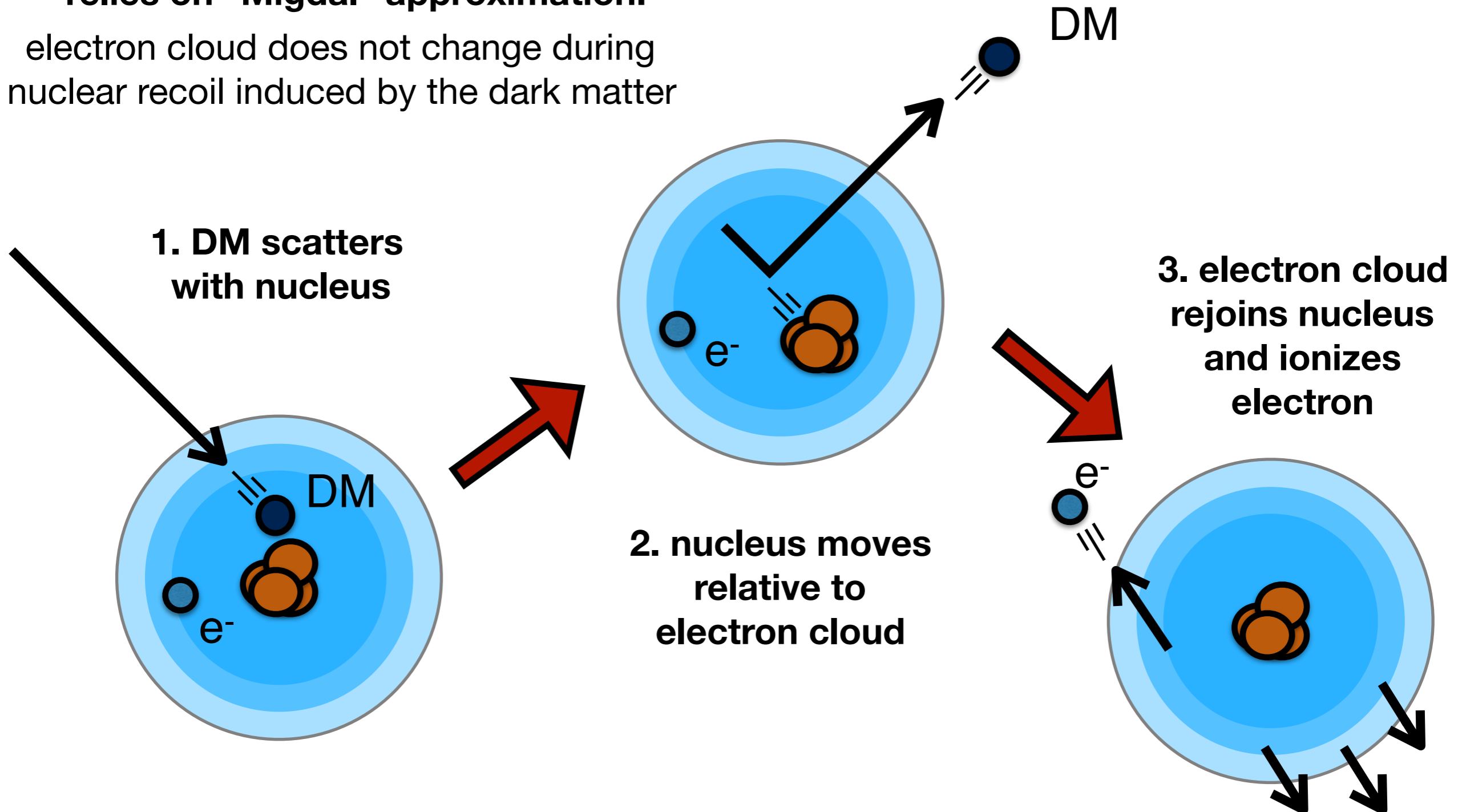
Baxter, Kahn, Krnjaic [1908.00012]

Liang, Zhang, Zheng, Zhang [1912.13484]



“Migdal” scattering

* relies on “Migdal” approximation:
electron cloud does not change during
nuclear recoil induced by the dark matter



“Migdal” scattering

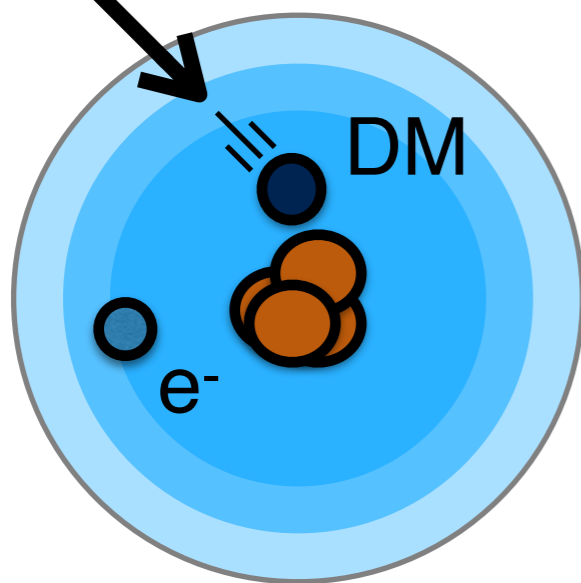
*

relies on “Migdal” approximation:

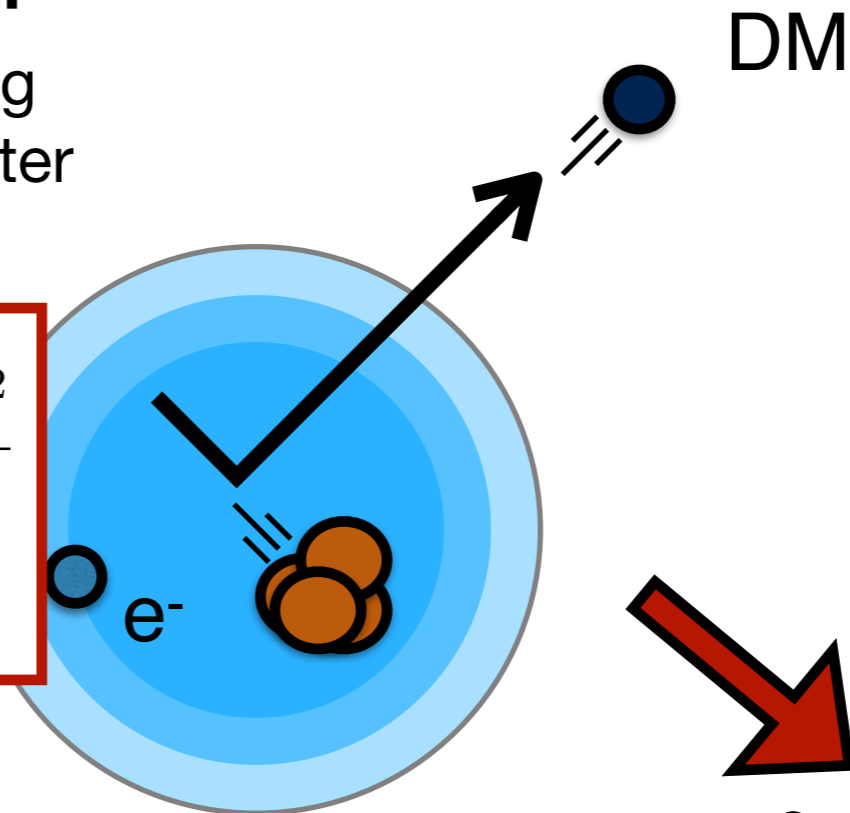
electron cloud does not change during nuclear recoil induced by the dark matter

$$\frac{d\sigma_N}{dE_R} \simeq \frac{1}{32\pi} \frac{m_A}{\mu_N^2 v_X^2} \frac{|F_A(q_A^2)|^2 |\mathcal{M}(q_A)|^2}{(m_A + m_X^2)^2}$$

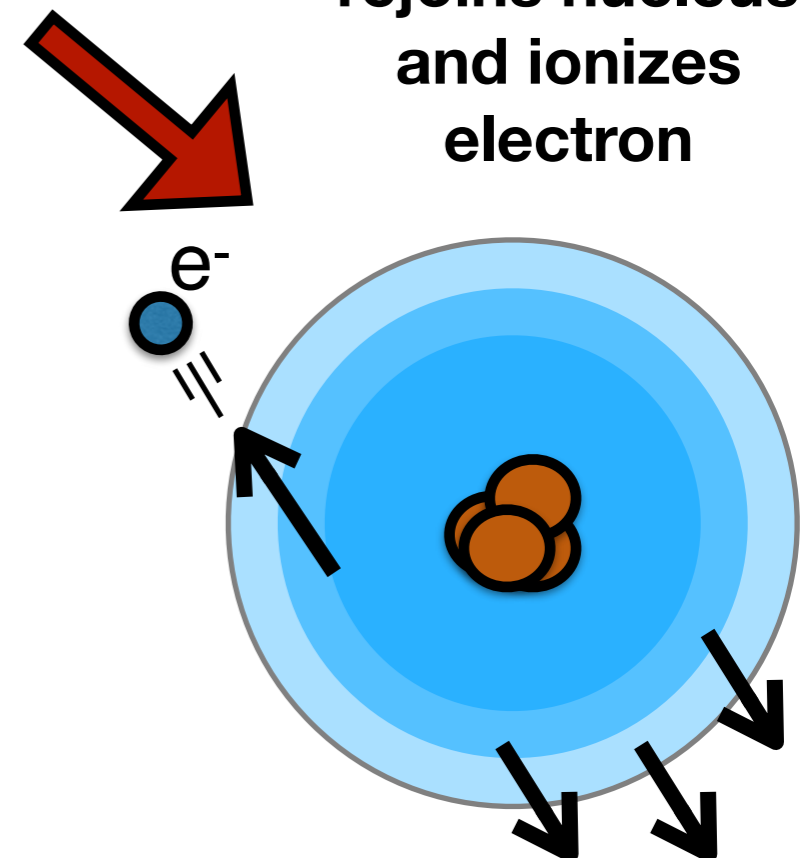
DM-nucleus scattering



2. nucleus moves relative to electron cloud



3. electron cloud rejoins nucleus and ionizes electron



“Migdal” scattering

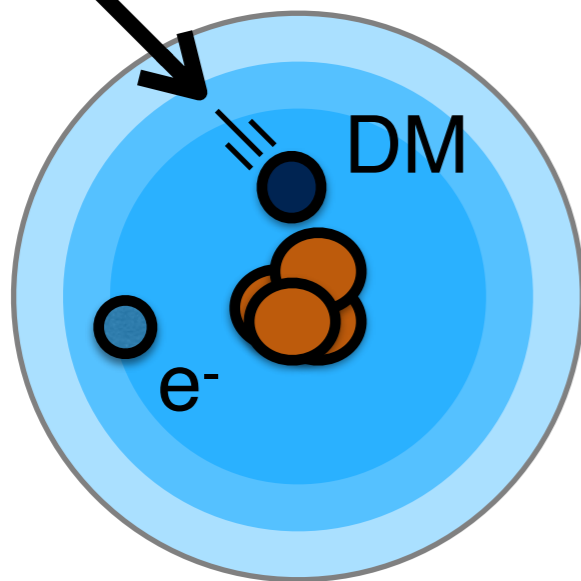
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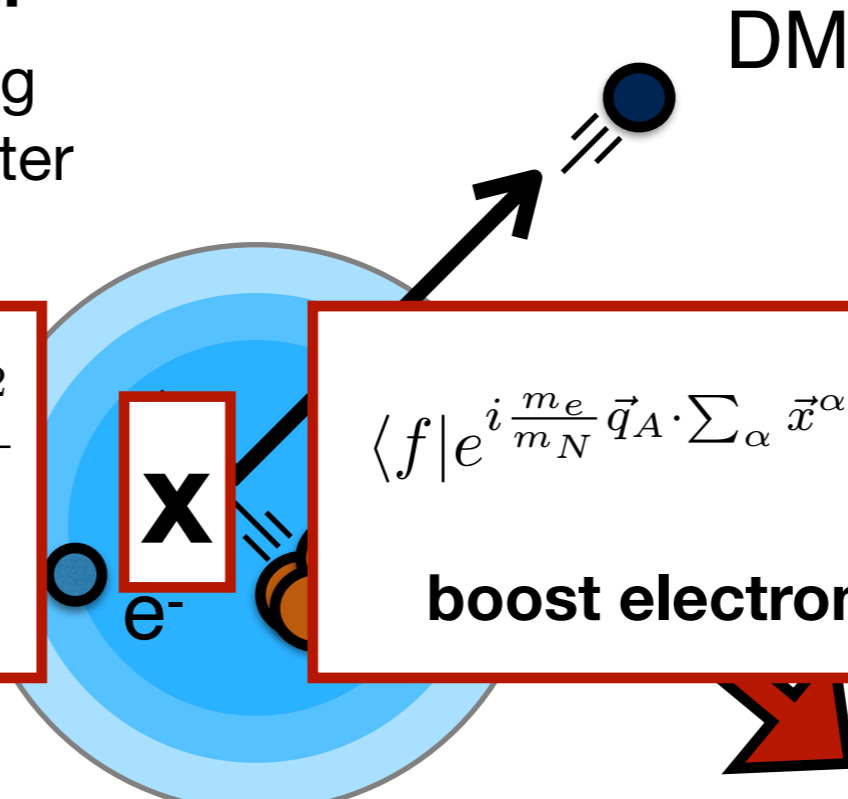
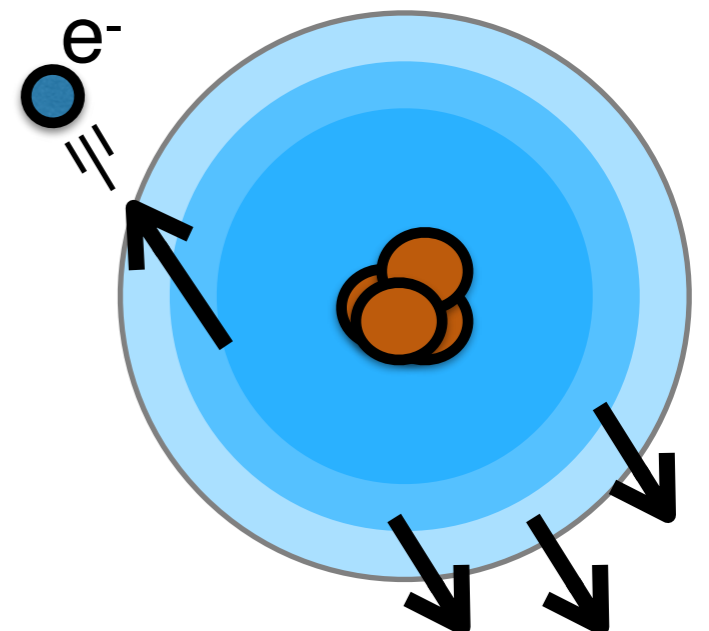
DM-nucleus scattering

$$\langle f | e^{i \frac{m_e}{m_N} \vec{q}_A \cdot \sum_{\alpha} \vec{x}^{\alpha}} | i \rangle \simeq \frac{i}{e} \frac{m_e}{m_N} \vec{q}_A \cdot \vec{d}_{fi}$$

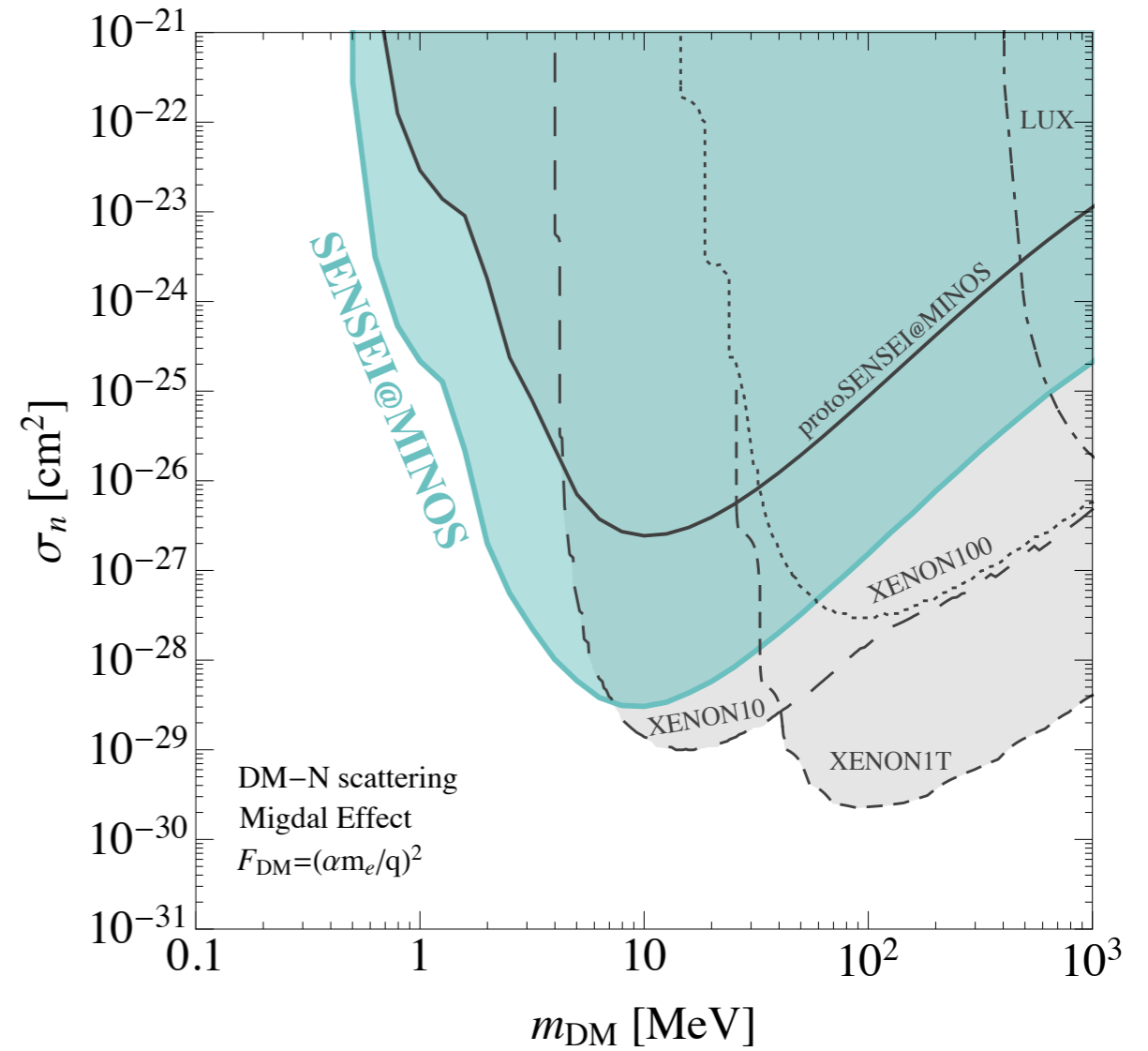
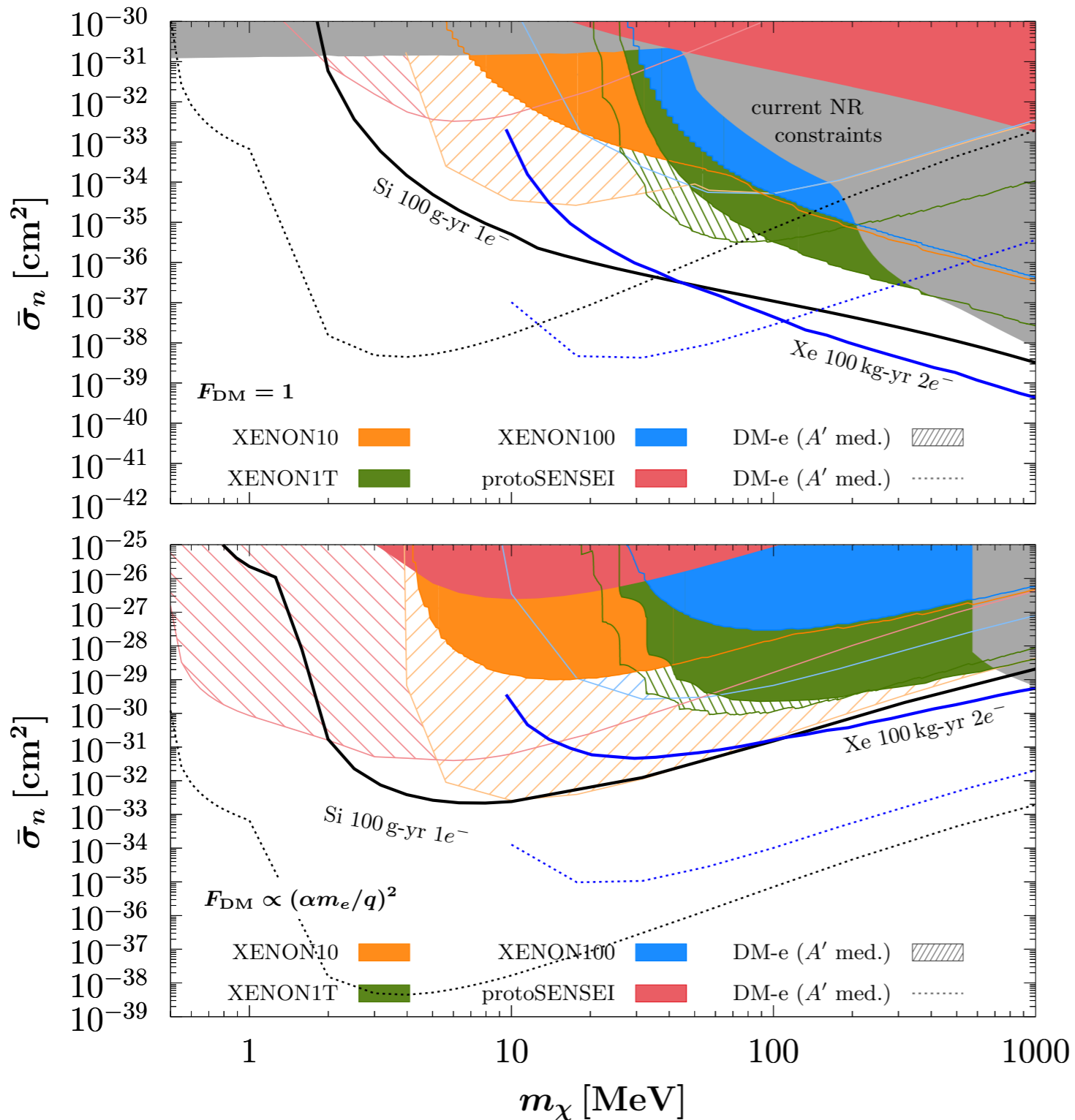
**boost electron cloud from $i > f$
electron**



**2. nucleus moves
relative to
electron cloud**



“Migdal” scattering



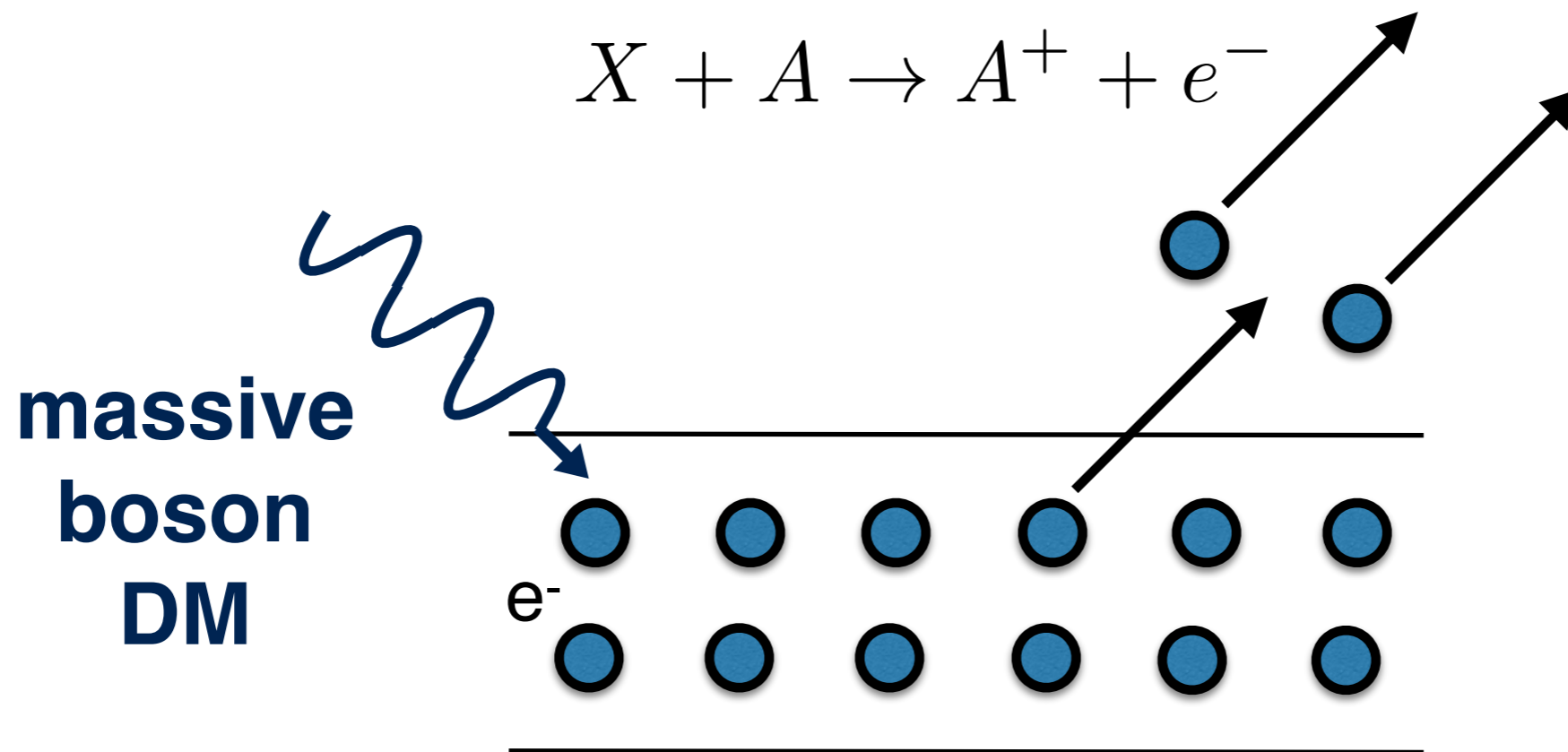
SENSEI collaboration [arXiv:2004.11378], submitted to PRL

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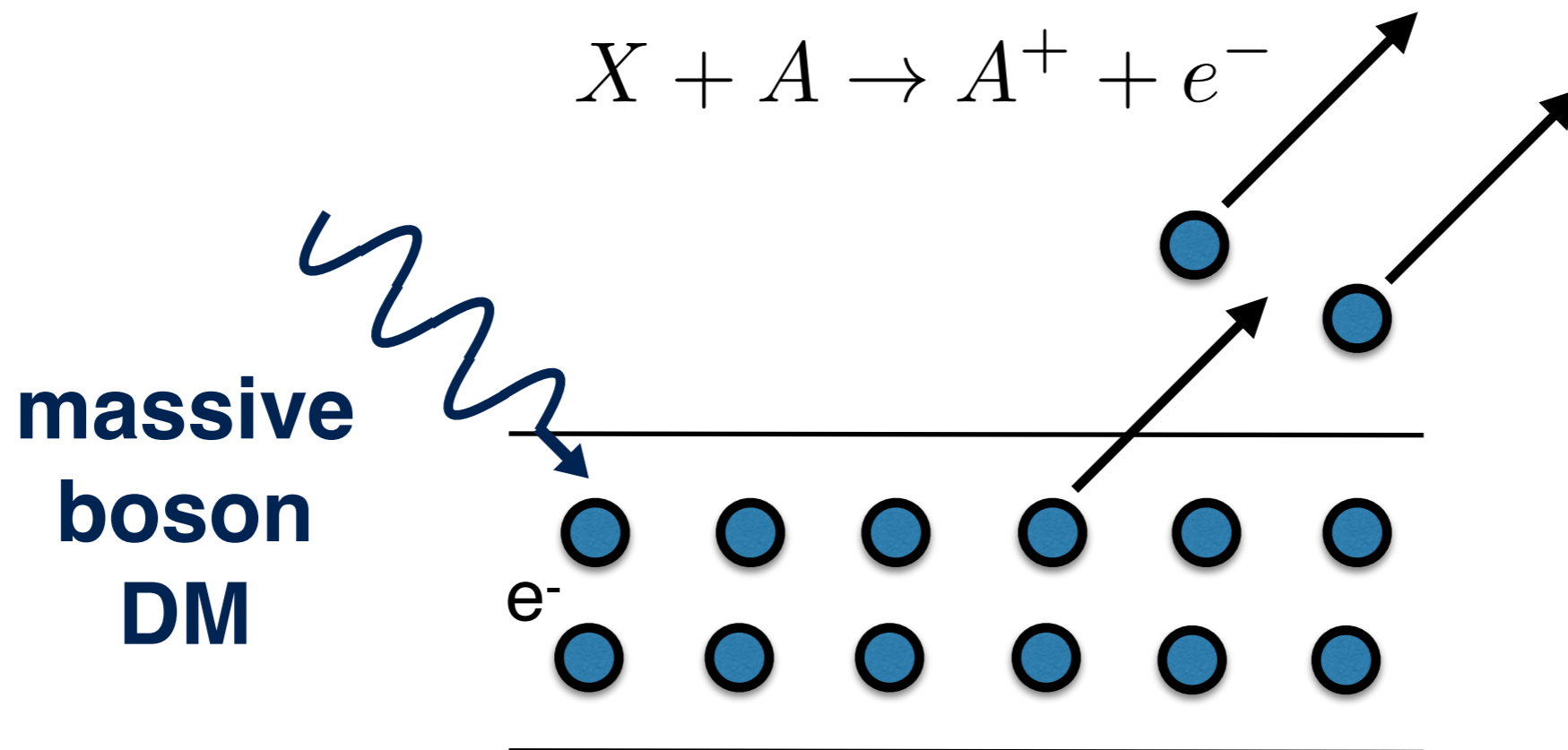
see talk by M. Baryakhtar
for other examples

photoelectric effect



**absorb all of the energy
the incoming dark matter**

photoelectric effect



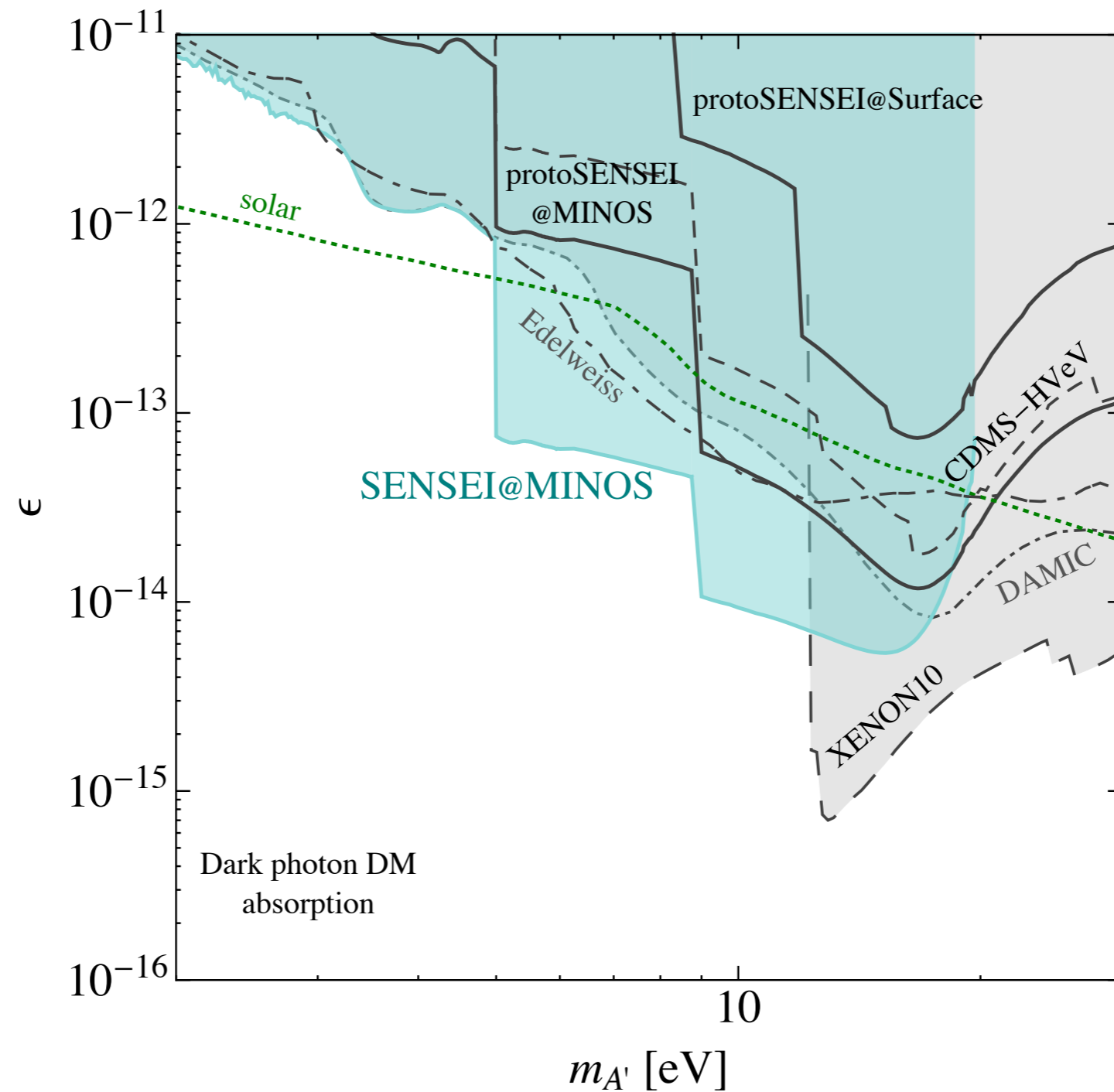
photon
 $|\vec{q}| = \omega$
 bosonic dark matter
 $|\vec{q}| = m_X v_{\text{DM}} \sim 10^{-3} \omega$

$< |\vec{q}_e|$

can relate massive boson absorption to photon absorption

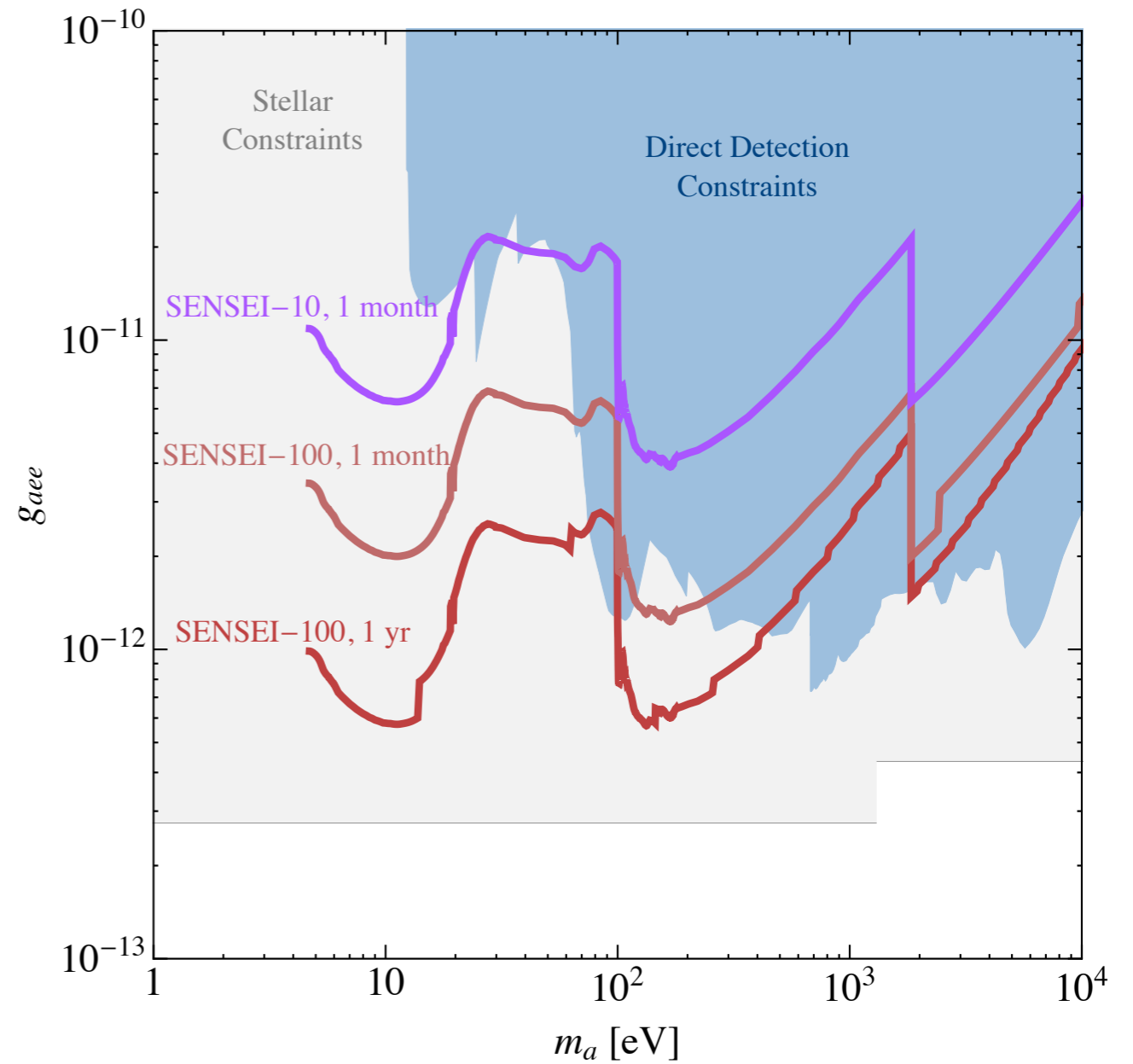
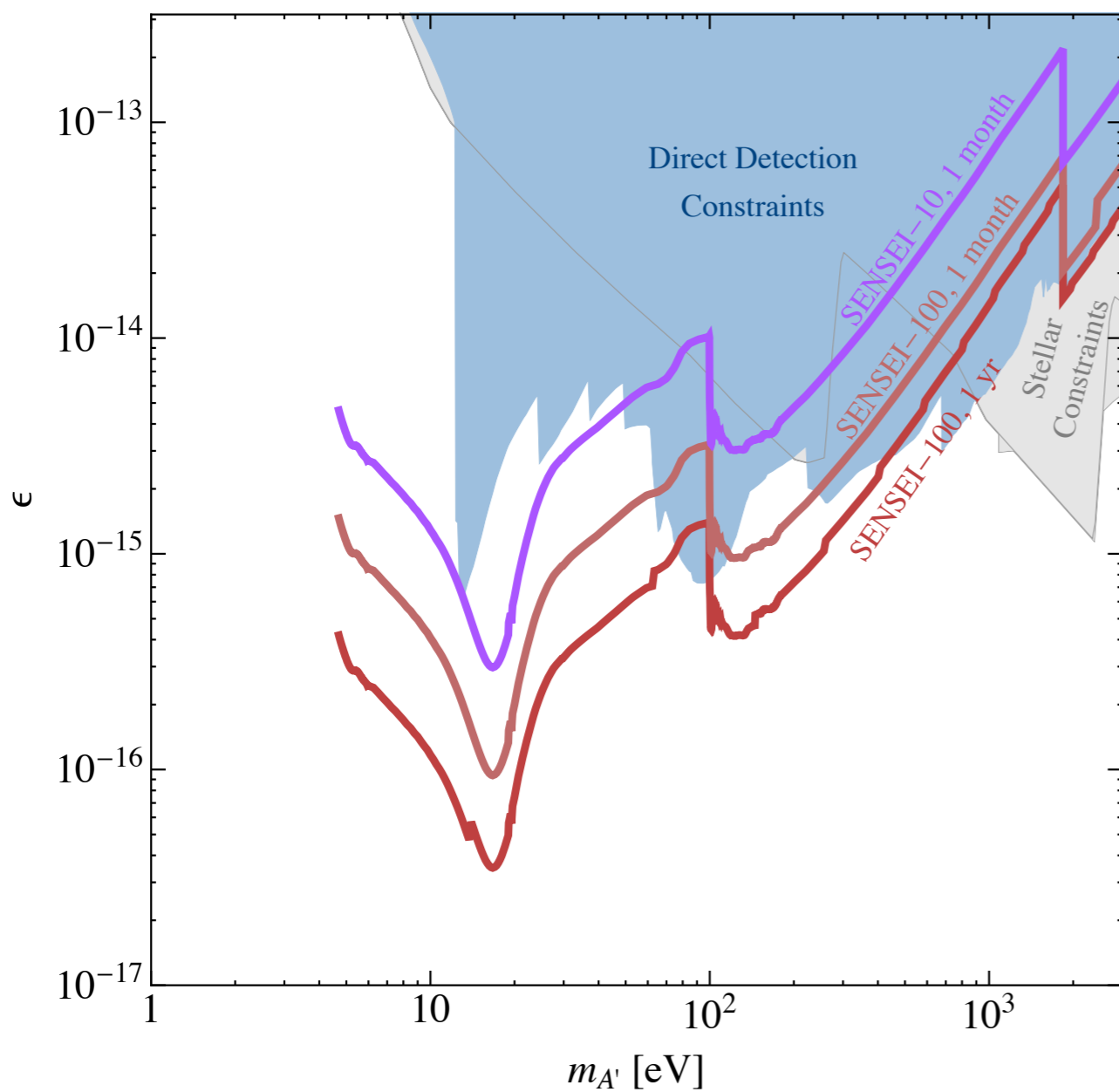
$$\sigma_{\text{DM}}(\omega) \propto \sigma_{\text{PE}}(\omega)$$

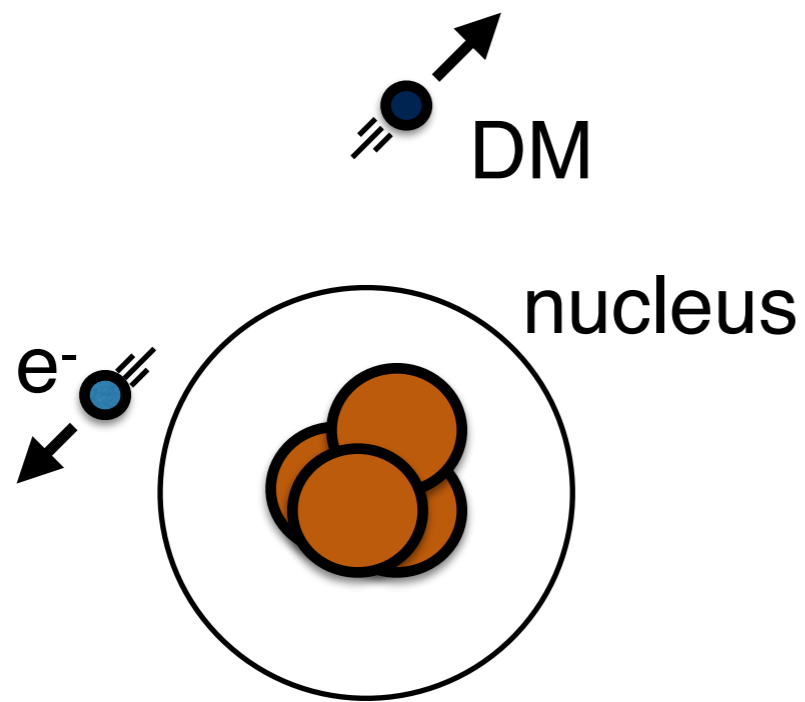
Dark Photon DM



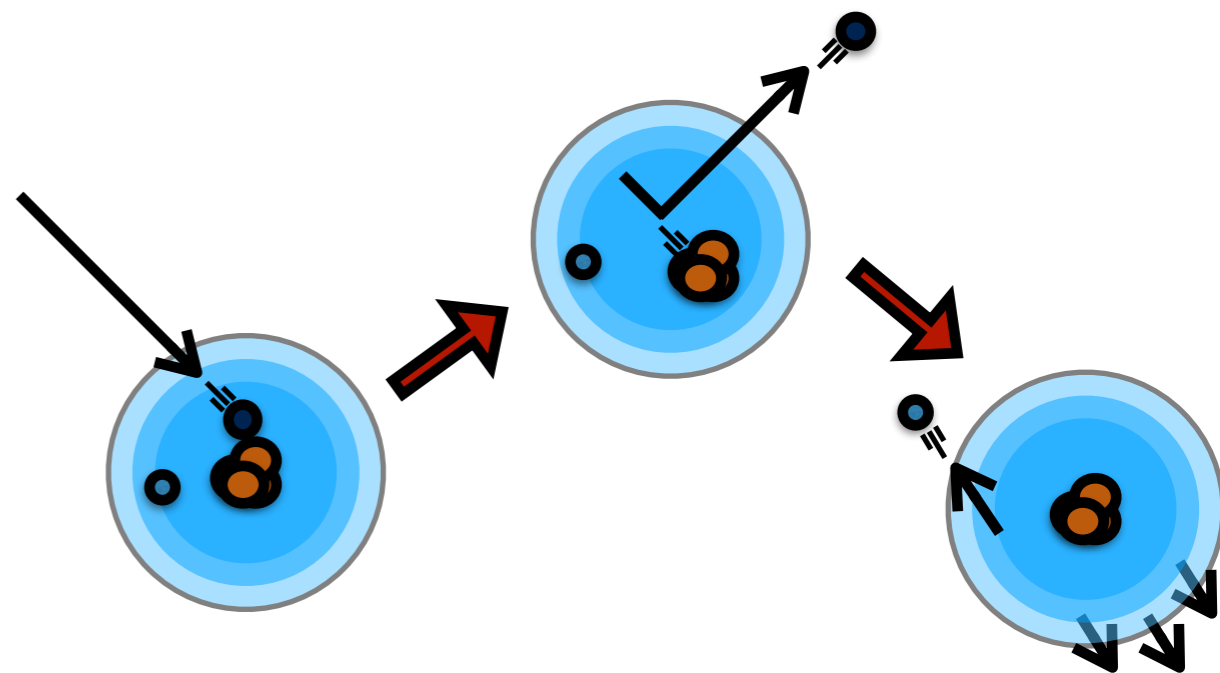
physics potential

bosonic absorption

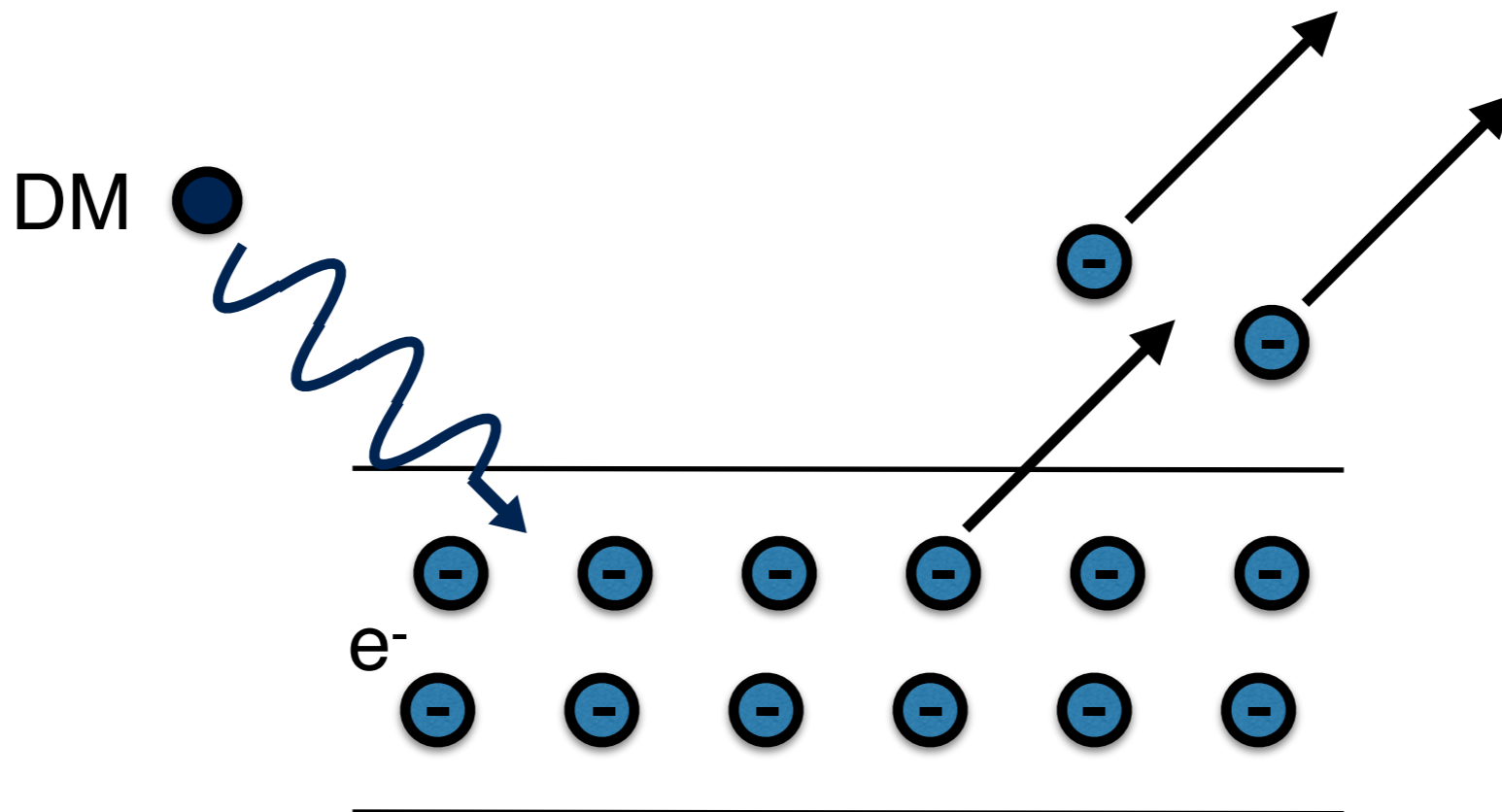




dark matter-electron scattering



dark matter-nucleus scattering



dark matter absorption

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see talk by M. Pyle

collective modes

- phonons: collective excitation of atoms in a crystal
 - acoustic: “in-phase”
 - optical: “out-of-phase”
- magnons: collective excitation of electron spin
 - sensitive to spin-dependent interactions (DM-coupling to electron spin)
- sensitivity to \sim keV DM masses
- may require new materials — area of active research
- there are also plasmons (collective excitation of electrons)...

see work by e.g. A. Caputo, A. Esposito, E. Geoffray, Y. Kahn, S. Knapen, G. Krnjaic, S. Griffin, T. Lin, T. Melia, A. Mitridate, A. D. Polosa, S. Rajendran, S. Sun, T. Trickle, Z. Zhang, K. Zurek, ...

dark matter candidates

