



Canada's national laboratory  
for particle and nuclear physics  
and accelerator-based science

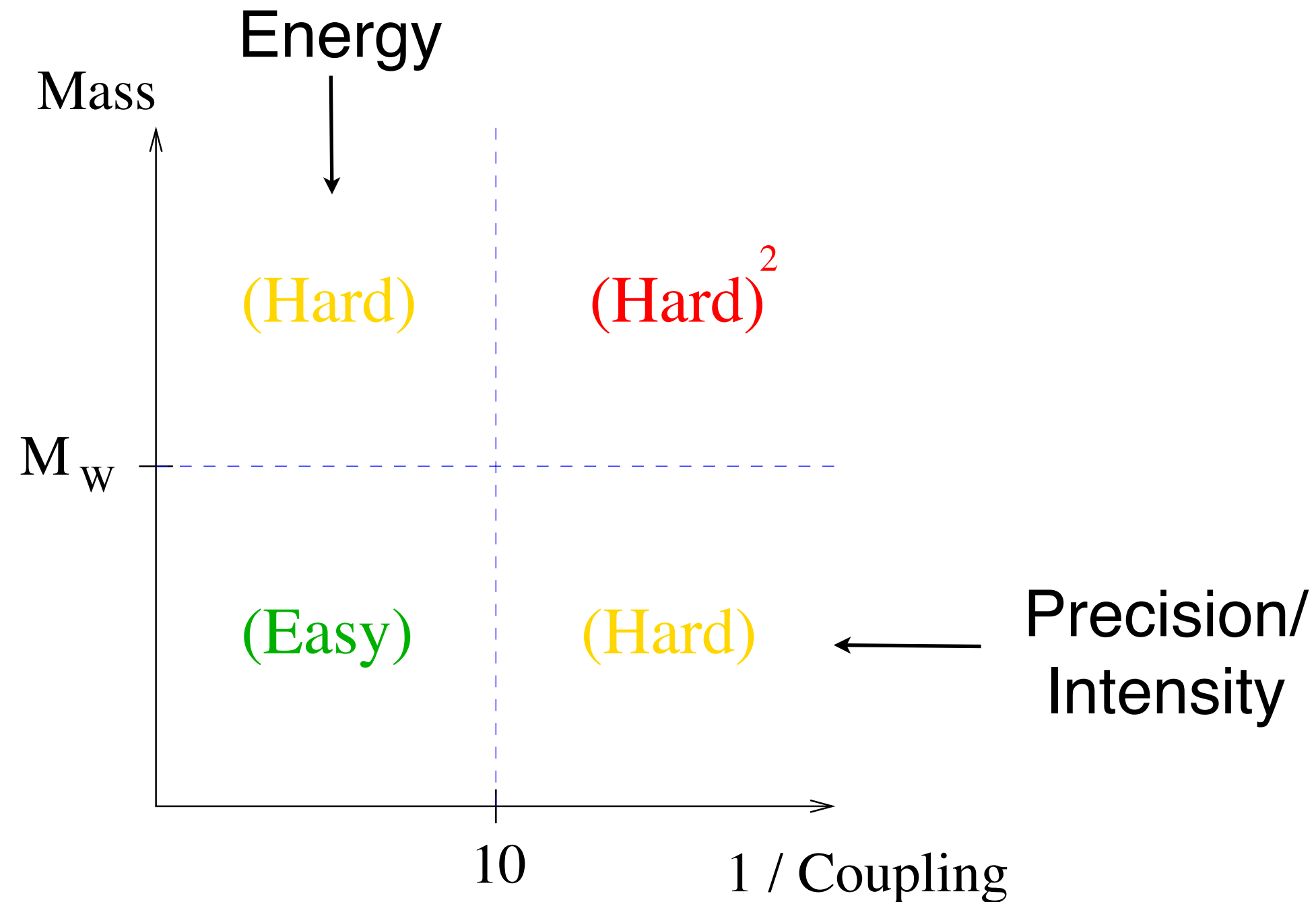
# New Directions

## TRIUMF Science Week 2017

[Summary by D.E. Morrissey]

- Goals:
  - test fundamental physics to greater precision
  - discover new phenomena
- Tools:
  - higher energy
  - higher precision and intensity
- Higher energy usually requires large colliders (LHC, ILC, ...).  
Increasing precision/intensity may be possible in smaller exps.

- Searching for new phenomena:

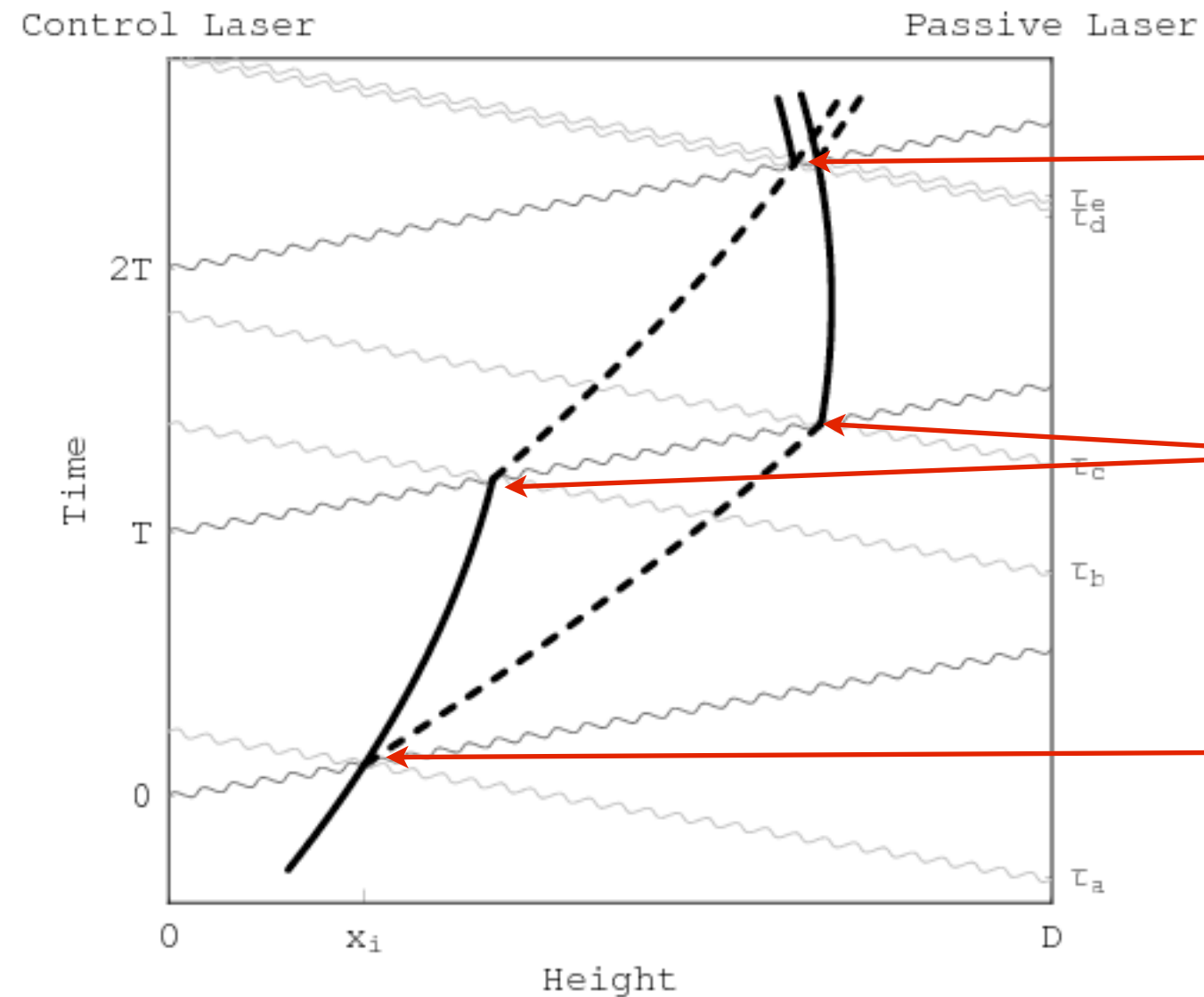


- Atomic methods for fundamental physics
- Searches for light dark sectors
- Searches for very light dark matter or axions
- Positronium and muonium
- Tests of sterile neutrinos
- Precision Higgs measurements beyond the (HL-)LHC
- Cosmology: CMB, 21cm radiation, ...

- Atomic interferometry (AI) is the precision frontier.
- Can be used to probe gravity effects: [Dimopoulos, Graham, Hogan, Kasevich 2008]
  - launch cloud of cold atoms upwards into vacuum
  - split into superposition of momentum states with laser pulse
  - unsplit momentum states with laser pulse
  - recombine (spatially) and measure the phase shift via interference
- Applications include tests of GR and searches for grav. waves.

- Atomic interferometry is the precision frontier.
- Can be used to probe gravity effects:

[Dimopoulos, Graham, Hogan, Kasevich 2008]

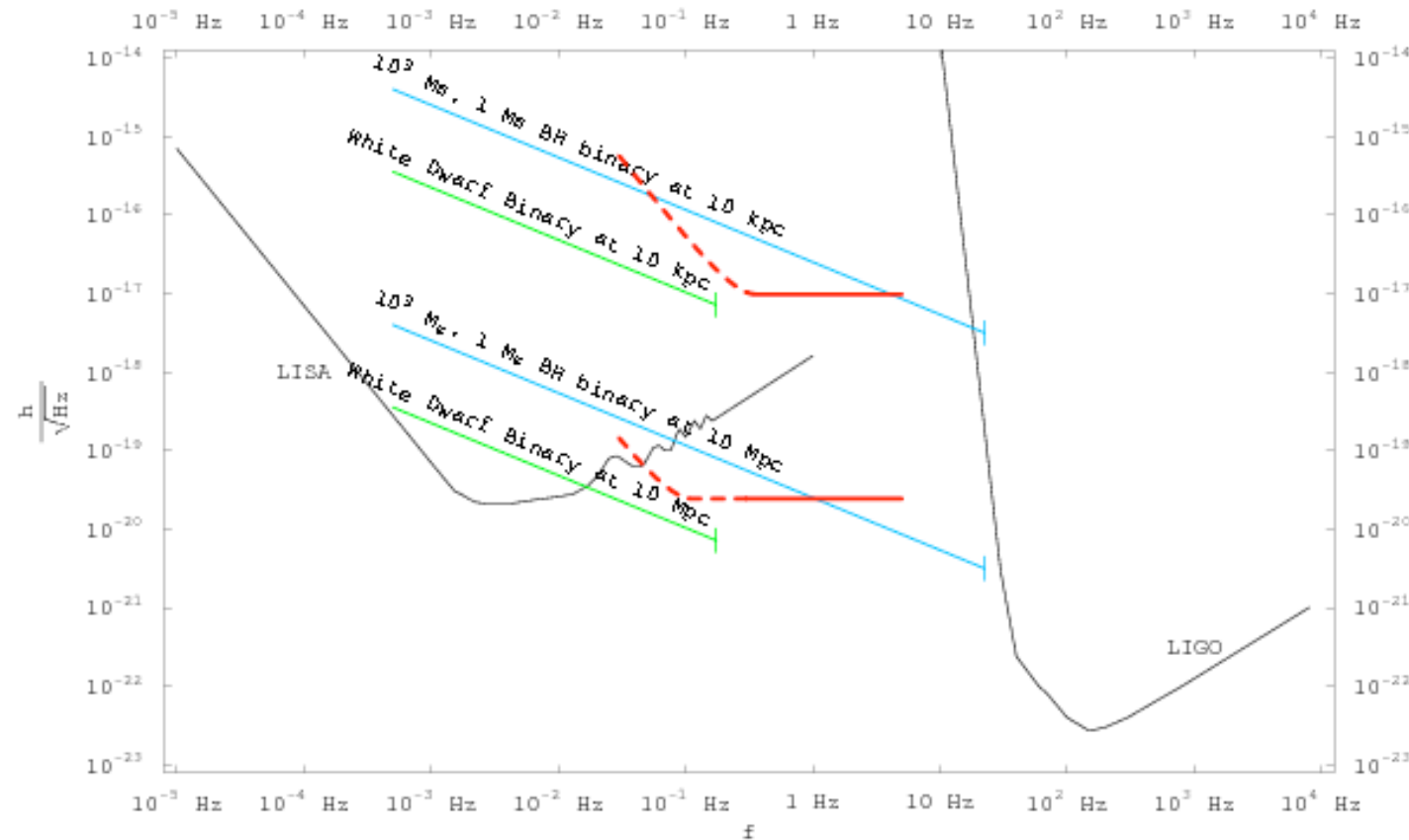


recombine and interfere

unsplit momentum states

split momentum states

- Applications:
  - precision tests of GR (equivalence principle, self-coupling,...)
  - detection of gravity waves (AGIS) or varying constants



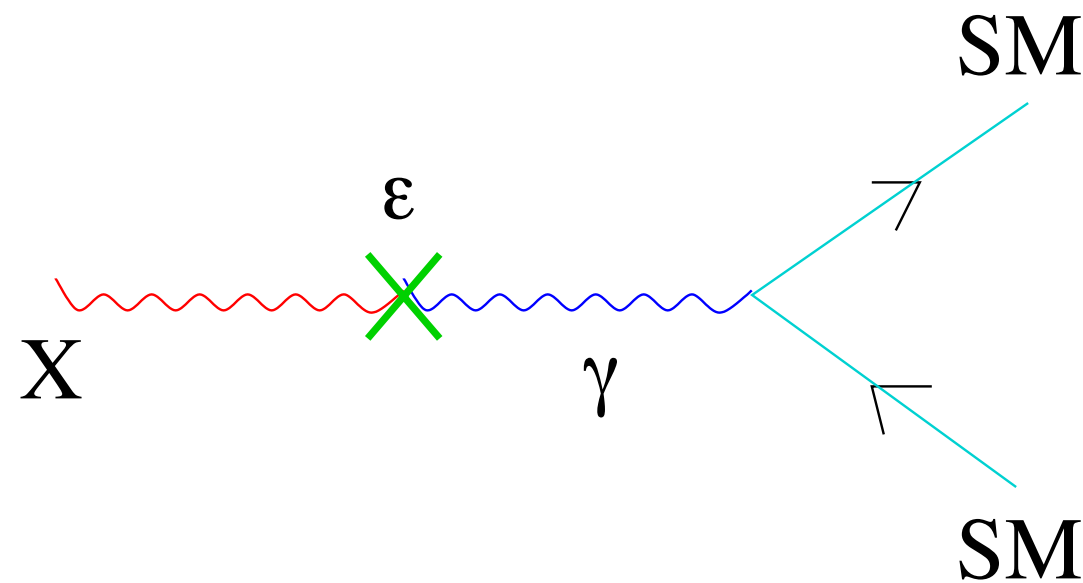
[Dimopoulos, Graham, Hogan, Kasevich 2008]

- two 10m tall evacuated apparatus
- 1km separation

- Many other methods and possibilities... [e.g. Arvanitaki *et al.* 2016]

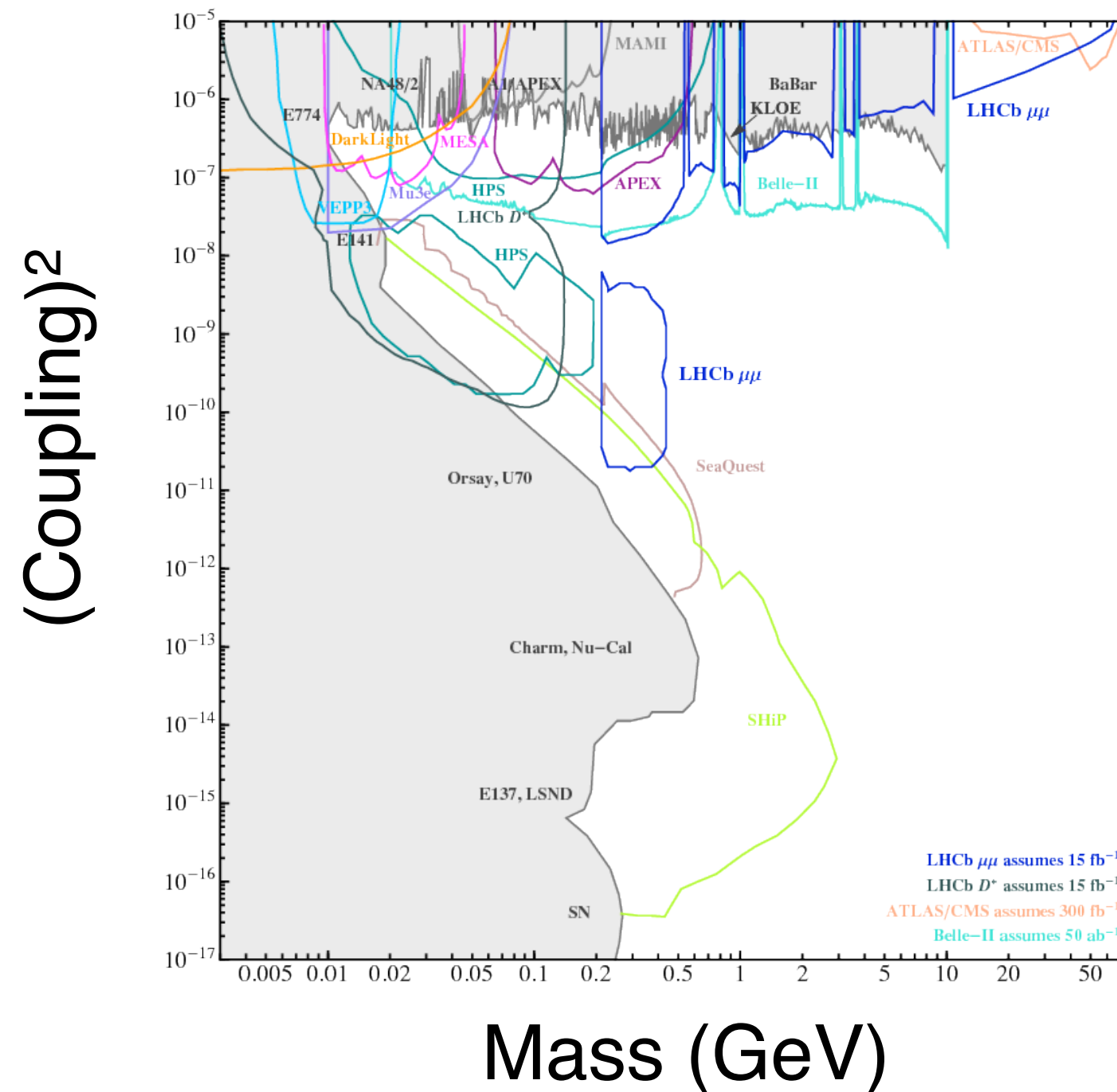
- DM may interact through new forces.  
*e.g.* new U(1) gauge boson  $\rightarrow$  “dark photon” (usually massive)
- A dark photon might decay to invisible dark particles.
- It could also decay visibly through kinetic mixing:

[Holdom 1986]



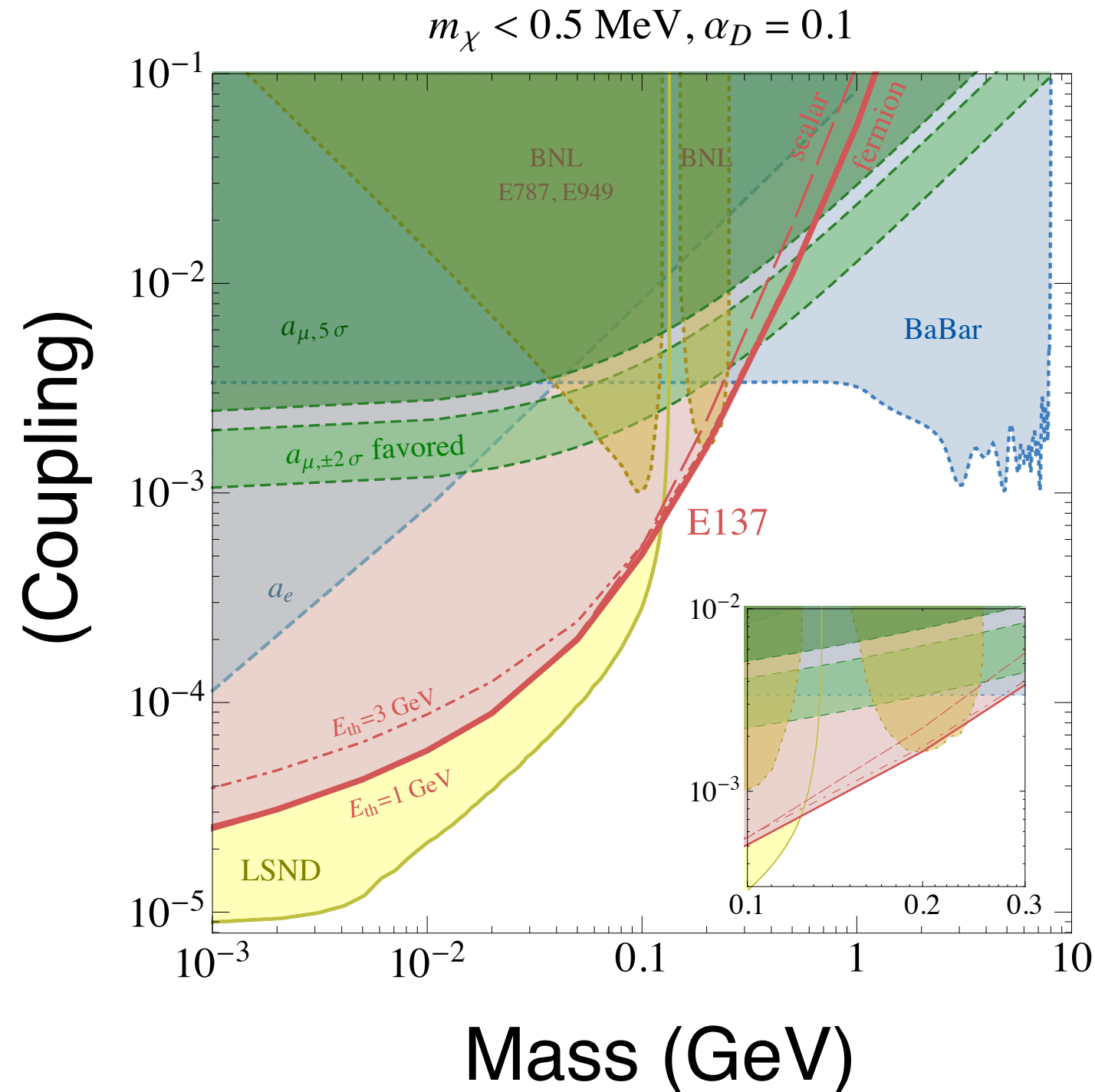


- Visible dark photon bounds:



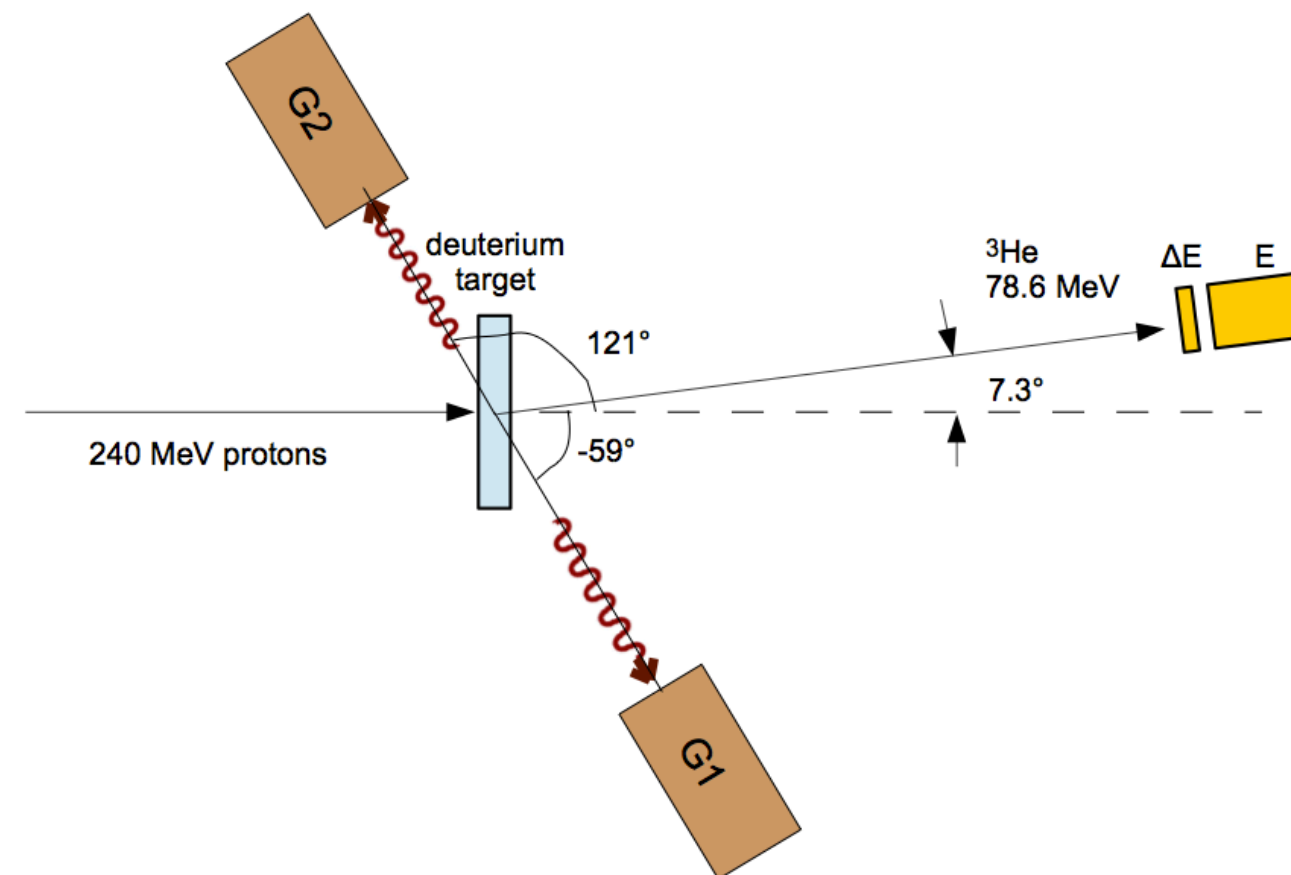
[Ilten, Soreq, Thaler, Williams, Xue 2016]

- Invisible dark photon bounds:



[Batell, Essig, Surujon 2014]

- Detection ideas:
  - dark photon from ARIEL beam dump (Doria)
  - invisible/long-lived dark photon in T2K/HyperK near detector (Ritz)
  - dark photon from neutral pion decay in  $p + d \rightarrow {}^3\text{He} + \pi^0$

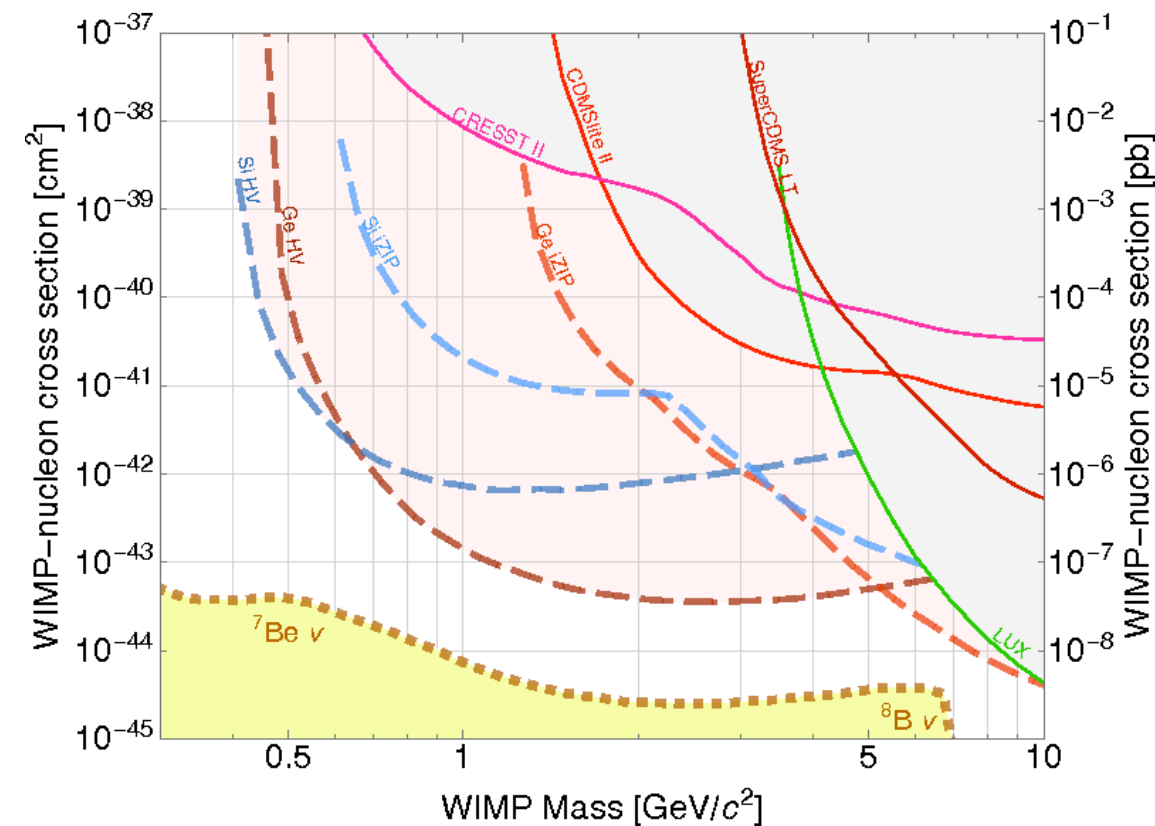


[Stan Yen]

- Most DM direct detection experiments lose sensitivity for mDM below about 10 GeV.

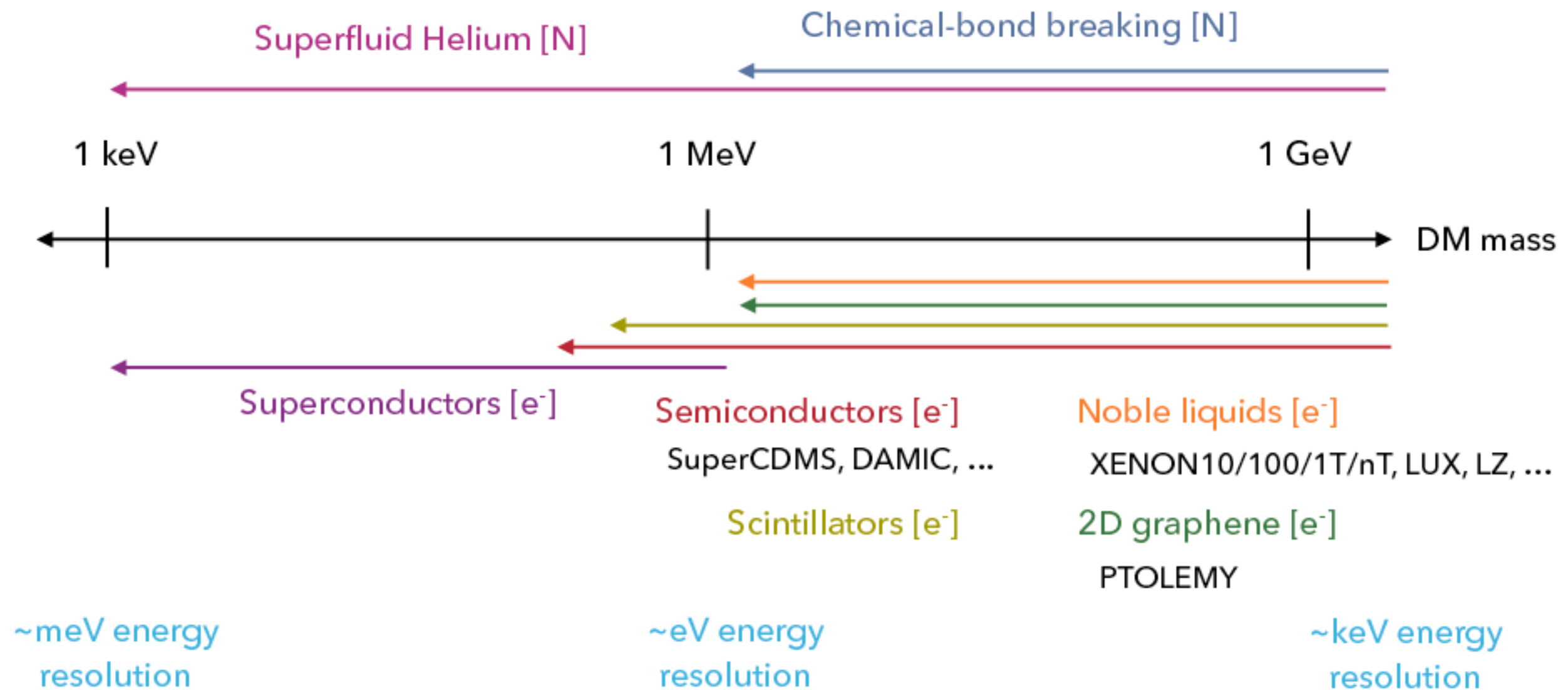
$$E_{recoil} \lesssim 10 \text{ keV} \left( \frac{m_{DM}}{10 \text{ GeV}} \right)$$

- SuperCDMS:



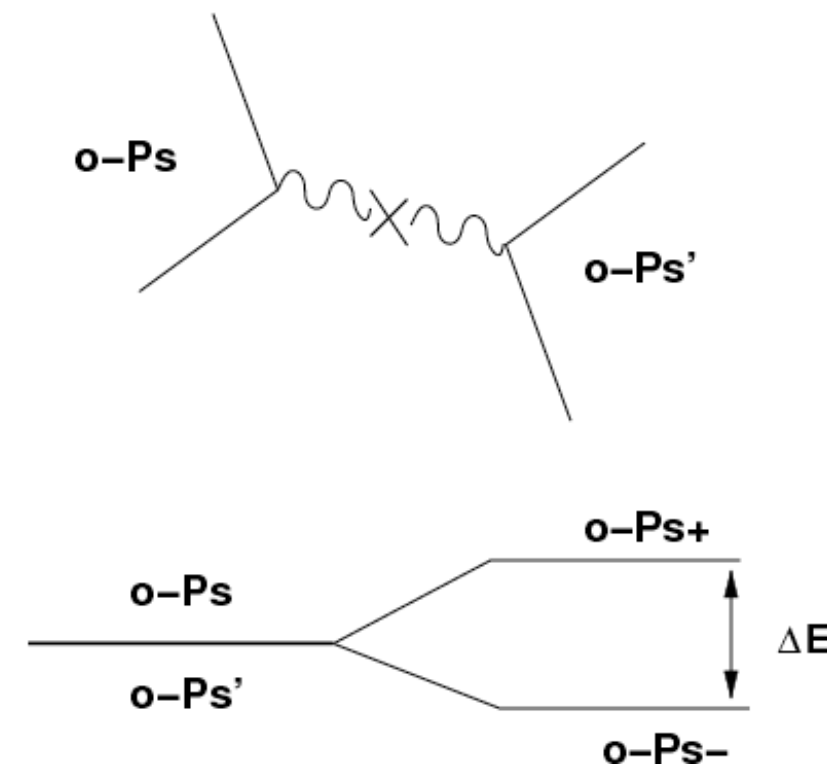
- Thermal DM can be as light at 5 keV, non-thermal even lighter!

- Sub-GeV DM searches need new detector technologies:



[Alexander *et al.*, Dark Sectors Community Report 2016]

- ARIEL (or cyclotron) could potentially make lots of positrons!  
Use to make positronium.
- Applications: [Makoto Fujiwara]
  - precision spectroscopy of positronium
  - test higher orders of QED through multi-photon decays
  - search for rare dark photon decays
  - test mixing with mirror positronium in mirror world scenarios



[Crivelli *et al.* 2010]

- LSND and Ga anomalies suggest a new neutrino mass scale near 1eV corresponding to oscillations over short distances.

$$\Delta (\text{phase}) \sim \left( \frac{\Delta m^2}{\text{eV}^2} \right) \left( \frac{L}{1 \text{ m}} \right) \left( \frac{1 \text{ MeV}}{E} \right)$$

$$\Delta m_{atm}^2 \simeq 2.5 \times 10^{-3} \text{ eV}^2 \qquad \Delta m_{sol}^2 \simeq 7.4 \times 10^{-5} \text{ eV}^2$$

- Minimal 3+1 approach: 3 SM-like neutrinos, 1 “sterile” neutrino.
- Idea: search for short distance oscillations within detector using neutrinos from  $^{12}\text{B}$  decays after production in ARIEL.  
(Similar to IsoDAR proposal based on  $^8\text{Li}$ .)

[John Behr]

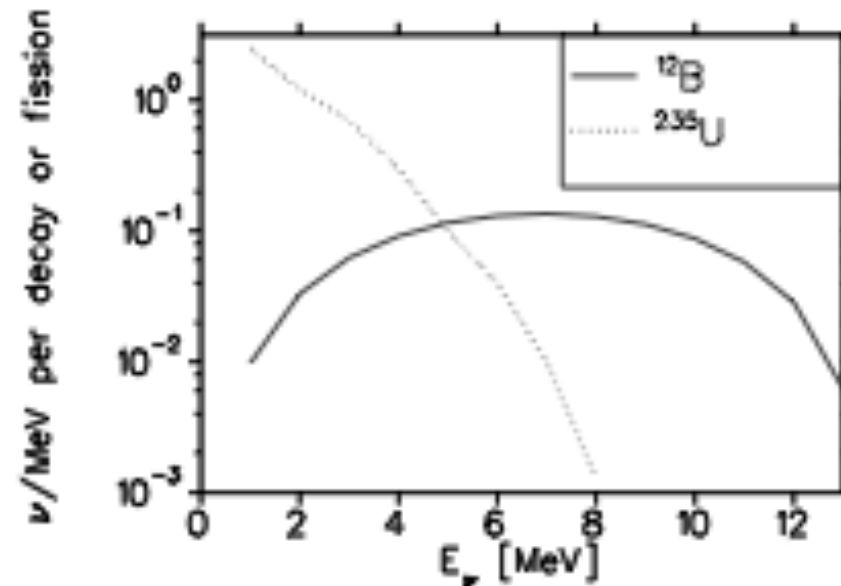
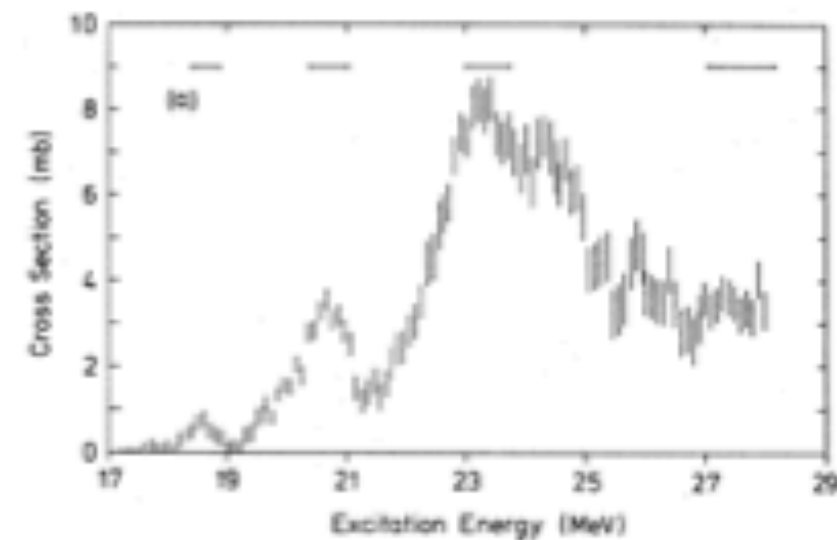
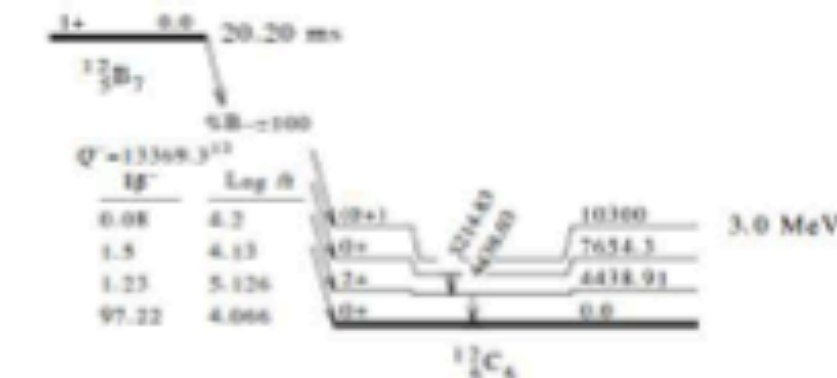
## TRIUMF In-target $\nu$ source with e-linac

ISOL extract 10% at best

→ in-target

$^{13}\text{C}(\gamma, p)^{12}\text{B}$

$t_{1/2}$ : chop 20 ms to exclude prompt bkg



$5 \times 10^{-4} \text{ }^{12}\text{B} / e^-$   
 ISODAR  $2 \times 10^{-3} \text{ }^8\text{Li} / p$   
 (ISODAR 7Li blanket → 10x)

35 MeV  $e^-$  3 mA →

$3 \times 10^{13} \bar{\nu} / s$

50 MeV  $e^-$  15 mA →

$3 \times 10^{14} \bar{\nu} / s$

Espinoza, Lazauskas, Volpe PRD (2012)

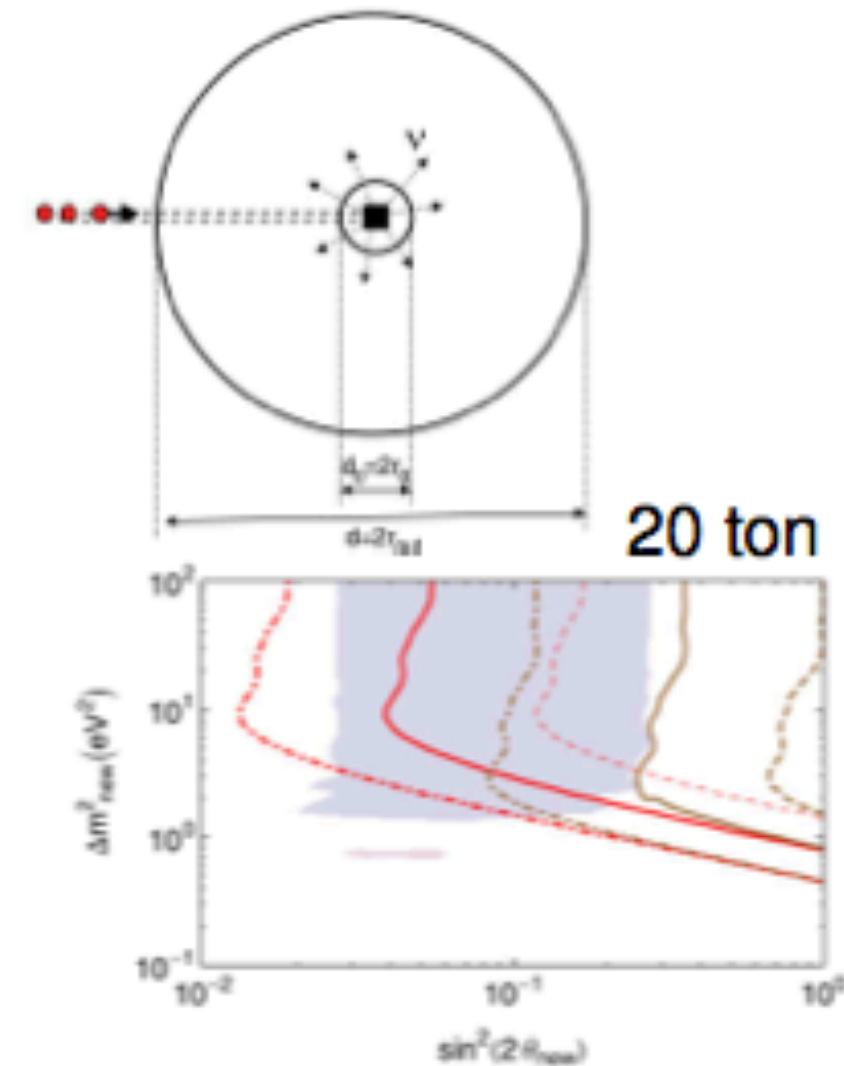


FIG. 6 (color online). Exclusion plots with binned analysis of the simulated data, obtained by varying the ion intensity:  $10^{14}$  ions/s (dash-dotted),  $10^{13}$  ions/s (solid), and  $10^{12}$  ions/s (dotted).

[John Behr - Monday]



- Understanding the Higgs boson is key to elementary particles!
- The Higgs portal opportunity:

1. Small decay width:

$$\Gamma(h_{SM}) = 4.1 \text{ MeV} \simeq (3 \times 10^{-5}) m_h$$

⇒ new decay modes can have significant branching ratios

2. Lowest dimension gauge-invariant operator:

$$\mathcal{L}_{eff} \supset (\text{BSM}) |H|^2$$

⇒ less suppressed compared to other connectors to BSM

- Rare Higgs decays realize this opportunity!

- Understanding the Higgs boson is key to elementary particles!

- Number of Higgs bosons produced: [Gori 2016]

<ul style="list-style-type: none"> <li>• LHC:</li> </ul>	$1.5 \times 10^7$	$\left\{ \begin{array}{l} 300 \text{ fb}^{-1} \\ 14 \text{ TeV} \end{array} \right\}$
	$1.5 \times 10^8$	$\left\{ \begin{array}{l} 3000 \text{ fb}^{-1} \\ 14 \text{ TeV} \end{array} \right\}$
<ul style="list-style-type: none"> <li>• ILC:</li> </ul>	$3 \times 10^5$	$\left\{ \begin{array}{l} 500 \text{ fb}^{-1} \\ 500 \text{ GeV} \end{array} \right\} \oplus \left\{ \begin{array}{l} 200 \text{ fb}^{-1} \\ 350 \text{ GeV} \end{array} \right\} \oplus \left\{ \begin{array}{l} 500 \text{ fb}^{-1} \\ 250 \text{ GeV} \end{array} \right\}$
	$2 \times 10^6$	$\left\{ \begin{array}{l} 4000 \text{ fb}^{-1} \\ 500 \text{ GeV} \end{array} \right\} \oplus \left\{ \begin{array}{l} 200 \text{ fb}^{-1} \\ 350 \text{ GeV} \end{array} \right\} \oplus \left\{ \begin{array}{l} 2000 \text{ fb}^{-1} \\ 250 \text{ GeV} \end{array} \right\}$

- LHC better for very clean channels, ILC better for others.



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Thank you!  
Merci!

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