QUEST-DMC: Probing Dark Matter with Nanowires, Superfluid Helium-3 and Quantum Sensors Paolo Franchini







GUINEAPIG 2023

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12 July 2023



200 days of French

Write this in French



Write this in English



Fill in the blanks Le cheval est gentil et les vaches sont gentilles



Outline

- The dark matter hypothesis
- Direct dark matter detection
- QUEST-DMC dark matter search programme
- Bolometry in Helium-3 with nanowires
- Measurement of the deposited energy
- Estimated sensitivity
- Future prospects



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"I was conscious that I knew practically nothing..."



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Socrates: 0 books

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Dark matter hep-th: 1300 papers on arxiv

Direct dark matter detection: candidates



Direct dark matter detection: candidates



The dark matter hypothesis

- Evidences from gravitational interactions
 - Spiral and elliptical galaxies: rotation curves
 - Galaxy clusters: gravitational lensing
 - Cosmic Microwave Background
 - Colliding galaxy clusters: x-rays

- No electromagnetic interaction ("dark")
- Stable over billion years
- Non-baryonic: an unknown particle?





The dark matter hypothesis

• Matter and energy content of the universe

• Dark matter is 85% of mass

• Philosophical implications



Paradigm shift?



Copernican revolution



Direct dark matter detection: WIMPs

Nucleus

Μ

WIMP

- Signal: scattering of DM candidates off a target
 - rare processes...
 - ... but many WIMPs
- Backgrounds:
 - Cosmic rays
 - Radioactive environment
 - Radioactive contamination
 - Neutrinos



Recoil energy

~M/1'000'000

Direct dark matter detection: WIMPs



Rep. Prog. Phys. 85 (2022) 056201

https://arxiv.org/abs/2207.07640

QUEST-DMC collaboration and ecosystem



QUEST-DMC programme

- Beyond Standard Model physics investigation
 - Quantum sensors
 - Helium at ultra-low temperatures



1) What is the nature of Dark Matter?

Detection of sub-GeV dark matter with a quantum-amplified superfluid He3 calorimeter

2) How did the early universe evolve?

Phase transition in extreme matter ↔ early universe









- Spin dipendent WIMP-nucleon cross section: 10⁻³⁷ cm² @ 1 GeV/c²
- eV recoil energy threshold
- background < 1 event/kg/day/keV_{DEP}





Helium-3

- Spin ¹/₂ (!)
- Superfluid (1972) below 2.5 mK
- P-wave pairing and multiple superfluid states
- ³He-B:

- Solid (bcc) Pressure (MPa) 3 Superfluid A phase 2 Superfluid B phase Normal liquid Gas 0.0001 0.001 0.01 0.1 10 100 Temperature (K)
- He3 as a fermionic condensate (similar to BSC theory)
- Cooper pairs: composite bosons, 100nm size
- Pair of bound quasiparticles with 10⁻⁷ eV energy and an effective mass

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Target



Wet dilution refrigerator





- Still (He4)
 - Heat exchangers
 - **Mixing chamber** •

Mixing

Still

Heat Switch

- He3 concentrate and diluted phases
 - Dilution process absorbs heat
 - Diluted He3 back up in the still
 - Cooling of concentrated in the way



Target

Nuclear demagnetisation refrigeration



Dark Matter events in He3

1.0

0.8

fraction 6.0

0.2

- Collision WIMP-He3 atom
 - Heat: quasiparticle excitations (10⁷/eV)
 - Light: from de-excitation



Nuclear Recoil

triple



Vibrating nanowire

Bolometer response

Detector

27

- Wire oscillating in magnetic field in a He3 cm³ box
- Damping force on the oscillator, enhanced by Andreev reflections
- Voltage response
- Measure **energy deposition** as variation of the resonance width ∆f



Lancaster nanowires

- 1. 200 um copper matrix with 1 um Nb-Ti wires
- 2. Draw the cable in multiple dies
- 3. Etch the copper and replace with water
- 4. Microscope + tweezers
- 5. Replace water with IPA
- 6. Let dry

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7. Mount on a PCB









Lancaster nanowires



1/500 of a cat's whisker

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Detector

Cells with wires





Quantum sensor

Bolometer in He3

• Deposited energy as variation of the damping force on the resonator



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Readout

Lock-in amplifier

- Extract signal (in a defined frequency band) from a noisy background using a reference signal (RMS noise ~ 10 nV)
 - Amplitude
 - Phase



Readout

DC SQUID

Readout

- Superconducting QUantum Interference Device
- Magnetometer, 10⁻¹⁴ T (brain: 10⁻¹³T)
- Magnetic flux into electrical voltage



- Voltage drive applied inductively
- Wire has Z(w) impedance
- Output current I,
- Resulting flux read by SQUID
- Output voltage **V**



Readout sensitivity

• Error on the energy measurement ↔ DM energy threshold



Conventional: 39 eV SQUID: 0.71 eV

Readout



Background

- Cosmic rays
- Radioactive decays
 - Environment
 - Contamination
 - Materials

Alphas Betas Gammas Neutrons Muons

- Neutrinos (irreducible)
 - → Goal: 1 background event/kg/day

Background: radiogenic

- Simulation of the decays
 - Energy deposited in the cm³ He³ cell

- Estimated activity
 - Radiopurity database
 - Screening of the materials



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Background: simulation



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Background: screening

Boulby Undeground Laboratory (UK) screening measurements:

Material	Up 238 U	Lower ^{238}U	²¹⁰ Pb	Upper 232 Th	Lower 232 Th	$^{235}\mathrm{U}$	$^{137}\mathrm{Cs}$	⁴⁰ K	60 Co	^{54}Mn
Concrete	$< 1.60 \times 10^5$	1.50×10^{4}	1.00×10^7	7.57×10^{3}	7.57×10^{3}	$<7.20{\times}10^3$	800	4.20×10^{4}	< 700	0.00
Aluminium	8.33×10^{3}	15.3	70.7	356	334	60.5	< 0.940	< 3.12	< 1.10	0.00
Superinsulation	679	< 200	$< 3.90 \times 10^{3}$	200	200	4.93	0.00	3.50×10^{3}	400	0
Stainless Steel	16	2.5	82.2	3.1	3.90	0.120	2.00	< 6.20	< 5.20	1.70
Steel	< 12.4	12	1.20×10^{4}	4.88	4.88	3.00	2.00	34.1	30.0	1.00
Araldite	< 3.60	< 4.80	14.5	< 3.40	< 2.20	0.0260	2.00	25.5	8.00	0.00
Stycast	< 10.5	< 9.50	< 14.9	< 12.8	< 6.20	0.0762	2.00	122	10.0	0.00

• Current setup, estimated radiogenic activities:

Component	Expected cor	unts [0-10 keV]	Uncertainty
	/kg/day	/cell/day	
Cosmic ray	1.05×10^5	3.31	$11 \ \%$
Radiogenic ER	8.31×10^4	2.61	14~%
Solar ν ER	1.51×10^{-2}	4.76×10^{-7}	2~%
Solar ν NR	6.37×10^{-4}	2.01×10^{-9}	2~%
TOTAL	1.88×10^5	5.92	



Bolometers in He3



• Bolometer operation limitation



Bolometer in He3: events



Extract:

• Rate of background events

DATA!

- Energy spectrum
- Energy threshold



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preliminary **Sensitivity projection** Signal simulation **Sensitivity for Detector** dark matter response model Background simulation 10⁵ 10¹ WIMP mass [GeV/c2] - 0.01 103 0.05 10-1 - 0.1 Rate [events/kg/day] Rate [events/kg/day] - 0.5 cosmic ray — 10 radiogenic — 5.0 neutrino NR 10-3 preliminary ---- 10.0 neutrino ER —— SD WIMP 10-5 10-3 arXiv:2106.06207 10^{-7} 10⁻⁵ 10^{-2} 10-1 10⁰ 10-3 10¹ 10² 10³ 10^{-2} 100 10¹ 10^{-1} 10² Energy [keV] Energy [keV]

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preliminary **Sensitivity projection** Signal simulation Detector Sensitivity for dark matter response model Background simulation 10³ 10¹ [%] 10⁻¹ oreliminary 10-3 Lock-in amplifier 10^{-5} SQUID readout ••••• OP shot noise 10^{-1} 101 10³ 10⁵ Energy [eV]



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lacksquare

Conclusion and outlook

- Simulation and analysis machinery in place
- Produced a first sensitivity limit, based on actual constructed detector cells and modelled energy reconstruction validated on data (<u>submitted to Nature</u>)
- Next:
 - Compare the model with first round of data
 - Implement light detection in the cell
 - Add cosmic rays tagging

Conclusion and outlook







- SQUID operated at RHUL at LN temperatures
- To be operated at uK tempeatures in Lancaster

Conclusion and outlook

- Work towards SQUID readout of a nanowire
- Start operating He3 cells with nanowires





Great potential for quantum technologies to open up a new window on the dark matter universe

