





Progress Toward Low-Mass Dark Matter Detection with Superfluid He (HeRALD) and Polar Crystals (SPICE)

David Osterman, on behalf of the SPICE/HeRALD Collaboration GUINEAPIG July 13, 2023



Outline



- 1. Intro to the collaboration and its goals
- 2. TES calorimetry and energy threshold
- 3. HeRALD description
- 4. SPICE description
- 5. Understanding and mitigating low energy backgrounds

We are looking for light DM SPICE HERALD

Lower $M_{DM}^{} \rightarrow higher \, n_{DM}^{} \rightarrow less \, need$ for high exposure

Traditional radiogenic backgrounds are less important \rightarrow can build cheap experiments quickly

Different target materials probe different models

Can produce world-leading limits by operating gram-scale detectors with different target materials and days-long exposure



[1] Snowmass 2021 (arXiv:2209.07426)

The SPICE/HeRALD collaboration





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Two goals of both experiments

SP/CE HeRALD



1. Good energy threshold

QET: Quasiparticle-trap-assisted Electrothermal-feedback Transition-Edge-Sensor



Two goals of both experiments SPICE

1. Good energy threshold

2. Low backgrounds



[2] Adari, Prakruth, et al. "EXCESS workshop: Descriptions of rising low-energy spectra." *SciPost Physics Proceedings* 9 (2022): 001.





TES calorimetry and energy threshold

Transition Edge Sensors (TESs)



Superconducting metal (typically tungsten) biased at T_0 slightly less than the transition temperature s.t. $R_0 = R_N/3$

Small E deposit causes small T increase which causes sharp R increase

Electrothermal feedback makes the temperature stable

V-biased \rightarrow Joule heating = V²/R





TES

Al fin

Al fins funnel energy to the TES

QET: Quasiparticle-trap-assisted Electrothermal-feedback Transition-Edge-Sensor





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Optimizing AI energy efficiency



Athermal phonons in the Si break Cooper pairs, forming quasiparticles (QPs)

QPs diffuse to the TES, which measures the change in R(T)



[3] Kurinsky, Noah. *The low-mass limit: Dark matter detectors with eV-scale energy resolution*. Stanford University, 2018.

Optimizing AI energy efficiency

Athermal phonons in the Si break Cooper pairs, forming quasiparticles (QPs)

QPs diffuse to the TES, which measures the change in R(T)

QPs cannot diffuse backward from lower energy to higher energy





Figure 6-3: A depiction of the physical structures of the quasiparticle gap (bottom) and their resulting quasiparticle energies (top). Quasiparticles in the aluminum film gradually diffuse to regions of lower gap energy (right) and become trapped.

[4] Hertel, Scott A. "Advancing the search for dark matter: from CDMS II to SuperCDMS." *Ph. D. Thesis* (2012).ç

Optimizing AI energy efficiency

QP diffusion is critical, but there is a tradeoff: higher AI surface coverage means

- Higher efficiency of capturing athermal phonons from the substrate
- QPs need to diffuse longer distances, and are more likely to be trapped on imperfections in the AI, decreasing the overall energy efficiency

Need to optimize the AI fin length according to the characteristic QP diffusion length (ℓ_d) in AI





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"Banana" experiments to measure QP diffusion



[4] Hertel, Scott A. "Advancing the search for dark matter: from CDMS II to SuperCDMS." *Ph. D. Thesis* (2012).

2023 HeRALD banana experiment (laser)

150nm thick AI fins fabbed at ANL Preliminary measurements of $\ell_{d} \sim 500$ um \rightarrow Longer AI fins?

→Improvement in active AI coverage?



vacuur

helium

800 1000

Al fin

Fits to sweep down center of fin - vacuum vs. He

Preliminary



100

integral

bul

normalized

g 10⁻¹

 10^{-2}

-400 -200



This material is based upon

work supported by the U.S. Department of Energy, Office of

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Optimizing TES energy resolution



Energy resolution scales strongly with $\rm T_{c}$ and TES volume

Current state-of-the-art $\rm T_c$ for tungsten TES is 40mK

 $T_c = 15mK W TES$ has been demonstrated $\rightarrow 19x$ improvement in resolution?

Can also decrease the TES volume for further improvement

$$\sigma_E \stackrel{\sim}{\propto} T_c^3 \sqrt{rac{V_{TES}V_{crystal}}{l_{fin}}}$$



The HeRALD Experiment

Advantages of superfluid 4He

Low mass makes 4He sensitive to energy deposits down to ~1 meV

$$T_{\rm NR} \leq \frac{2m_\chi^2 m_N v_\chi^2}{(m_\chi + m_N)^2}$$

0.62 meV to quantum evaporate He atom, 10meV absorption \rightarrow x10 gain







Advantages of superfluid He

- 1: 1 kg-day, 40eV threshold
- **2:** 1 kg-yr, 10eV threshold
- 3: 10 kg-yr, 100meV threshold
- 4: 100 kg-yr, 1meV threshold
- **4':** 100 kg-yr, 1meV threshold, off-shell phonon processes





[5] Hertel, Scott A., et al. "Direct detection of sub-GeV dark matter using a superfluid He 4 target." *Physical Review D* 100.9 (2019): 092007.

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HeRALD v0.1













The cryogenic photodetector (CPD)

The sensor (CPDv2) is a 3" diameter, 1mm thick, 10g Si wafer

~700 tungsten Transition Edge Sensors (TESs) with $T_c = 51 mK$

~2.3eV resolution (σ) for energy in Si





Target He volume with single sensor suspended above

Cesium stops superfluid He from creeping onto the sensor

High current (>7A) dispensers deposit Cs onto sensor post

Baffles separate Cs region from sensor region















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Three signal channels

16eV photons - from quickly (10ns) decaying He_2^* singlet excimers

 He_2^* triplet excimers - propagate ballistically at 1m/s; long-lived (13s halflife) → quench on walls

Quasiparticles - phonons and rotons, the excitations of superfluid He

calorimeters



Initial HeRALD R&D data





S1: singlet scintillation photons S2: He evaporation some Δt later

He thickness @(1cm) v_{phonon} @(100m/s)

→ Δt 𝖉(100us)



Initial HeRALD R&D data





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Initial HeRALD R&D data

Two clear populations from calibration Xray energies:

- 5.9keV (55Fe)
- 1.5keV (AI fluorescence)

Scintillation:

Matches expectation from light yield and solid angle

Evaporation:

Roughly matches expectation, but lots more to study!



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Initial HeRALD R&D data

$\Delta t \ delay \ before \ evaporation \ signal \ is \\ depth-dependent$

As expected: 55Fe depth spread is ~uniform Al is only near the bottom





Initial HeRALD R&D data





Intended signal region <20eV (only QPs), but we can still look at ER/NR above 20eV

Preliminary 252Cf observations: Higher evaporation:scintillation ratio Larger triplet fraction



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Initial HeRALD R&D paper by HD Pinckney should be on arXiv next week!



The SPICE Experiment

SPICE targets and their motivations



Sapphire (AI_2O_3)

- Polar crystal \rightarrow optical phonons (very low effective mass)
- Oxygen nuclei are light \rightarrow good kinematic matching to light DM

 SiO_2

- Polar crystal
- Great dark photon coupling

GaAs

- Polar crystal
- Scintillation and phonon signals \rightarrow NR/ER discrimination down to very low energies

Sapphire (Al₂O₃) target

Low mass oxygen nuclei good NR scattering targets

Polar unit cell \rightarrow optical phonons down to ~100meV

Prototype detector (below) demonstrates we can make devices out of this material









SiO₂ target

Great coupling to dark photons, high quality factor (see arXiv: 1910.10716)

TESs on SiO₂ substrates (right) work!







GaAs target

Scintillation and phonon signal \rightarrow ER/NR discrimination down to eV-scale signals

GaAs scintillation has been observed (from device below)









Understanding and mitigating low-energy backgrounds

The low-energy excess



The primary background in all phonon-based experiments at eV scales



The low-energy excess



The primary background in all phonon-based experiments at eV scales



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SPICE stress studies: purposefully creating SPICE HeRALD Stress events

Two identical TES-based athermal phonon detectors - 1cm² by 1mm thick silicon substrate

One **glued** down to Cu substrate (high stress), one **suspended** from wire bonds (low stress)



High stress



Low stress

SPICE stress studies: purposefully creating stress events



SPICE HeRALD

SPICE stress studies: What could be causing stress events?

Tungsten film stress? (upper right) -Plot is from chip with two TESs and the rest of the surface covered with Al to remove substrate phonons -From stress-induced microfractures in W? -Can be mitigated by moving to other TES materials: IrPt, AIMn, etc.

Al film stress? (lower right) -Strong evidence for Al deformation over time at low temperatures -Point of plot: cooling Al caused the release of some stress



SPICE



HeRALD LEE mitigation effort: coincidence discrimination

- Goal: use two detector channels (separate wafers) to eliminate this excess by coincidence
 - -An event in the He will be visible in both channels
 - -A spontaneous phonon event in a photodetector will show up only in that one channel



SPCE

He RALD

HeRALD v0.1 - 2-channel runs happening soon! SPICE HERALD



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HeRALD v0.2 - 4-channels!







Summary



HeRALD:

- Demonstrated superfluid He film stopping with Cs, and measured scintillation, evaporation, and triplet events in superfluid 4He target (on arXiv next week!)
- Measured improved QP diffusion in AI fins (paper in production)
- Next: Discriminate LEE stress events with coincidence

SPICE

- Demonstrated sapphire, GaAs, and SiO₂ target materials
- Measured LEE-type events from stress
- Next: Eliminate LEE stress events with different TES materials



Extra slides

Recent dark matter exclusions for direct detection SPICE HERALD



Figure 7: Combined Spin-independent dark-matter nucleon scattering cross section space. Currently 90% c.l. constraints are shown in shaded beige [240–255] (data points taken from [239]) while the reach of currently operating experiments are shown in green (LZ, XENONnT, PandaX-4T, SuperCDMS SNOLAB, SBC). The limits from 2013 are shown as a gray line [256]. Future experiments are shown in blue (SuperCDMS, DarkSide-LowMass, SBC, 1000 ton-year liquid xenon, ARGO) and yellow (Snowball and Planned \times 5). The neutrino fog for a xenon target is presented in gray as described in Sec. 4.3.2. Plot reproduced from Ref. [2].

[4] Jodi Cooley, Tongyan Lin, W Hugh Lippincott, Tracy R Slatyer, Tien-Tien Yu, Daniel S Akerib, Tsuguo Aramaki, Daniel Baxter, Torsten Bringmann, Ray Bunker, et al. Report of the topical group on particle dark matter for snowmass 2021.arXiv preprint arXiv:2209.07426, 2022.

Different models and materials SPICE HERALD (





| Light dark photon mediator (Sec. III, Fig. 1) | | | | |
|---|---------------------------------------|--|--------------------------------|--|
| Detection channel | Quantity to maximize to reach | | Best materials | |
| | lower m_{χ} | lower $\overline{\sigma}_e$ | Dest materials | |
| (Optical) phonons | ω_O^{-1} (Eq. (24)) | quality factor Q defined in Eq. (27) | SiO_2 , Al_2O_3 , $CaWO_4$ | |
| Electron transitions | E_g^{-1} (Eq. (28)) | depends on details of electron wavefunctions | InSb, Si | |
| Nuclear recoils | $(A\omega_{\min})^{-1}$ (Eq. (29)) | $(Z/A)^2 \omega_{\min}^{-1}$ (Eq. (31)) | diamond, LiF | |
| Hadrophilic scalar mediator (Sec. IV, Figs. 2, 3) | | | | |
| Detection channel | Quantity to maximize to reach | | Best materials | |
| | lower m_χ | lower $\overline{\sigma}_n$ | Dest materials | |
| (Acoustic) phonons | $c_s/\omega_{ m min}~({ m Eq.}~(36))$ | Light mediator: ω_{\min}^{-1} (Eq. (35)) | diamond SiO_2 | |
| | | Heavy mediator: c_s^{-1} or $\omega_{\rm ph}^{-1}$ or $A\omega_{\rm ph}$ | all complementary | |
| | | depending on m_{χ} (Eqs. (37), (38), (39)) | | |
| Nuclear recoils | $(A\omega_{\min})^{-1}$ (Eq. (29)) | Light mediator: ω_{\min}^{-1} (Eq. (40)) | diamond, LiF | |
| | | Heavy mediator: A (Eq. (43)) | CsI, Pb compounds | |

arXiv: 1910.10716 Griffin et. al. 2021

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The inspiration for HeRALD

Inspiration comes from HERON (right), a detector that used superfluid He as a target to detect neutrinos, a very light particle

HeRALD uses the same principle: a light target (He atoms) to detect a light particle (light dark matter)





[5] J.S. Adams. The heron project. Low Energy Solar Neutrino Detection, 2001.

Quantum evaporation from phonons/rotons

Phonons and rotons are the excitations of superfluid He





Quantum evaporation from phonons/rotons

Phonons and rotons are the excitations of superfluid He

Kapitza resistance causes reflection at solid interfaces

He atom quantum evaporation at superfluid surface within a certain incident angle







CPDv3 features

- channel 1
- channel 2; wafer edge
- electrical wire bonds
- - QETs in horizontal rows
- QETs around CPD half edges
- gold thermalization bond pads
- hanging bond pads
- heater QETs



map of features



CPDv3 specs vs CPDv2



| | CPDv2 (wal-e) | CPDv3 (split) |
|---|---------------------|-------------------------------|
| Al lead (on each QET) width | ?? (something >4um) | 6um |
| N _{QET} | 673 | 678 (339 per half) |
| R _{N,eff} (all TESs in parallel) | 200mΩ | 397m Ω (for each half) |
| Active surface coverage | 0.68% | 0.67% |
| Passive surface coverage | 0.18% | 0.17% |

CPDv3 features

678 QETs (339 per half) - same QET specs as CPDv2

1 gold bond pad per half - for thermalization

1 "heater QET" per half - for heating and also as a second signal channel if necessary

3 hanging bond pads per half - if we wish to hang the CPD with 2 mil Al wire bonds

76.2mm diameter 1mm thick wafer with 22mm flat edge at top and 11mm flat edge at bottom

Note: Substrate in right image is turquoise. The black outline outside the substrate is a guiding feature on the Al and W masks to help with alignment. This guiding feature will appear in the respective colors of the individual mask design layers.







CPDv3 vs CPDv2

v3







Background rates for exclusion curves



SPICE HeRALD

[12] Hertel, Scott A., et al. "Direct detection of sub-GeV dark matter using a superfluid He 4 target." *Physical Review D* 100.9 (2019): 092007.

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