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Mathematisch-Naturwissenschaftlich-
Technische Fakultät

Determining absolute VUV fluxes for assessing the relevance of photon-surface interaction in ion sources

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20th ICIS, 17.-22.09.2023, Victoria, Canada

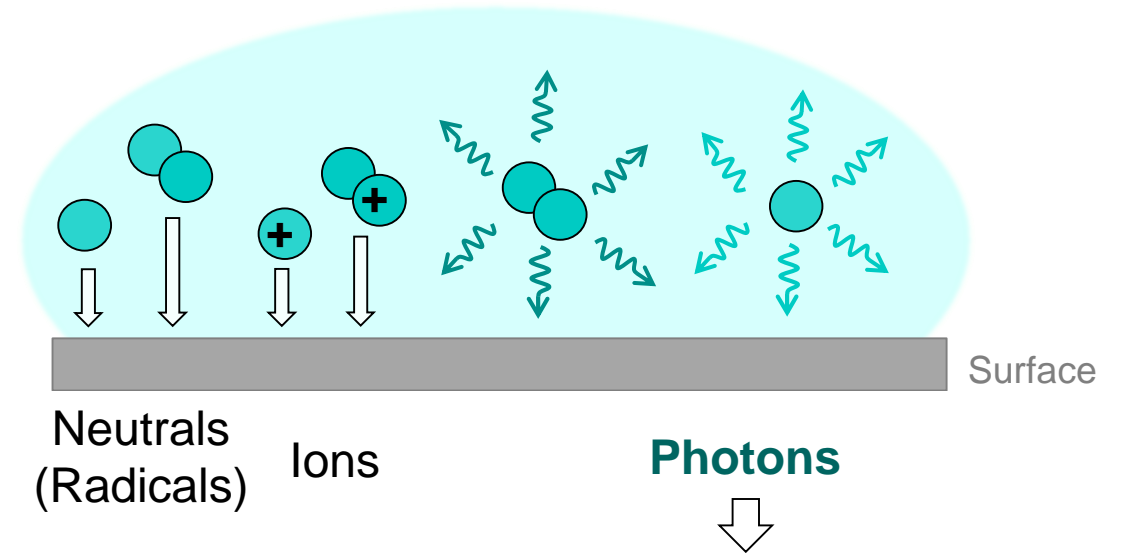
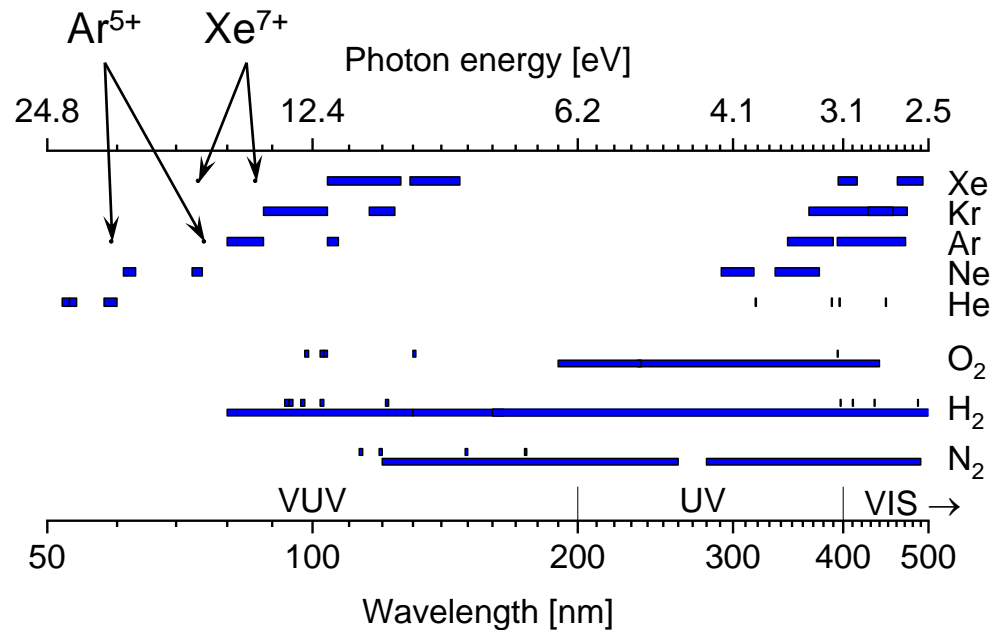


DFG Deutsche
Forschungsgemeinschaft

Funded by the Deutsche Forschungsgemeinschaft (DFG, German Research Foundation) – Project FR 3881/1-1.

VUV emission in ion sources

- Resonance transitions of neutrals and ions
- Insight into basic plasma physics
- Energetic fluxes towards plasma-facing components

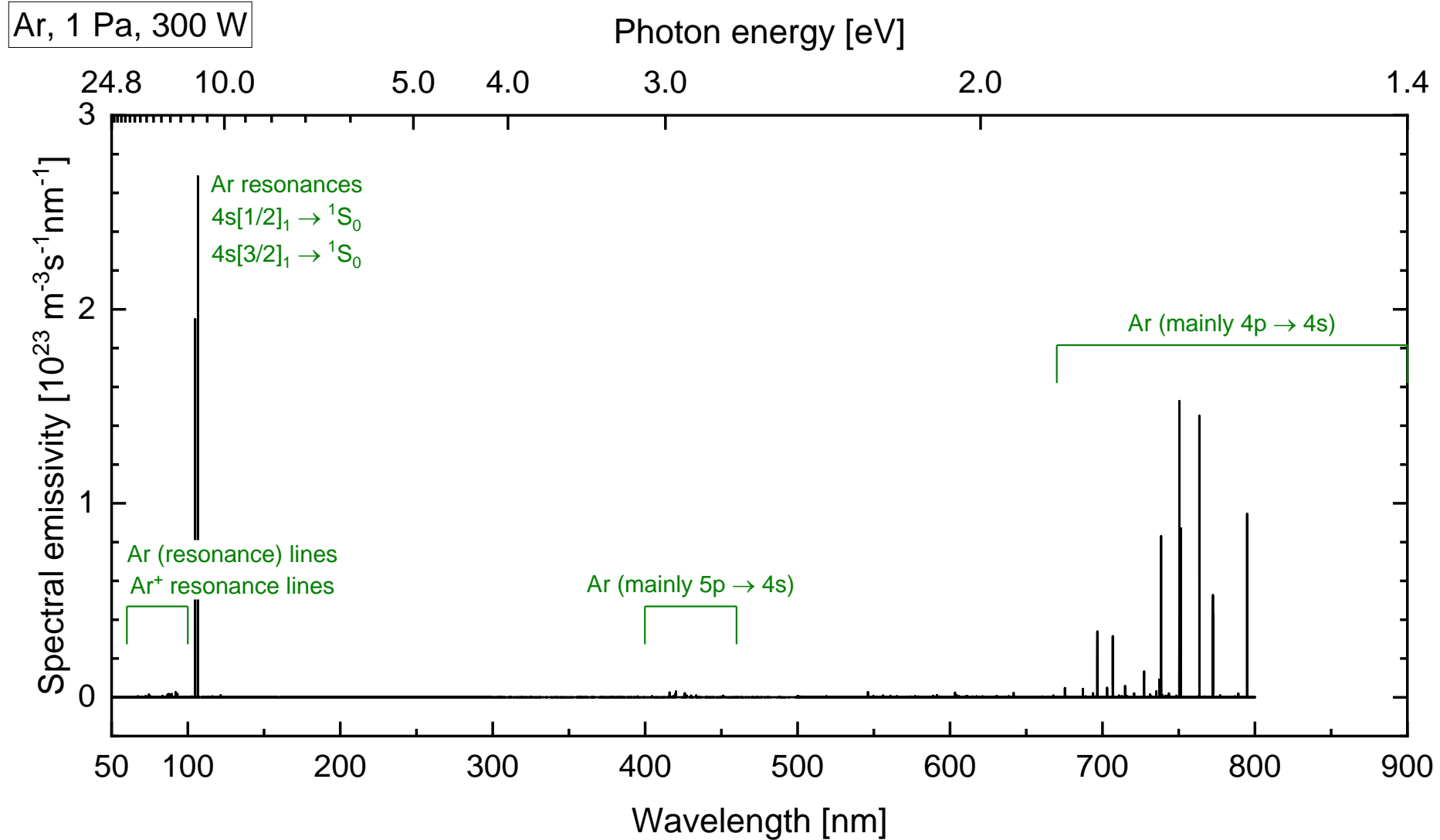


Impact depends on
photon energy and absolute flux

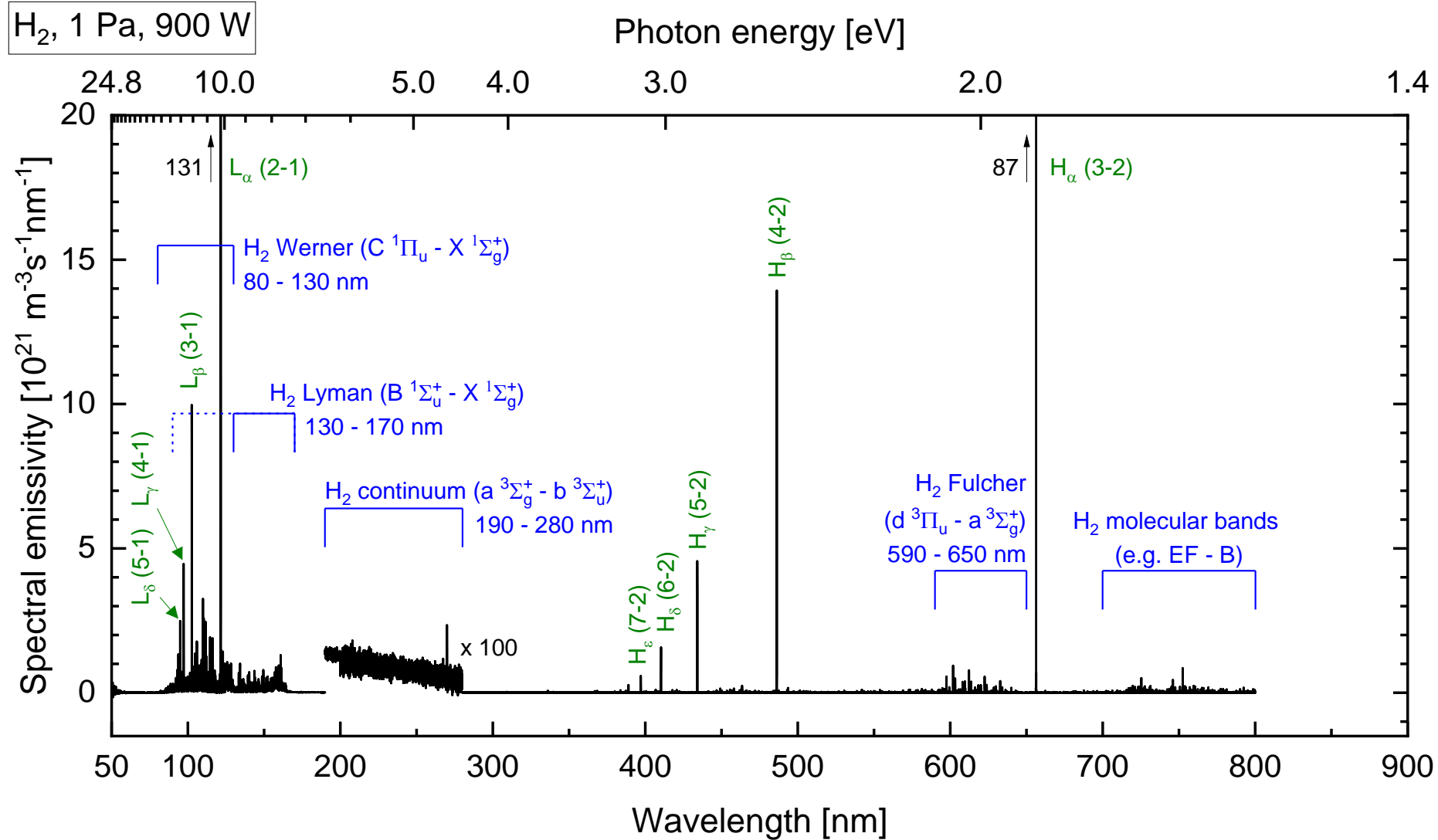
Challenges for VUV spectroscopy:

Direct vacuum connection +
Complex calibration procedure

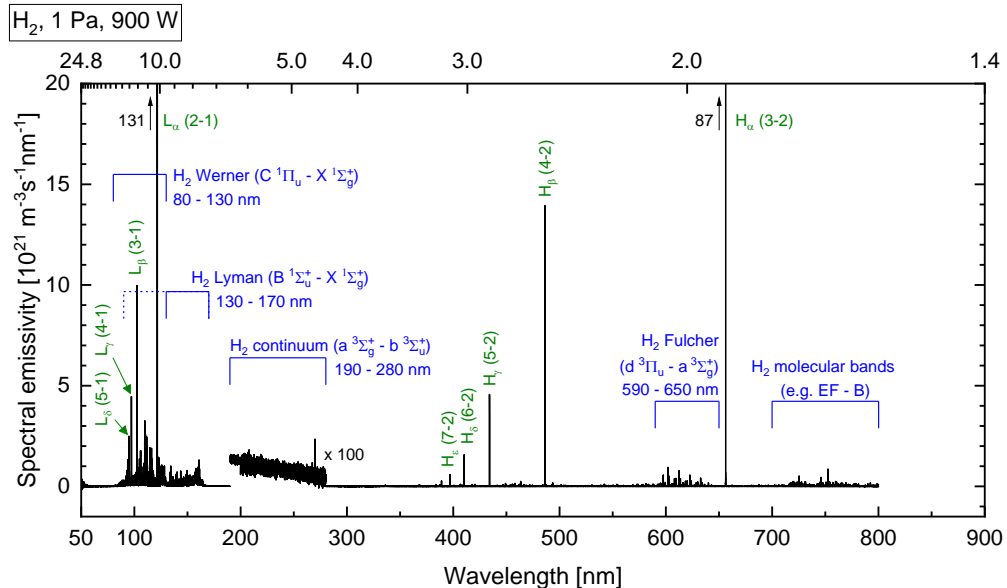
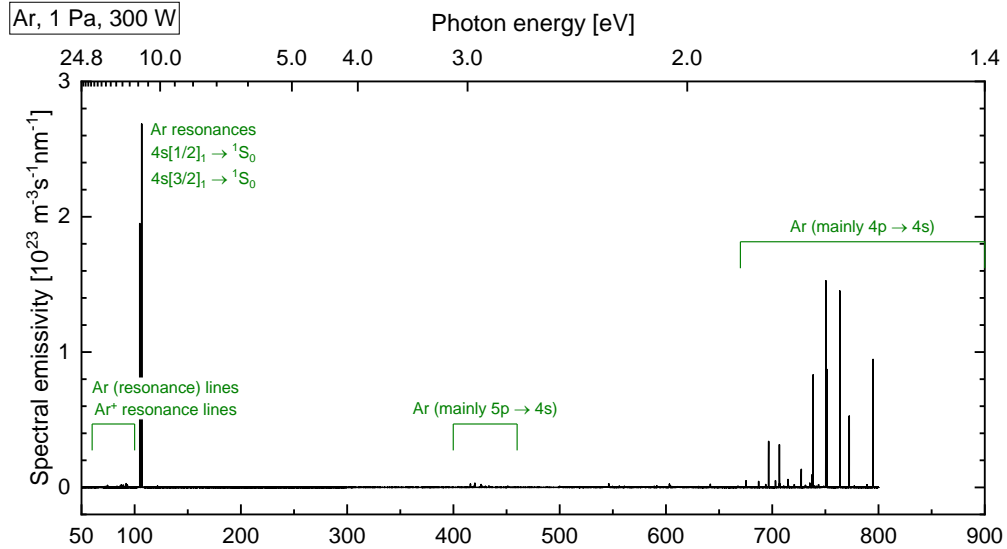
Detecting VUV emission



Detecting VUV emission



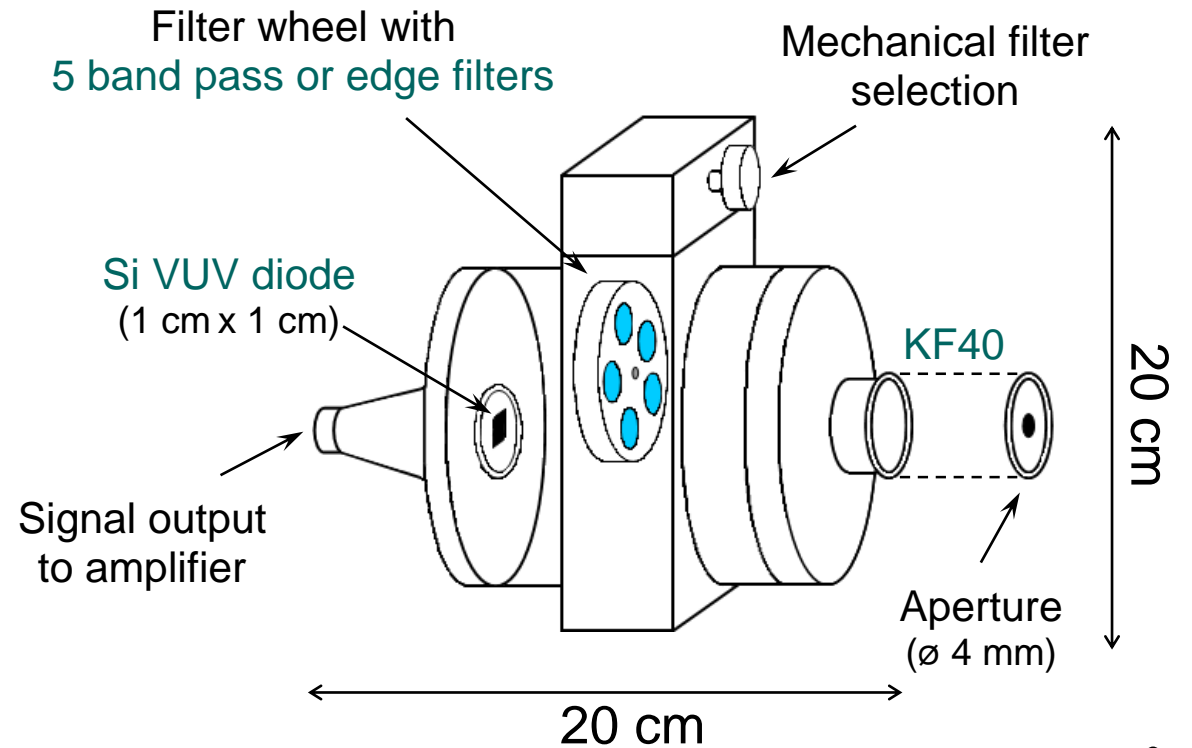
Detecting VUV emission



VUV diode & optical filters

- Simple setup
- Portable & easily attachable
- No separate pumping required

KOMPPULA & TARVAINEN
 PSST 24 (2015) 045008

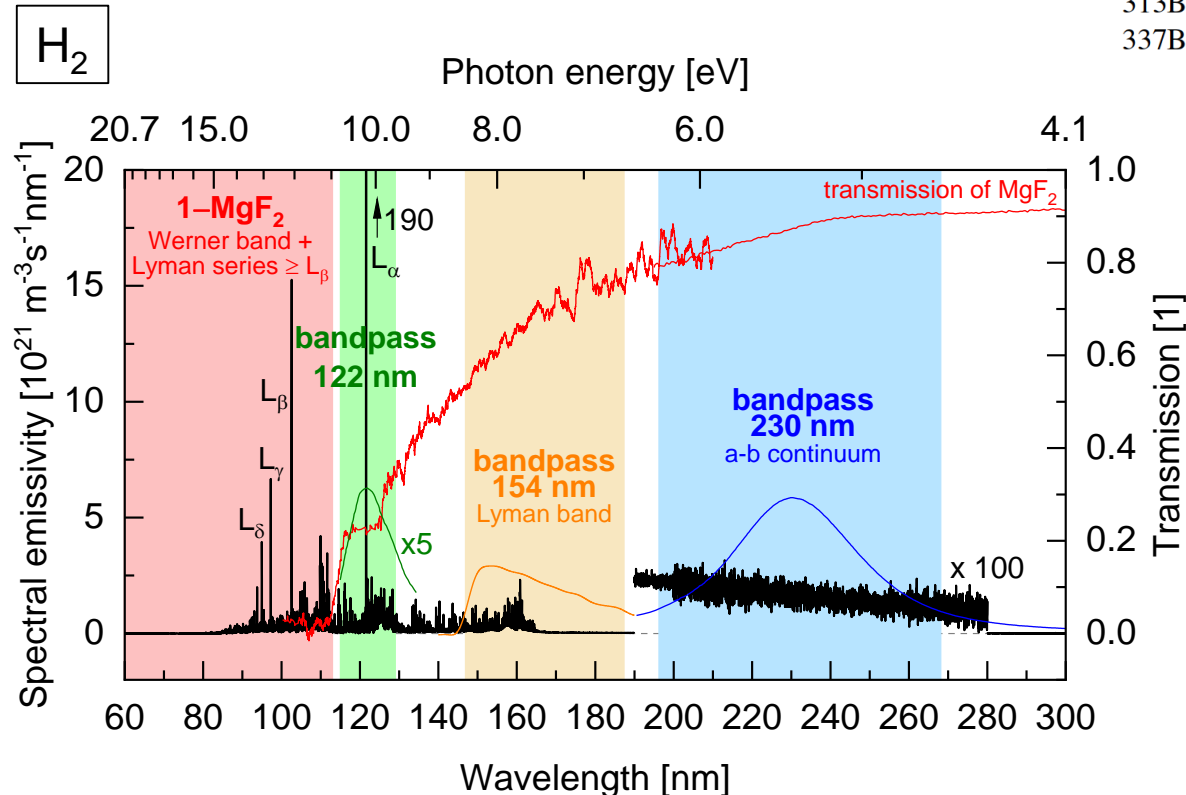


FRIEDL et al., *Meas. Sci. Technol.* **34** (2023) 055501



Detecting VUV emission

VUV diode system – available filter sets



Band pass filters			Long pass filters	
Filter name	Central wavelength	FWHM	Filter name	Cut-on wavelength
122BP	122 nm	14 nm	MgF ₂	113 nm
154BP	154 nm	30 nm	Fused silica (FS)	153 nm
230BP	230 nm	37 nm	BK7	281 nm
313BP	313 nm	10 nm	400LP	397 nm
337BP	337 nm	10 nm	500LP	500 nm

↓

H ₂	Emission system	Filter combination	Filter interval
	H ₂ Werner band (C-X), H lines $\geq L_\beta$	1-MgF ₂	≤ 113 nm ≥ 11.0 eV
	H line L _{α}	122BP	115–129 nm 10.2 eV
	H ₂ Lyman band (B-X)	154BP	147–187 nm 8.4–6.6 eV
	H ₂ Continuum (a-b)	230BP	196–268 nm 6.3–4.6 eV

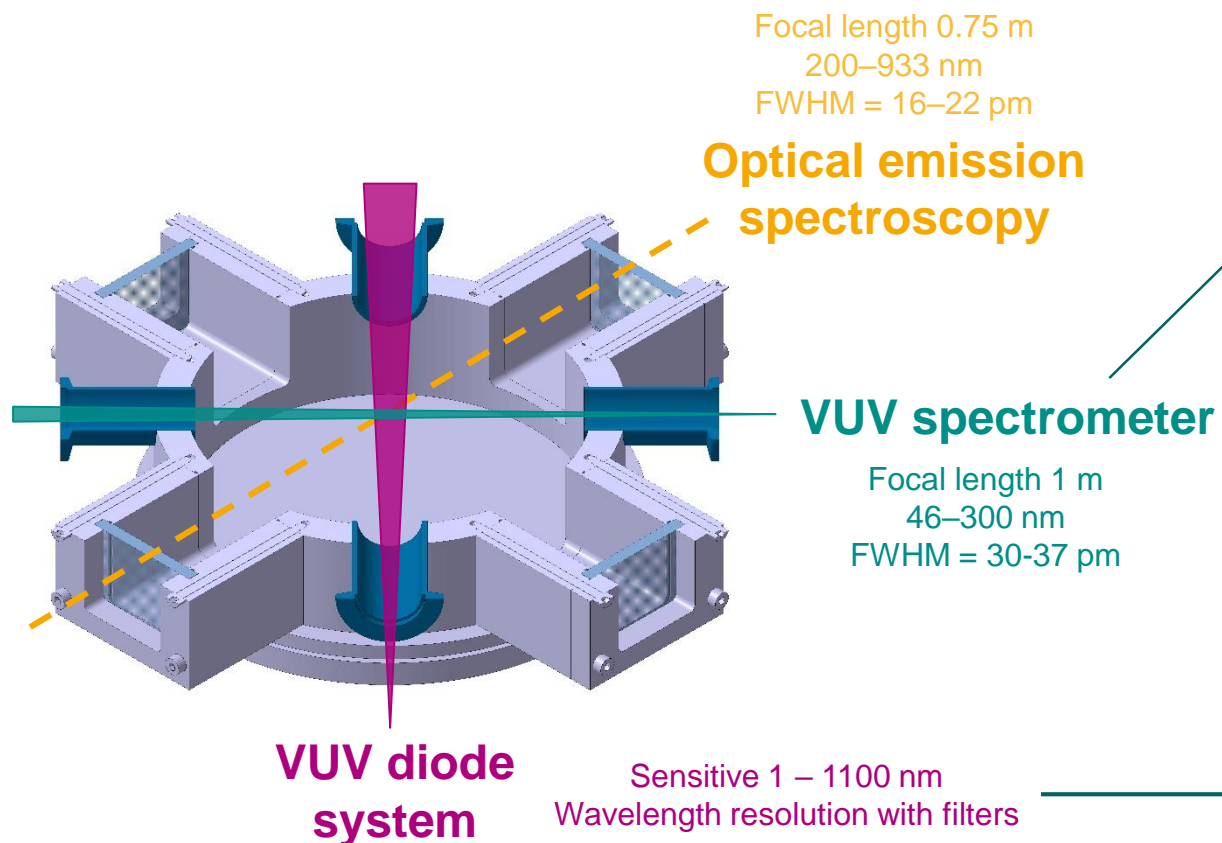
➤ Available for Ar, H₂, N₂, O₂, and mixtures

Detecting VUV emission

VUV diode system – absolute calibration

PlanICE (planar ICP, 2 MHz, 2 kW)

FRÖHLER-BACHUS et al., *JQSRT* **259** (2021) 107427



PMT: 116.5–300 nm
CEM: 46–116.5 nm

Absolute radiometric calibration

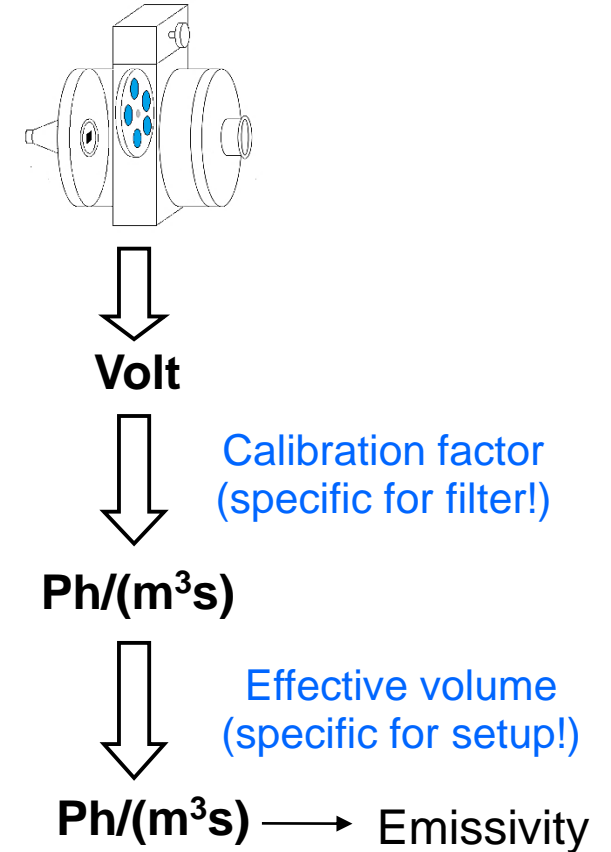
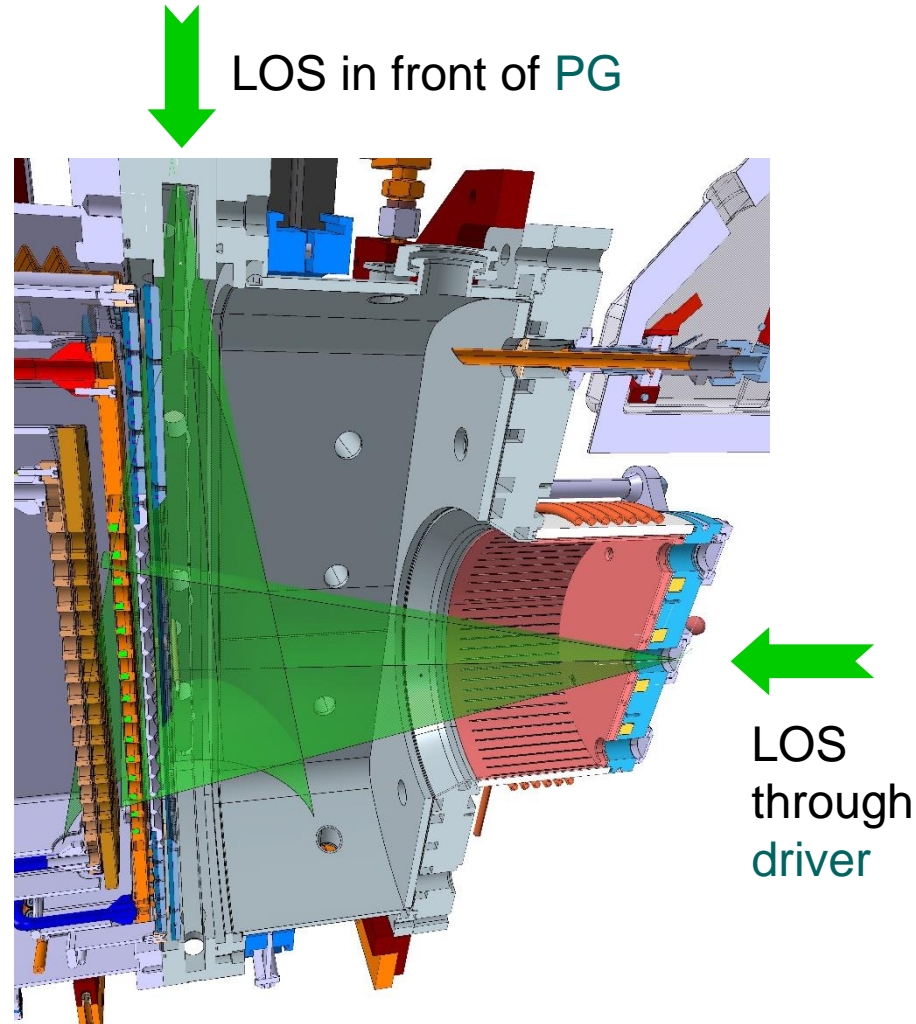
(→ high current hollow cathode,
branching ratios in N₂,
D₂ arc lamps, Ulbricht sphere)

- Gas and filter specific calibration factors for Ar, H₂, N₂, O₂ (etc.)
- Energy-resolved absolute VUV flux measurements up to photon energies of 27 eV
- Accuracy better than 25%,
Dynamic range > 4 ord.magn.

VUV emission in ion sources

Demonstration at negative hydrogen ion source of BUG

- H₂ plasma generated in driver
- H⁻ produced at caesiated plasma grid (PG) with low WF
- Plasma surface interaction at PG drives strong dynamic of ion source performance, where VUV: affects surface WF & induces emission of photoelectrons
- VUV emission never measured before in BUG

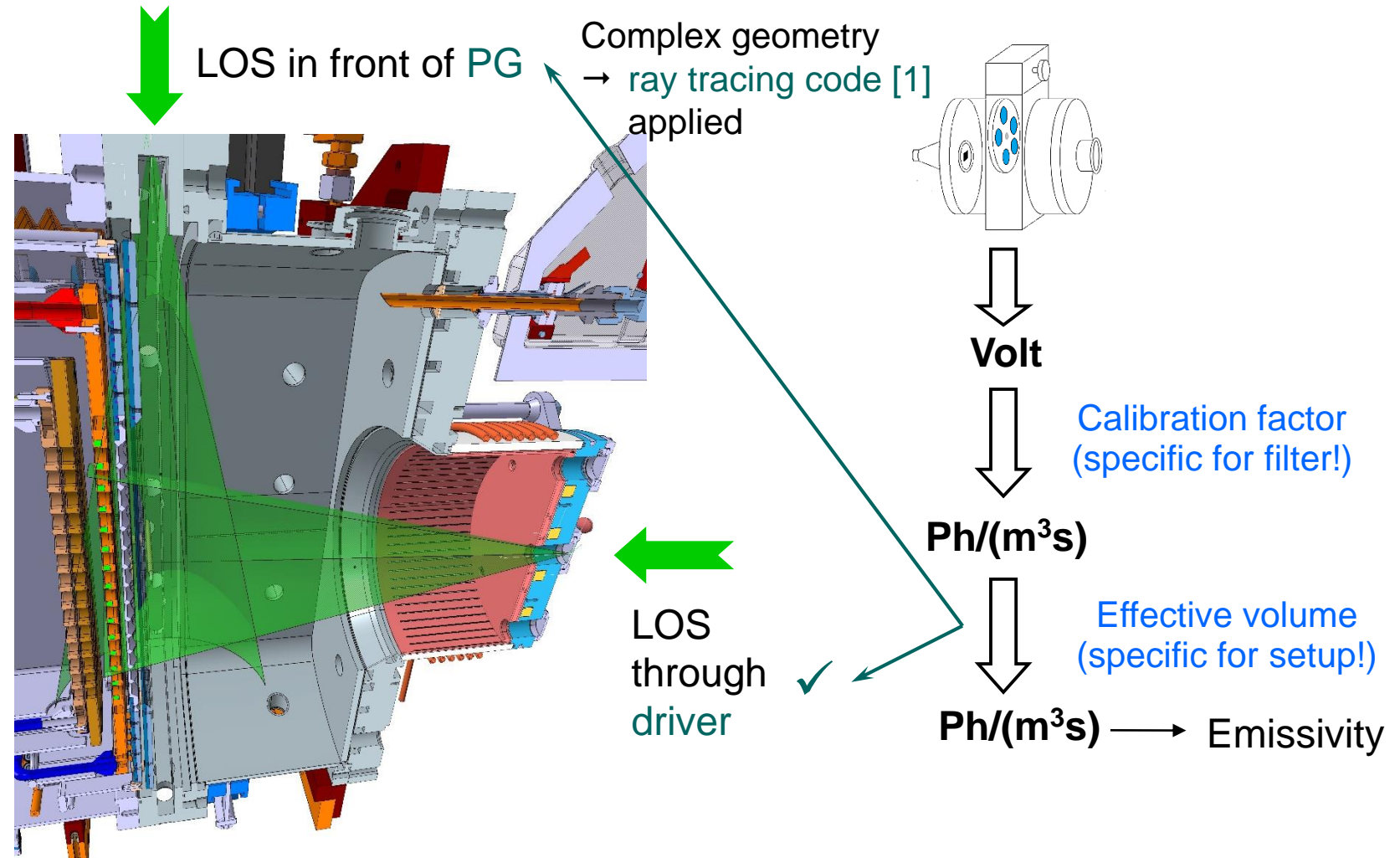


VUV emission in ion sources

Demonstration at negative hydrogen ion source of BUG

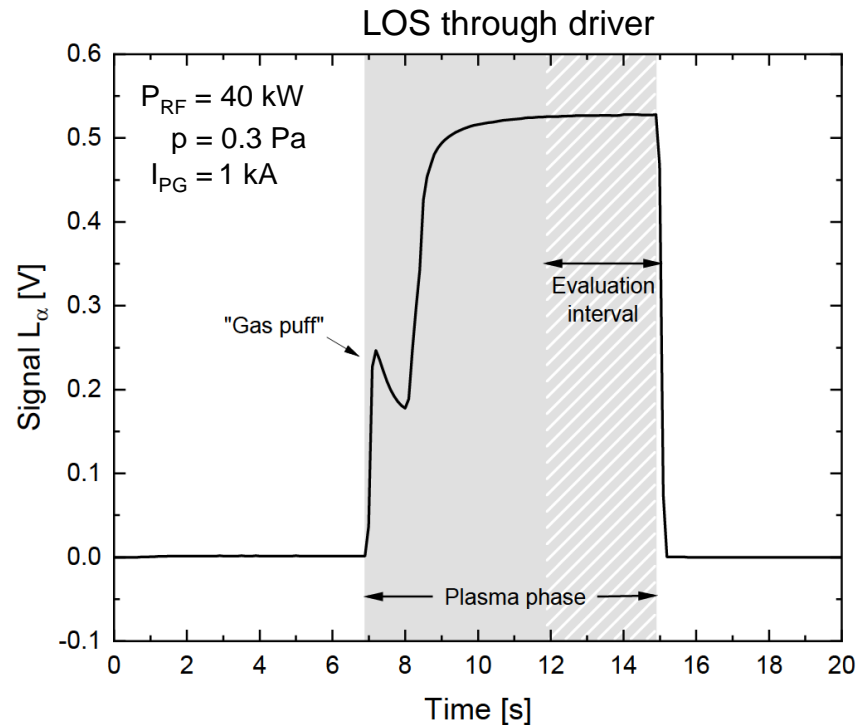
[1] HURLBATT, *J. Phys. D* **53** (2020) 125204

- H_2 plasma generated in driver
- H^- produced at caesiated plasma grid (PG) with low WF
- Plasma surface interaction at PG drives strong dynamic of ion source performance, where VUV: affects surface WF & induces emission of photoelectrons
- VUV emission never measured before in BUG

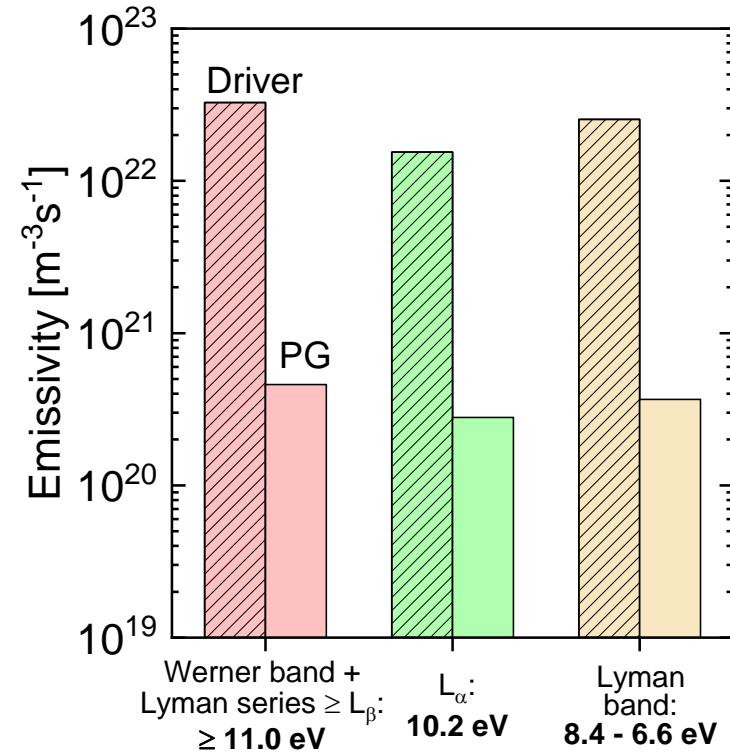


VUV emission in ion sources

Demonstration at BUG (1)



- ✓ Low noise
- ✓ High temporal resolution (here 100 ms)



$P_{RF} = 40 \text{ kW}$
 $p = 0.3 \text{ Pa}$
 $I_{PG} = 1 \text{ kA}$

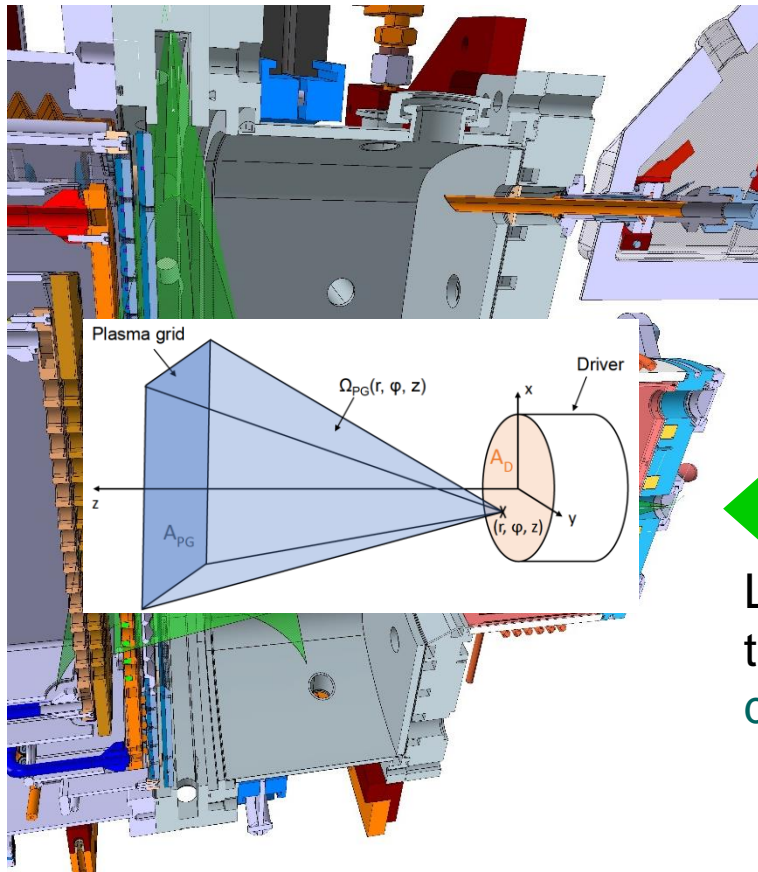
Driver: $T_e = 9.8 \text{ eV}$, $n_e = 5 \times 10^{17} \text{ m}^{-3}$
 PG: $T_e = 2.3 \text{ eV}$, $n_e = 6.6 \times 10^{16} \text{ m}^{-3}$

- ✓ VUV in driver almost 2 ord.magn. more intense
- ✓ Dependence on plasma parameters agrees with CR modeling (Yacora)

VUV emission in ion sources

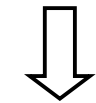
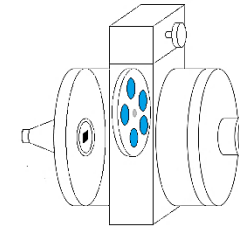
Demonstration at BUG (2)

↓ LOS in front of PG ✓

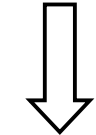


← LOS through driver

Beam widening
 → solid angle calculation required
 → flux reduced by factor ≈ 10

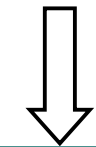


Volt



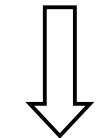
Calibration factor
 (specific for filter!)

Ph/(m³s)



Effective volume
 (specific for setup!)

Ph/(m³s) → Emissivity

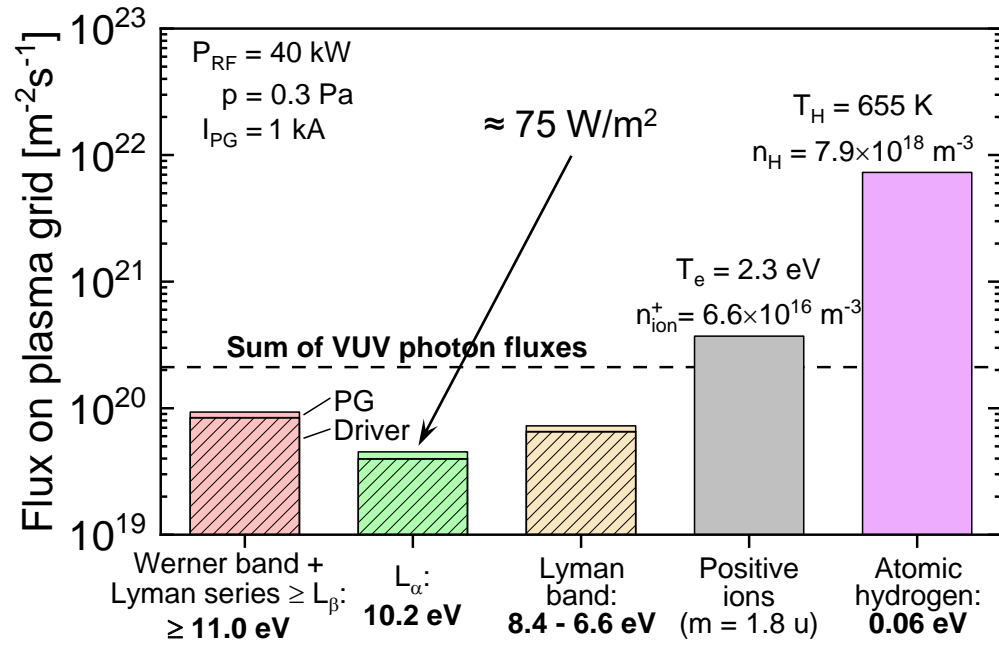


Geometry (Volume/Surface)

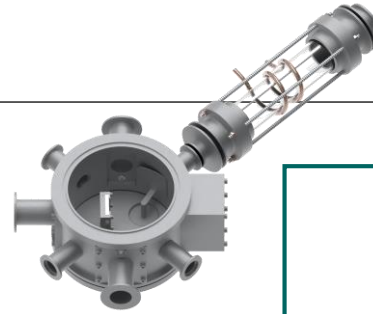
Ph/(m²s) → Photon flux

VUV emission in ion sources

Demonstration at BUG (3)

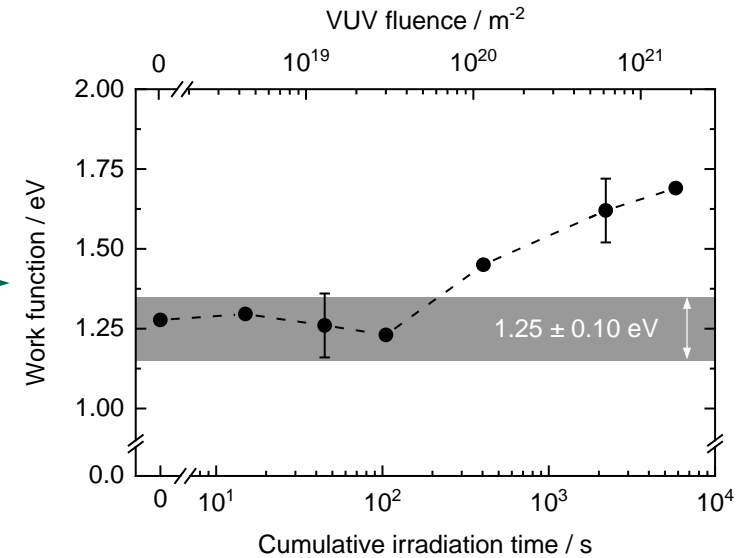
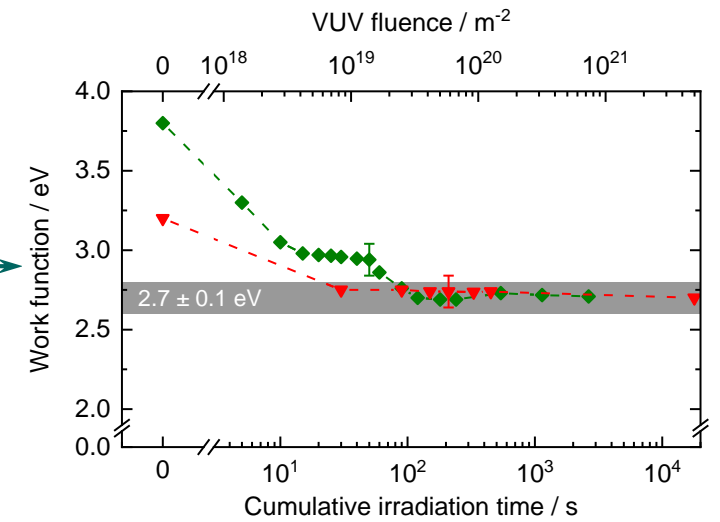


- VUV flux dominated by driver emission
- Atomic flux dominant but low energy
- Total VUV flux comparable to ion flux
- VUV fluxes high enough to impact the WF



HEILER, *PhD*, UniA 2022
 & NIBS, Padova 2022

VUV flux $< 10^{18} \text{ m}^{-2}\text{s}^{-1}$

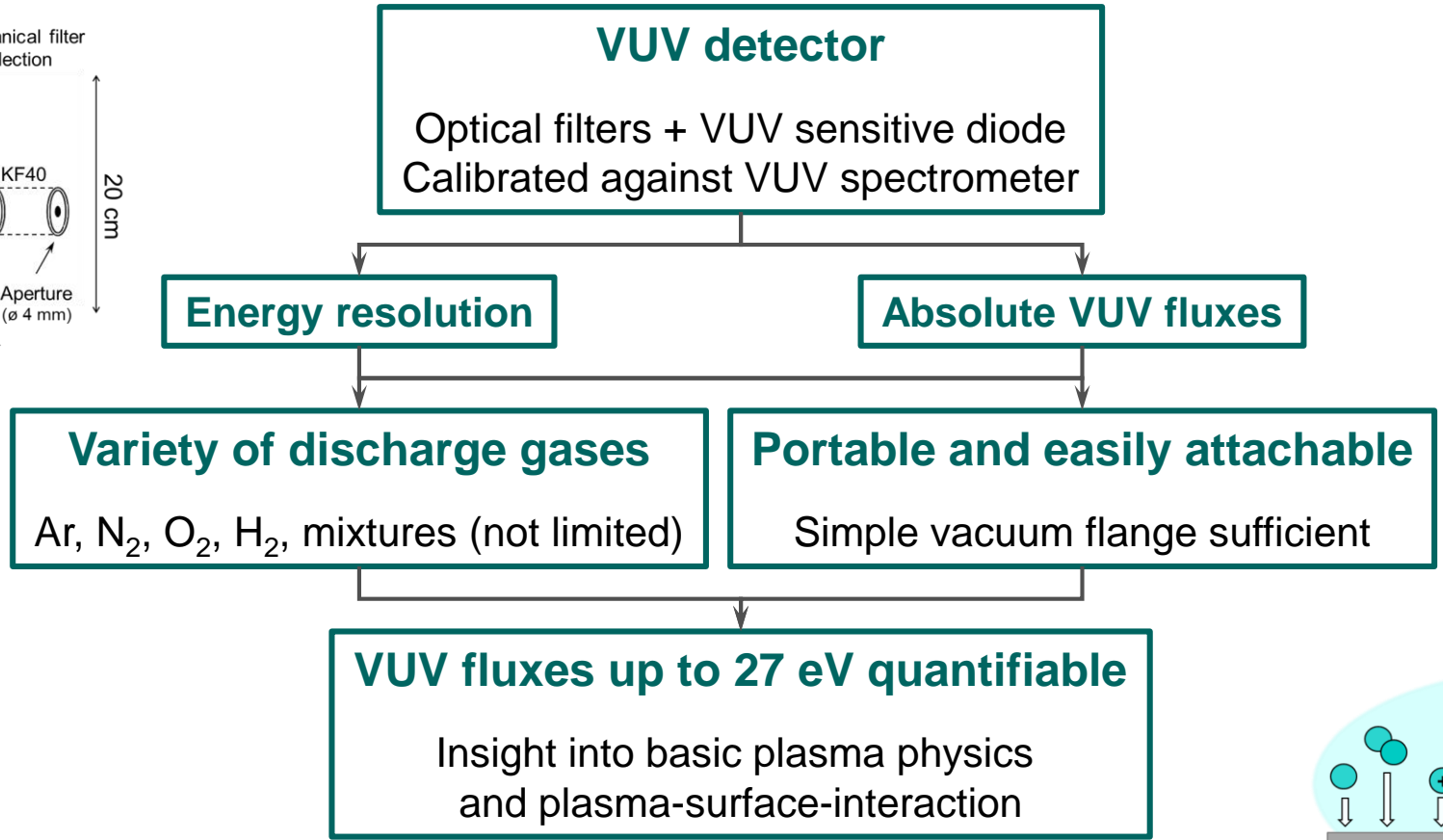
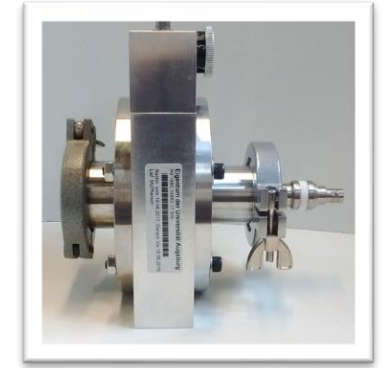
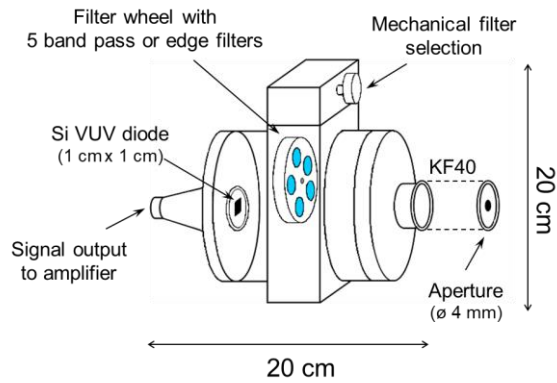


WÜNDERLICH et al., *ICPIG*, Iași 2015

VUV flux $> 10^{22} \text{ m}^{-2}\text{s}^{-1}$

Photoelectrons negligible

Conclusion



at BUG

- Comparable photon to ion fluxes onto PG
- Photon energies >11 eV
- Flux magnitude sufficient to change the WF

