# Reducing our dependence on slow simulators with deep learning

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## ML for classification and beyond

There is a lot more ML can do than classify examples!



map noise to structure









- Simulation dependence in collider HEP
- Generative models: accelerate simulations with deep learning
- Generalized numerical inversion: machine learn

with prior independence

- Weak supervision: learn directly from (unlabeled) data
- CWoLa hunting: model agnostic **anomaly detection**

#### Simulation at the LHC

10000000000 m reegeleeleelee ...... and a contract and a mmmmmm mmm Recence Spanning 10<sup>-20</sup> m up to 1 m Inspired by Sherpa 1.1 can take O(min/event) paper - can you spot the differences?

#### Simulation at the LHC

.....

Part I: accelerate the existing simulation

LEELELELELEL

Recent

egeleeleelee!

Inspired by Sherpa 1.1 paper - can you spot the differences?

## Spanning 10<sup>-20</sup> m up to 1 m can take O(min/event)

State-of-the-art for material interactions is Geant 4.

Includes electromagnetic and hadronic physics with a variety of lists for increasing/decreasing accuracy (at the cost of time)

This accounts for O(1) fraction of all HEP computing resources!



Goal: replace (or augment) simulation steps with a faster, powerful generator based on state-of-the-art machine learning techniques

# This work: attack the most important part: Calorimeter Simulation

We are not trying to generate an entire event (O(1000) particles)) all at once - it would be **very had to validate!** Instead, generate a single particle shower (before electronics) and appeal to combinatorics. We are not trying to generate an entire event (O(1000) particles)) all at once - it would be **very had to validate!** Instead, generate a single particle shower (before electronics) and appeal to combinatorics. We are not tryin event (O(1000) r would be **very h**and generate a single electronics) and a

### N.B. calorimeter energy deposits factorize (sum of the deposits is the deposit of the sum) but digitization (w/ noise) does not!





## Now to the machine learning

A generator is nothing other than a function that maps random numbers to structure.



#### Our structure: calorimeter images

## Calorimeter images



Grayscale images: Pixel intensity = energy deposited



## Calorimeter images

Challenge: multiple layers with non-uniform granularity and a causal relationship?

N.B. images are O(1000) dimensional





## One possibility: GANs

Generative Adversarial Networks (GAN): *A two-network game where one maps noise to images and one classifies images as fake or real.* 



## Introducing CaloGAN



## Introducing CaloGAN

Mode collapse: learns to generate one part of the distribution well, but leaves out other parts.

#### help avoid 'mode collapse'

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**Discriminator network** 

## Locally connected layers

Due to the structure of the problem, we do not have translation invariance.

# of filters

stride

Classification studies found fully connected networks outperformed CNNs

However, convolutional-like architectures are still useful to e.g. reduce parameters

## Locally connected layers



## Results: average images

**Geant4** 





## Energy per layer



## Warning: challenge with GANs

Unlike for classifiers, it is not easy to figure out which GAN is a good GAN - trying to learn a O(1000) generative model and not a single likelihood ratio!

...this is a place where science applications can make a big impact on ML.



## "Overtraining"



A key challenge in training GANs is the diversity of generated images. This does not seem to be a (big) problem for CaloGAN.



## Extrapolating



GANs are not designed to extrapolate, but in some cases, they can smoothly go on!

works here until there is no new physical principles which turn on at some energy

## Conditioning

Fix noise, scan latent variable corresponding to energy



Fix noise, scan latent variable corresponding to x-position



## Timing



<b>Generation Method</b>	Hardware	Batch Size	milliseconds/shower
GEANT4	CPU	N/A	1772 -
CALOGAN	CPU Intel Xeon E5-2670	1	13.1
		10	5.11
		128	2.19
		1024	2.03
		1	14.5
		4	3.68
	GPU	128	0.021
	NVIDIA K80	512	0.014
		1024	0.012 🔶

(clearly these numbers will change as both technologies improve - this is simply meant to be qualitative and motivating!)

## Collaboration workflow

Integrating these techniques into a full detector simulation is another layer of complication, but is possible and hopefully worth the effort!



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GAN studies by Aishik Ghosh and others



One source of MC dependence is the same as classification:

→ mis-modeling dependencies between features

However, there is a new source: dependence on the feature priors.



This is really new for regression for classification, if p(xlsignal) and p(xlbackground) are mis-modeled, you get the optimal answer as long as their ratio is correct.

An example that you can have in mind is energy calibration.



We want to predict the true energy given the measured energy

(and possibly other features - more on that soon)

...however what I'm about to say applied more generally (though the impact is biggest when the resolution is poorest)

Suppose you have some features x and you want to predict y.

detector energy true energy

One way to do this is to find an f that minimizes the mean squared error:

$$f = \operatorname{argmin}_g \sum_i (g(x_i) - y_i)^2$$

Proof in the lectures from yesterday! Then, f(x) = E[y|x].

Why is this a problem?



## $f(x) = E[y|x] = \int dy \, y \, p(y|x)$

## $E[f(x)|y] = \int dx \, dy' \, y' \, p_{\text{train}}(y'|x) \, p_{\text{test}}(x|y)$

this need not be y even if  $p_{train} = p_{test}(!)$ 

ATLAS and CMS use a trick to be prior-independent:

**Numerical inversion** *instead of predicting y from x, predict x from y and then invert the function* 

... put another way:

learn f:y  $\rightarrow$  x and then for a given x, predict f<sup>-1</sup>(x)

by construction, f is independent of p(y) and thus f<sup>-1</sup> also does not depend on p(y), as desired.

This procedure is independent of the prior p(y) but may not close exactly, i.e.  $E[f^{-1}(x)|y]$  may not be y.

...under mild assumptions, it does close for the mean absolute error, but usually has some non-closure for the squared error.

Also, the calibration procedure can distort the underlying distribution, i.e. if you start with a Gaussian, you almost never end up with exactly a Gaussian.

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The detector response of jets depends on many properties of the jet. Ideally, the calibration can include this!



The current ATLAS approach to including more features is to repeat NI sequentially:

$$p_{\mathrm{T}}^{\mathrm{reco}} \mapsto \hat{p}_{\mathrm{T}}^{\mathrm{reco}} = f_{\theta_n}^{-1} \left( \cdots f_{\theta_2}^{-1} \left( f_{\theta_1}^{-1} \left( p_{\mathrm{T}}^{\mathrm{reco}} \right) \right) \cdots \right)$$



This works well when the jet response is independent of  $\theta_i$  given  $\theta_j$ .



For reasons discussed earlier, we can't include correlations by learning y given x and all the  $\theta$ 's.

However, it would still be great to use machine learning to automatically and efficiently make use of correlated information.

We cannot use numerical inversion out-of-the-box because we now have a many-to-one function.


Since we are not (necessarily) interested in calibrating the  $\theta$ 's, we can generalize NI as follows:

(1) Learn a function f to predict x given y and all the  $\theta$ 's. (2) For every combination of  $\theta$ , invert f. (3) Calibrate via  $x \rightarrow f_{\theta}^{-1}(x)$ 

Step (2) is intractable, so replace it with another learning step: predict y given  $f(y,\theta)$  and  $\theta$ .

# GNI in action



#### Consider two features:



# GNI in action



#### $\hat{R}$ is the calibrated E[x|y] / y



Only the simultaneous approach removes the full residual dependence!





Adding more features (with more interdependencies) will lead to more dramatic improvements.

We can also extend this approach to calibrate other observables and even simultaneously calibrate some of the  $\theta$ 's (even more generalized NI!)



## **Background and Motivation**

Usual paradigm: train in simulation, validate on data, test on data.

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If data and simulation differ, this is sub-optimal!





 $h(n_{trk}, Track Width) \rightarrow [0,1]$ 

## Background and Motivation

Usual paradigm: train in simulation, validate on data, test on data.



 $h_{MC}(n_{trk}, Track Width) \rightarrow [0,1]$ 

## Background and Motivation

Usual paradigm: train in simulation, validate on data, test on data.





(N.B. f & g from simulation and selection can't bias Q and G - more on that later)

Determine the performance of the WP in data.

How did we get this?

dijets =  $f_q \times \mathbf{Q} + (1-f_q) \times \mathbf{G}$ 

Z+jets =  $g_q \times Q + (1-g_q) \times G$ 

2 equations, 2 unknowns (Q, G)

 Image: WP in data:

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Esignal, data, Eback, data

Can correct the MC to have the same performance as data.



Once we have scale factors (& their uncertainty), we can ensure that our analysis will be accurate.

...so what is the problem?

remember my claim from earlier:

If data and simulation differ, this is sub-optimal!

This is an accuracy versus precision problem. It is "easy" to achieve accuracy through calibration, but the results may not be the best one possible. In this 2D feature space, we can actually derive h<sub>data.</sub>



h<sub>MC</sub>(n<sub>trk</sub>, Track Width) ≠ h<sub>data</sub>(n<sub>trk</sub>, Track Width)



To stress this point, suppose that  $h_{MC}$  is the random classifier:

 $h_{MC} = 0$  if you pick a random number x in [0,1] and x <  $\varepsilon$ 1 otherwise

We can calibrate this classifier in data, but clearly, it is sub-optimal !!



#### One more slide about why it matters



J. Barnard, E. Dawe, M. Dolan, N. Rajcic, Phys. Rev. D 95 (2017) 014018

Especially important for **deep learning** using subtle features → hard to model!

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W boson radiation pattern - same physics, different simulators!



One of the biggest challenges with any MC-based method is that it can't use information that the MC doesn't know about.

One solution is to train directly on data !

In general, this is not possible since data are unlabeled. However, in a wide range of cases, it is possible to work with less.

There is an interesting connection between what I'm calling "weak supervision" and the topic of "label noise".

The setup: suppose you have (at least) two mixed samples, each composed of two classes (say q and g).

Requirement:

The two classes are well-defined i.e. q in sample 1 is statistically identical to q in sample 2).

Remember this plot?

dijets =  $f_q \times Q + (1-f_q) \times G$ Z+jets =  $g_q \times Q + (1-g_q) \times G$ 

two equations, two unknowns (Q, G)

We often know f, g (from ME + PDF) much better than full radiation pattern inside jets.

ً18∄ ≣\_ = d/(d+g 0.9 16 ΔΤΙ Δ.S Discriminant for Data-Driven Tagger 0.8 L dt = 4.7 fb<sup>-1</sup>,  $\sqrt{s}$  = 7 TeV 14 anti-k, R=0.4,  $\ln l < 0.8$ 0.7 12 160 GeV<p <210 GeV 0.6 10 0.5 quark vs gluon 8 0.4 jets in data 6 0.3 4 0.2 2 0.1 0 **0** 0.05 0.150.1 0 **Track Width** 

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This doesn't work well when you have more than 2 observables because the templates become sparse.

## Method 1: Learn from Proportions



L. Dery, BPN, F. Rubbo, A. Schwartzman, JHEP 05 (2017) 145

proportions

## N.B. Don't need 100% fraction accuracy

Even though the proportions are required as input, if they are slightly wrong, you can end up with the correct classifier.



T. Cohen, M. Freytsis, B. Ostdiek, JHEP 02 (2018) 034

#### Works in low-dimensions



## A note about training statistics



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how different are the proportions for the two mixed samples

# Method 2: Learning without Proportions



E. Metodiev, BPN, J. Thaler, JHEP 10 (2017) 51

#### Works in low-dimensions





0.6

0.8

0.5<sup>L</sup>

0.2



As with LLP, need sufficient effective statistics

0.8

 $f_{1} (= 1 - f_{2})$ 

0.6

0.4

Can't learn when the two proportions are the same.

#### Methods Overview

Property	$\mathbf{LLP}$	CWoLa
Compatible with any trainable model	$\checkmark$	$\checkmark$
No training modifications needed	X	$\checkmark$
Training does not need fractions	X	$\checkmark$
Smooth limit to full supervision	X	$\checkmark$
Works for $> 2$ mixed samples	$\checkmark$	?

## Next step: what about high dim.?

There are many O(1)-dimensional ML problems for jets, but since the full radiation pattern is higher dimensional, need to go to bigger!

We'll use jet images as a testing ground, still focusing on quarks versus gluons.





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Translated Pseudorapidity  $\eta$ 



The CWoLa approach works out-of-the box - can use welltested CNN architecture with usual cross-entropy loss.

On the other hand, LLP requires significant work on the technical implementation / optimization.

$$\ell_{\text{WMSE}} = \sum_{a} \left( f_a - \frac{1}{N} \sum_{i=1}^{N} h(\mathbf{x}_i) \right)^2 \qquad \ell_{\text{WCE}} = \sum_{a} \text{CE} \left( f_a, \frac{1}{N} \sum_{i=1}^{N} h(\mathbf{x}_i) \right)$$

#### Works in many-dimensions!



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P. Komiske, E. Metodiev, **BPN**, M. Schwartz, Phys. Rev. D 98, 011502(R), arXiv:1801:10158

#### A note about training statistics



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As usual, it is likely that the best approach will use all of the available information, including some input from simulation.

...this could be as simple as pre-training in MC and then running weak supervision





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I will leave you with one last, but very exciting topic.

One of the most important goals of HEP is to search for new particles. However, we have not found anything (significantly) unexpected in a while ... we need simulationindependent ways of searching for new particles !

anomalies, i.e. something unexpected

N.B. The approach discussed here is not the only one - see also M. Farina, Y. Nakai, D. Shih, 1808.08992 & T. Heimel, G. Kasieczka, T. Phlen, J. Thompson, 1808.08979 for an alternative approach based on auto-encoders.

### Remember CWoLa ...



*Classification Without Labels* 

Solution: Train directly on data using mixed samples



E. Metodiev, BPN, J. Thaler, JHEP 10 (2017) 51

#### Weak/unsupervised learning for anomalies

Can we take this idea one step further to look for something unexpected?

## = CWoLa Hunting\*





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Minimal assumption: if there is a signal, it is localized in one known dimension.

\*Image from this article. This Koala is actually being freed - I do not condone violence against these animals!
## Weak/unsupervised learning for anomalies

Mixed sample 1: signal region

Mixed sample 2: sideband region

Train a classifier to distinguish the two mixed samples.



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If there is a signal, there will be something to learn and the signal will be enhanced. If no signal, nothing to learn.

## Weak/unsupervised learning for anomalies

Need to be careful about testing/training on the same data.

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## Weak/unsupervised learning for anomalies



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Using a classifier trained to distinguish a signal region from a sideband, make progressively harsher cuts on the NN output

### CWoLa hunting vs. Full Supervision



If you know what you are looking for, you should look for it. If you don't know, then CWoLa hunting may be able to catch it!



(1) Need an observable X (e.g.  $m_{JJ}$ ) where the signal is localized and the background is not.

(2) Identify features Y (e.g. jet substructure) that are ~independent of X, but can be useful for identifying a broad range of new particles.

actually, we don't need independence, we just need them to not allow us to sculpt bumps.





#### •Generative models: accelerate

simulations with deep learning

Generalized numerical inversion:

machine learn with prior independence

• Weak supervision: learn directly from

(unlabeled) data

• CWoLa hunting: model agnostic

### anomaly detection



# Conclusions and Outlook

Accelerating simulation/reducing simulation dependence can improve performance and may even help us to understand something new and fundamental about nature!





to structure

