Opportunities for Fundamental Symmetries Experiments with Neutrons and Nuclei

A. R. Young

North Carolina State University Triangle Universities Nuclear Laboratory



Outline

- Beta Decay
 - ¹⁹Ne and nuclear mirror decays
 - The β -asymmetry with UCN
 - Other experiments?
- Spin precession with neutrons
 - n-n' precession experiments
 - SR spin-dependent force

Two Drivers for High Precision Measurements with Mirror Decays

- i) Progress on limiting uncertainties of the electroweak corrections
- ii) Nuclear mirror decay analysis problem identified and solved



Precision of hadronic component of inner radiative corrections ~ factor of two improvement!



Seng, Gorchtein, Patel, Ramsey-Musolf, PRL **121**, 241804 (2018)

IMPACT: |V_{us}| now nominally provides limiting uncertainty for unitarity

4σ shift in EWRC correction – ~0.1% in unitarity sum!







37K



Lepton Flavor Universality

Crivellin and Hoferichter, 2002.07184



Hardy 2015 (modified)

EWRC

Nuclear Structure



Improve lattice constraints with improved K/ π ratio for $K\ell 3$ with improved measurement of $\pi^+ \rightarrow \pi^0 e^+ \nu$

Opportunities for Progress

V_{us}?

Nuclear Structure?

- Neutron lifetime and angular corr (no nuclear structure)
- Improved measurements of mirror nuclei
 - target \leq 0.1% precision (probe effects at scale of EWRC shift)
 - A=10 to A =20 multiple theory approaches, few superallowed cases



Marciano (INT 2019)

$$\begin{split} \mathsf{R}_{\mathsf{A}} = & \Gamma(\mathsf{K}_{-} > \mu v(\gamma)) / \Gamma(\pi_{-} > \mu v(\gamma)) \\ & vs \\ \mathsf{R}_{\mathsf{V}} = & \Gamma(\mathsf{K}_{\mathsf{L}}_{-} > \pi e v(\gamma)) / \Gamma(\pi_{-}^{+} > \pi_{0}^{0} e^{+} v(\gamma)) \end{split}$$

Improve lattice constraints with improved K/ π ratio for $K\ell$ 3 with improved measurement of $\pi^+ \rightarrow \pi^0 e^+ v$



V_{us}?

Nuclear Structure?

BSM?

- e/μ K, π decay ratios
- High precision neutron and mirror decay measurements:
 - can provide excellent
 broad band sensitivity
 to BSM
 - in some cases not limited by EWRC corr. (will provide example!)

- Neutron lifetime and angular corr (no nuclear structure)
- Improved measurements of mirror nuclei
 - target \leq 0.1% precision (probe effects at scale of EWRC shift)
 - A=10 to A =20 multiple theory approaches, few superallowed cases



Marciano (INT 2019)

R_A=Γ(K->μν(γ))/Γ(π->μν(γ)) vs R_V=Γ(K_L->πev(γ))/Γ(π⁺->π⁰e⁺ν(γ))

Improve lattice constraints with improved K/ π ratio for $K\ell$ 3 with improved measurement of $\pi^+ \rightarrow \pi^0 e^+ v$

Status of Mirror Decays, Issues Resolved!



Nuclear mirrors in good shape!

L. M. Hayen: GT recoil order effects and radiative corrections 1906.09870
$$\begin{split} |V_{ud}|_{0^+ \to 0^+} &= 0.97370(21) \\ |V_{ud}|_n &= 0.97399(78) \\ |V_{ud}|_{n+mir} &= 0.97396(63) \end{split}$$

Current status sets stage for drive to higher precision!

TRIUMF is wonderfully positioned to have a major impact on the status of the charged current

How to proceed?

Optimizing impact:

- Identify opportunities to push to 0.01% (superallowed) precision
- "Modest" technical advances, building on local expertise

Areas where competition "thin" –only 2 actually "on the books" at 0.2% sensitivity or below

Angular correlation measurements with polarized nuclei

Large enhancements of sensitivity to coupling to ratio of F to GT decay amplitudes ρ for n (~4) and ¹⁹Ne (~13) due to accidental cancellations in asymmetry

To reach current precision of superallowed with asymmetry measurement :

- neutron $\sim 0.1\%$
- ¹⁹Ne ~0.3%



LH, A. Young, 2009.11364



 $\Gamma = 1 + A(E) \langle J \rangle \beta \cos \theta$

How to proceed?

Optimizing impact:

- Identify opportunities to push to 0.01% (superallowed) precision
- "Modest" technical advances, building on local expertise
- Areas where competition "thin"

Angular correlation measurements with polarized nuclei

Large enhancements of sensitivity to coupling to GT decay amplitudes for n (~4) and ^{19}Ne (~13) due to accidental cancellations in asymmetry

To reach current precision of superallowed wi th asymmetry measurement :

- neutron ~0.1%
- ¹⁹Ne ~0.3%

Already Achieved by TRINAT group!



Decay Scheme and Advantages of ¹⁹Ne



- Decay scheme ~100% to G.S
- All decay properties adequately known for 0.02% measurement except ρ
- Optically trappable

Table 2: Parameters used to calculate $\mathcal{F}t_0$ and A_0 .

Constant	Value	Units	Reference	
$K/(\hbar c)^6$ 8	$8120.278(4) \times 10^{-10}$	GeV ⁻⁴ s	(44)	
$G_F/(\hbar c)^3$	$1.16637(1) \times 10^{-5}$	${\rm GeV^{-2}}$	(44)	
Δ_R^V	$2.467(22) imes 10^{-2}$		(11)	
Q_{EC}	3.23949(16)	MeV	(45)	
f_V	98.648(31)		(46)	
f_A/f_V	1.0012(2)		(47)	
$\delta^V_C - \delta^V_{NS}$	$0.52(4) imes 10^{-2}$		(25)	
BR	0.999878(7)		(22)	
P_{EC}	0.00101(1)		(25)	
$t_{1/2}$	17.2578(34)	S	(48, 49)	
F_{σ}^{V}	-148.5605(26)		(24, 50)	
W_0	2.72850(16)	MeV	(45)	
m_e	0.510998910(13)	MeV	(51)	
M	19.00188090(17)	amu	(45)	



Most Precise Experiment to Date...





(featured Stern-Gerlach polarizer)

In 1993-95, Young participated in the thesis experiment of G. Jones at Princeton, a measurement of the beta asymmetry of ¹⁹Ne

Successfully developed a model of Si(Li) detector response and detailed simulation of decay and calibration data to produce consistent analysis (thesis D. Combs)

¹⁹Ne Beta Decay Results

Result: no evidence for 2^{nd} class currents, a roughly 0.15% measurement of ρ



Systematic	Correction (10^{-4})	Uncertainty (10^{-4})
Monte Carlo Corrections:		
Above threshold in both detectors:		
Backscatter correction	+14.5	± 3.6
Energy loss correction	-2.0	± 0.5
Above threshold in a single detector:		
Backscatter correction	+3.1	± 0.8
Energy loss correction	-0.9	± 0.2
Below threshold in both detectors:	-0.5	± 0.1
Polarization	_	+5.7 -0.0
Spin relaxation	+5.3	± 5.3
Energy non-linearity	_	± 0.5
Dead time	-0.5	± 0.4
Pileup	+0.6	± 0.4
Background subtraction	+0.2	± 0.2
Statistical	_	<u>+</u> 2.6
Total	+19.8	+6.5, -8.7

 $A_0 = -0.03872(65/-87)_{sys}(26)_{stat}$

 $\rho = 1.6014(+21/-28)_{sys}(8)_{stat}$

Impact

Mirror decays now fully consistent with neutron and superallowed (n and mirrors now make an impressive crosscheck of superallowed



Ratio of ¹⁹Ne to n Ft₀ has impressive sensitivity to tensor Fierz terms:

$$R = 1 + 0.051\epsilon_S - 6.1\epsilon_T$$

 $\Lambda_T > 5.5 TeV$

Over 20 TeV with 0.02% precision!

Note: Fierz term changes sign from beta to positron decays! Also handy for other measurements

Ongoing Research...

Laser-trapped species include Alkali metals (³⁷K) and meta-stable & noble gas atoms (¹⁹Ne)





TRINAT measurement improves all major sources of uncertainty in Princeton measurement by factor of 10, improvements on the way...

Options for improvement may exist for neutrons too!

Motivation for Measurement of Beta Asymmetry with UCN

Polarization: "Potential barrier" polarization demonstrated effective alternative to supermirror/³He cell technology with P \geq 99.5% and ultimate uncertainties at or below 0.1% level



(note: neutron magnetic moment is negative)

Neutron generated backgrounds: small number of neutrons and low capture probability (long residency time) lead to order of magnitude improvement relative to (then) current cold neutron beams experiments



The Original Concept for UCNA



With **116 dps in an open trap:** uncertainty in $\delta A_o/A_o < 0.2\%$ in ~1 year



Obtained between 1 and 2 UCN/cm³



Results



$$A_0 = -0.12054(44)_{\text{stat}}(68)_{\text{syst}}$$
$$\lambda \equiv \frac{g_A}{g_V} = -1.2783(22)$$

$$2011-2012$$
$$A_0 = -0.12026(54)_{\text{stat}}(67)_{\text{syst}}$$

2012-2013

$$A_0 = -0.12111(74)_{\text{stat}}(69)_{\text{syst}}$$

Critical Issue: scattering corrections

$$A_0 = -0.12054(44)_{\text{stat}}(68)_{\text{syst}} \qquad \textbf{0.67\%}$$

$$\lambda \equiv \frac{g_A}{g_V} = -1.2783(22)$$

	% Corr.		% Unc.	•
	2011-2012	2012-2013	U. m.	2010 Unc.
$\Delta_{\cos\theta}$	-1.53	-1.51	0.33	• 0.30
$\Delta_{\text{backscattering}}$	1.08	0.88	0.30 -	- 0.34
Energy Recon.			0.20 -	0.31
Depolarization	0.45	0.34	0.17 -	
Gain			0.16	
Field Nonunif.			0.11	Most dramatic
Muon Veto			0.03	improvement,
UCN Background	0.01	0.01	0.02	from shutter!
MWPC Efficiency	0.13	0.11	0.01	
Statistics			0.36	
Theory	Corrections [9	, 10, 24-27]		
Recoil Order	-1.68	-1.67	0.03	
Radiative	-0.12	-0.12	0.05	



(originally not in design, but required to obtain sufficient decay rate)

If it is possible to load a 10 liter volume (5cm x 200cm) to ~100 UCN/cm³ both issues can be addressed:



Fill through split coil gaps at bottom

Extending the Reach of neutron EDM Experiments with Precession Searches

- EDM apparatus are use successfully to place limits Lorentz invariance violation scenarios, axions and other dark matter and short-ranged interactions
- EDM experiments must be optimized for long storage time and the application of strong electric fields, often limiting their flexibility
- As more EDM experiments come on line and densities increase, planning capability for different cell configurations, inner surface topologies and field orientations may have a significant impact on the physics delivered by the experiment
- Two cases as examples: mirror neutron searches and ALPs

TRIUMF is wonderfully positioned to have a major impact on future precession limits

Neutron-Mirror Neutron Searches

- Motivated by fact that spin-conserving mirror world interaction breaks into two, coupled two state problems when fields are along the quantization axis in respective universes
- Expect generically that eigenstates are effected quadratically by perturbations, precession measurements generically possible! Berezhiani, Eur. Phys. J. C 64, 421-431 (2009)

• Started considering possible experiments with small working group (B. Franke, C. Swank, G. Pignol, S. Roccia)

still effectively sharing my notes...

Conducted to monitor magnetic field dependent oscillations with interaction <u>10¹</u>

$$\mathcal{H} = \begin{bmatrix} -\mu_n \mathbf{B} \cdot \boldsymbol{\sigma} & 1/\tau_{nn'} \\ 1/\tau_{nn'} & -\mu_n \mathbf{B'} \cdot \boldsymbol{\sigma} \end{bmatrix},$$

Results in field dependent oscillations

$$P_{BB'}^{nn'}(t) = \frac{\sin^2[(\omega - \omega')t]}{2\tau_{nn'}^2(\omega - \omega')^2} + \frac{\sin^2[(\omega + \omega')t]}{2\tau_{nn'}^2(\omega + \omega')^2}$$
(2)
+ $\left(\frac{\sin^2[(\omega - \omega')t]}{2\tau_{nn'}^2(\omega - \omega')^2} - \frac{\sin^2[(\omega + \omega')t]}{2\tau_{nn'}^2(\omega + \omega')^2}\right)\cos\beta$



Strongest limits come from storage expts (Serebrov)



A spin-dependent Hamiltonian is presented in Berezhiani et al, Eur. Phys. J. C 72:1974 (2012) for a spin-independent n-n' coupling:

$$H = \begin{bmatrix} \mu B \sigma & \epsilon \\ \epsilon & \mu' B' \sigma \end{bmatrix}$$
(1)

which can be expanded to

$$H = \begin{bmatrix} \frac{\mu}{2}B & 0 & \epsilon & 0\\ 0 & -\frac{\mu}{2}B & 0 & \epsilon\\ \epsilon & 0 & \frac{\mu'}{2}B' & 0\\ 0 & \epsilon & 0 & -\frac{\mu'}{2}B' \end{bmatrix}$$
 For fields along the z and z' axis (2)

for the wavefunction



Mirror neutrons have same mass, opposite magnetic moment

Splitting depends on magnitude of earth and mirror magnetic fields!

- Fields in both neutron and mirror neutron frames have very small spatial variations (true for normal fields for precession experiments)
- Mirror neutrons are not confined by cell walls
- Magnetic field arranged along the z axis in our universe (the experiment)
- Each precession measurement takes ~100-300 s (T_2 or storage time limit)
- Negligible n' amplitude before first spin flip (flip in guide with very short collision time, short times between collisions compared to precession measurement





The cell sweeps out a trajectory, where B_{\parallel}' is constant, and B_{\perp}' rotates at freqency Ω

Average field: (B_{\parallel}) cos θ along z (B_{\parallel}) sin θ along x

Fluctuating field: (B₁')cosθ along z Non-zero mirror field produces a diurnal variation in the coupling to n' and the precession frequency independent secular equation, $H\Psi = E\Psi$, can be written:

$$\begin{bmatrix} f-E & 0 & \epsilon & 0 \\ 0 & f-E & 0 & \epsilon \\ \epsilon & 0 & d-E & 0 \\ 0 & \epsilon & 0 & -d-E \end{bmatrix} \begin{bmatrix} \psi_{n+} \\ \psi_{n-} \\ \psi_{n'+} \\ \psi_{n'-} \end{bmatrix} = 0$$
(4)

The eigenvalue equation for this system is (with $f=\frac{\mu}{2}B$ and $d=-\frac{\mu'}{2}B'$:

$$(E^2 - f^2)(E^2 - d^2) - 2\epsilon^2(E^2 + fd) + \epsilon^4 = 0$$
(5)

With solutions (letting $\Gamma = \sqrt{(f-d)^2 + 4\epsilon^2}$):

$$E_{n\pm} = \pm \frac{1}{2} (\Gamma + (f+d))$$

$$E_{n'\pm} = \pm \frac{1}{2} (\Gamma - (f+d))$$

Neutron in limit $\epsilon \rightarrow 0$
(6)
Mirror neutron in limit $\epsilon \rightarrow 0$

Expected result! Precession energies depend on coupling to mirror sector and mirror magnetic fields!

Experimental signal

$$|a(t_f)|^2 = \left[|b|^2 \sin^2 \frac{\omega_1}{2} t_f - |a| |b| \sin \Delta \phi \sin \omega_1 t_f + |a|^2 \cos^2 \frac{\omega_1}{2} t_f \right]$$
$$= \frac{\alpha^2 M^2}{2} \left[1 - \sin \left(2 \left(\phi_o + \gamma \right) + \left(E_2 - E_1 \right) t_p / \hbar - \pi / 2 \right) \right]$$

The part of the phase that evolves with the total storage time is:

$$E_1 - E_2 = 2(f + d + \sqrt{4\epsilon^2 + (f - d)^2})$$

The effect of coupling to the mirror world is to change the effective size of the magnetic field! Effect is quadratic in couplings away from resonance, like n-n' oscillation

$$E_2 - E_1 \approx -2f\left(1 + \frac{\epsilon^2}{\left(f - d\right)^2}\right),$$

A Precession Experiment with Polarized Neutrons

From B. Franke, "By-products of nEDM Searches", Neutron Summer School 2018, Raleigh NC (2018):



Ingredients to extract f_n via the Ramsey method:

- 100% polarized ensemble
- Magnetic field, ideally on single homogeneous component
- very precise external clock
- count neutrons depending on polarization state

The nEDM search

The Ramsey's method of separated oscillating fields



Spin analyzer (depending on point in precession, spin somewhere between parallel and antiparallel spin analyzer axis) We have considered 3 kinds of experiments so far

- Absolute average change in measured effective magnetic field due to "pseudo-magnetic field" from mirror couplings (expect deviation, especially if B' on scales of few ~10⁻⁶ T where EDM experiments are performed with high precision)
- B-scaling measurements (look for non-linearity in B)
- Time varying fields

All three rely on the presence of **atomic co-magnetometer**, which can be used as a reference, and which is not affected by coupling to the mirror world (checked with Z. Berezhiani, this appears to be reasonable), experiment measures R: $R = \frac{\gamma_n}{\gamma_{Hg}}$

念TRIUMF

Results from a clock comparison



A Limit from Pseudomagnetic Field Constraints

The precession phase is parametrized in terms of energy differences for the energy states:

 $\Delta E = E_2 - E_1 = -\mu_n B_z - (\mu_n B_z) = 2|\mu_n|(B + b) = 2f(1 + \delta)$ Mirror field Spin down Spin up $\delta = \frac{\epsilon^2}{(f - d)^2}$ Given limits on a pseudomagnetic field of b < 5×10⁻¹³ T with B = 1×10⁻⁶ T so f = 6×10⁻¹⁴ eV

ε< 4.2×10⁻¹⁷ eV

For mirror fields significantly larger than 1 μ T, these limits will be less stringent

- Changing magnetic fields is difficult for nEDM experiments – they are carefully optimized for these fields (for PSI they are about 1 µT)
- If several measurements were made of precession at the current precision of nEDMs, then for mirror fields near 1 µT or lower, the limit will be at the uncertainty in determinations of R, equivalent to the relative uncertainty of the precession frequency (or the energy of the precessing state):

Convert EDM limit to effective magnetic splitting ($\Delta E = 2f + 2d_n E$):

 $\Delta E/2f = 2(\sim 1 \times 10^{-26} \text{ ecm} \times 10^4 \text{ V/cm})/(2 \times 6 \times 10^{-14} \text{ eV}) = 1.7 \times 10^{-9}$

$$\epsilon < 2.4 \times 10^{-18} \, eV$$

τ > 260 s

Abel et al., Phys. Rev. X 7, 041034 (2017)



Analysis of precession data (but not organized into EDM sets, just daily variations) could make interesting limits!

τ > ~149 s

- We want to add the transverse fields in the mirror dimension. This has already been done by Berezhiani, but in a basis that confused me – I think that this will result in small changes, including producing a new source of T₁ and T₂ losses (needs to be checked – in some scenarios may produce observable signature for mirror fields
- Precession measurements may add a new tool (if this is right!) to probe interactions with a mirror universe. The sensivitivity seems comparable to beam and UCN disappearance measurements.
 - New observables of precession phase and depolarization, simultaneously accessible
 with oscillations
 - Some data already taken can probably be cast into interesting limits (based on rough estimates)!
- Resonance scenarios may offer more sensitivity, but complicated (beat frequency effects) developing 4x4 matrix approach – C. Swank observed that can scan through resonace and adjust width of resonance independent from field using dressed fields!
- Large range of field orientations, dressed fields, etc...useful for these studies must be planned in advance!

Another concept

- Another opportunity SR forces!
- Very difficult to arrange experiments with different materials without disrupting experiments
- Also difficult to develop very large ratios of near-scale densities, attractor materials
- If precession cell designed appropriately can incorporate a microfluidic substrate which can increase the density ratios of probed materials by 1 to 2 orders of magnitude, and permit relative measurements (on reasonable time scales) without changing the measurement fields, and provides a control for effects!

Tuned to probe ranges $(i. e. V_5 = \alpha \frac{e^{-r/\lambda}}{r})$ near micron scale



Realization

Mercury flowing through array!



Si wafer wet-etch

No opportunity to use!



Assemblies non-magnetic (PSI flux-gate study)

Neutron Friendly

1 micron waviness, much smaller roughness

90% channel filling, 1 micron membrane 5cm x 5cm area, use non-wetting fluids (mercury, ethanol)

Summary

The availability of

- (1) Production capability of radioactive nuclei and neutrons
- (2) Expertise on fundamental symmetries measurments of laser trapped nuclei

Creates some very interesting opportunities (especially if densities near 100 UCN/cm³ are achieved) for next generation asymmetries experiments with reach comparable to superallowed decays and new BSM constraints

Having both ¹⁹Ne and UCNs adds some interesting long term possibilities for other experiments (CRES?), see Brent Graner's talk..

The room temperature EDM might be ideally suited for the flexibility required to take a dvantage of physics opportunities for new limits from precession experiments!

Planning for flexibility in field range, orientation, rf field production and cell configurations with an external vacuum chamber and access for fluidic ports/control may permit new physics reach on dark sector and short-ranged force limits!