Opportunities with 14:00 the ARIEL e-linac 15:00 WG summary



16:00

17:00

		📇 Print	Full screen	Filter
0	Electron beams and new physics	Brian BATEL		
	https://ca01web.zoom.us/j/69998354433 (password: 972986)			14:00 - 1
	Electron beam dump experiments		Dr.	Luca DORI
	https://ca01web.zoom.us/j/69998354433 (password: 972986)			14:30 - 1
0	Discussion			
	https://ca01web.zoom.us/j/69998354433 (password: 972986)			15:00 - 1
	Break			
	https://ca01web.zoom.us/j/69998354433 (password: 972986)			15:30 - 1
0	The DarkLight experiment		Jai	n BERNAUE
	https://ca01web.zoom.us/j/69998354433 (password: 972986)			16:00 - 1
	Electron beams and missing momentum		N	ikita BLINO
	https://ca01web.zoom.us/j/69998354433 (password: 972986)			16:30 - 1
0	Discussion			
	https://ca01web.zoom.us/j/69998354433 (password: 972986)			17:00 - 1

See also plenary talks by O. Kester & J. Dilling

David McKeen DND2020 Nov 6 2020





Electron Beams and New Physics



See also plenary talks by Rajendran & Baryakhtar

Where is the New Physics?

DM in thermal contact at early times has interesting implications

Light DM needs new interactions for this to be viable⇒portals &dark sectors

The Dark Sector Paradigm



Thermal Freezeout

Dark matter produced from reactions in the plasma during the Big Bang Requires *non-gravitational* dark matter interactions



• Viable mass range between MeV - 100 TeV in simplest scenarios



Renormalizable Portals



Vector portal ("dark photon") is a very nice benchmark \Rightarrow

⇐Lots of interest in a handful of "portals"

Benchmark Model: Vector Portal Dark Matter

SM
$$rac{\epsilon}{2}F_{\mu
u}F^{'\mu
u}$$
 χ

[Holdom] [Pospelov, Ritz, Voloshin] [Hooper, Zurek] [Arkani-Hamed, et al]

$$\mathcal{L} \supset |D_{\mu}\chi|^{2} - m_{\chi}^{2}|\chi|^{2} - \frac{1}{4}(F_{\mu\nu}')^{2} + \frac{1}{2}m_{A'}^{2}(A_{\mu}')^{2} - \frac{\epsilon}{2}F_{\mu\nu}'F^{\mu\nu} + \dots$$

- Dark photon mediates interaction between DM and SM
- 4 new parameters: $m_{\chi}, m_{A'}, lpha_D, \epsilon$
- Can obtain correct relic abundance



• Variations in cosmology and phenomenology obtained by changing mediator, or dark matter properties - important to explore all options



Dark Particle Production in Electron Fixed Target Experiments



Advantages:

- high collision luminosity
- forward kinematics
- large production rates
- clean experimental environment



via Bremsstrahlung

Electron beams typically useful for new vectors



[Andreas. et al]

[Izaguirre, Krnjaic, Schuster, Toro]

See talk by

N. Blinov

- Large missing energy signature (small energy deposition in ECAL, no energy deposition in HCAL)
- + 2.84 x 1011 EOT best limits on invisibly decaying dark photon below 300 MeV









Experimental Perspectives on Electron Beam Dumps

DND2020 Workshop

Luca Doria (doria@uni-mainz.de)

PRISMA Cluster of Excellence and Institut für Kernphysik Johannes-Gutenberg Universität Mainz

Prehistory of electron BD experiments



*Beam: 20GeV electrons on Al target *~200m decay length (mostly earth shielding) *Detector: scintillators + wire chambers

*Other experiments (originally for axion searches):

- SLAC E141: 9 GeV electrons on W E. M. Riordan et al., Phys. Rev. Lett. 59, 755 (1987).
- Fermilab E774: 275 GeV electrons on W A. Bross et al., Phys. Rev. Lett. 67, 2942 (1991).





Data reinterpreted theoretically by: Bjorken et al. Phys.Rev.D80:075018 (2009)

Electron Beam Dump Experiments

- Availability of high-current, CW accelerators
- Parasitic operation
- Complementarity wrt proton beam dumps (meson decays)
- Lower neutrino background
- Theoretically simpler signal (similar to QED processes)
- Double test: DM production (in the BD) AND interaction (in the detector)

BDX at JLab



https://www.jlab.org/accel/ops/ops_liaison/BDX/BDX.html







2. EB mode, (un-)polarized, 155 MeV, 150 µA: P2 *High stability, thick targets, long runs, high luminosities* (2.) MX-EB mode: (un-)polarized, 30–105 MeV, 10 µA: MAGIX Early MAGIX measurements, short runs



Phase 1



1000 (available!) PbF2 crystals Volume: 1x1x0.13 m³ 5x5 crystal sub-modules 1200 kg mass

Phase 2



Addition of Pb-Glass blocks Volume: 1m³ 4100 kg mass

Phase 3



Reach maximum volume: O(10m³)

DarkMESA

Beam Dump

- 20 X₀ Beam Dump
- Material: Aluminum (+ Water)
- Addition of a W plate?
- Energy on Dump: ~135 MeV
- 10⁴ h of operation; 10²² EOT

Experimental Area

- 70 X₀ (~8m) barite concrete
- ~ no neutrons at detector position 🏨

Ē

- no beam dump backgrounds
- No neutrinos



Detector Concept:

- 81 lead glass blocks
- 30x30x150cm each
- 5" PMTs or SiPM readout
- Other crystals under study

Background Rejection

- Beam on/off
- Comics Veto
- Segmentation



⇐Staged detector design

23 m



DarkMESA



Simulation

- GEANT4
- Experimental Halls
- Beam Dump
- Detector
- DM/e DM/p interaction
- MadGraph-4
- Dark Photon Production
- Input to GEANT4

Cosmic background sets a lower limit on threshold \Rightarrow

Estimated reach for dark photon DM is interesting



The ARIEL Case



Summary of Strengths and Challenges

Beam Properties:

High power BD (~100kW expected, more w/o ISOL target..500kW?), bremsstrahlung on Au (+Al) Low beam energy (30 MeV -> 50 MeV?) Have to stay close to BD for good acceptance -> backgrounds? Advantage: no muon/neutrino background

Detector:

Calorimeter / Noble liquid detector / Gas TPC ... ? Low DM masses -> Low threshold -> BKGs again (environment, BD, low-E neutrals) Veto system: cosmics, <u>low energy neutrons and photons</u> Timing? Challenging with CW beam (need sub-ns resolution) -> dedicated bunched beam?

Further studies:

Complex logistics: where to place the detector (separator room, new cave, new beamline, ...)? Enough space in the separator room? Radiation levels low enough?

Summary

- LDM is a quite generic possibility. Many models on the market: experiments needed!
- With a rapidly "heavy" DM window closing, "light" DM searches are gaining a lot of interest.
- Dark sector experiments discussed at major labs equipped with electron machines: SLAC, Cornell, DESY, ELSA, MAMI/MESA, Frascati, KEK, ... Lot of competition.
- BD-type experiments have the potential to explore unique parameter regions at low masses.
- An opportunity for the TRIUMF beams (protons could also be an option...?)
- Realistic full simulation study needed: beam dump + detector technology





There has been a lot of attention on experiment at ATOMKI⇒

X17

C Physics - Technology - Community - In focus

SEARCHES FOR NEW PHYSICS | NEWS

9

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Rekindled Atomki anomaly merits



Additional evidence of an unknown particle from a Hungarian lab give: to NA64 searches

27 NOVEMBER, 2019 | By Ana Lopes





Los Angeles Times

Did physicists discover a previously unknown fifth force of nature?

If it's real, it needs to be studied in gory detail. David McKeen, theoretical particle physicist at the University of Washington. If scientists ... Aug 19, 2016

The NA64 experiment at CERN (Image: CERN)

Atomki's new high-resolution LaBr3 spectrometer, which will record g excited nuclei. Credit: Atomki



 0^{+}







New results on ${}^{3}H(p,\gamma){}^{4}He$ arXiv:1910.10459 [nucl-ex]









Why believe it?

- ► This model has $\chi^2/d.o.f.$ of 1.07, significance of 6.8σ
- Bump, not last bin effect
- Remeasured with new detector: A J Krasznahorkay et al 2018 J. Phys.: Conf. Ser.1056 012028
- Compatible masses in ⁸Be and ⁴He, and compatible couplings (Feng et al. arXiv:2006.01151)
- Non-linearities in Isotope shifts (King-plots), observed (I. Counts et al., arXiv:2004.11383)
 - Hard to distinguish from higher order SM effects.



Why not believe it?

DM boson interpretation is proto-phobic to evade NA48/2 limits

- Actually: $\frac{\epsilon_p}{\epsilon_n}$ coupling below $\pm 8\%$. Z^0 is $\sim 7\%$
- Recently, alternative processes were proposed
 - ► arXiv:2003.05722v3 Hard $\gamma + \gamma$ process
 - arXiv:2005.10643 Anomalous Internal Pair creation









Dominated by accidental background



- Scaling with instantaneous luminosity:
 - $\blacktriangleright \text{ Signal } S \sim \mathcal{L}$
 - ▶ QED background $Q \sim \mathcal{L}$
 - Accidental background $A \sim \mathcal{L}^2$
 - Sensitivity $\frac{S}{\sqrt{Q+A}} \propto 1$ for $A \gg Q$
- Sensitivity almost independent of luminosity. Scale is set by bunch-clock / time resolution
- Out-of-time "coincidences" give accurate measure of acceptance including efficiency.





16

2

0

10

12

14



18

Invariant mass [MeV]

20

22

24



Tracking detectors

- Stack of three tGEMs, 25x40 cm, modified CERN design
- Readout via APVs and MPD4 (Same as SBS and PREX)
- Hampton group has built eight.



Experience: Møller at MIT HVRL



Trigger detectors

- Scintillator Hodoscope, 10 segments/spectrometer
- Needs timing resolution of < 500 ps</p>
- MUSE beam hodoscope: 2 mm thick scintillator, SiPM readout: < 100 ps</p>
 - Tested up to 8mm wide, 15 cm long.





Where it could be sited \Rightarrow

⇐ What the experiment would look like



- Minimal modification
- Could use exisiting beam dump
- Cleaner environment: Beam dump far away
- Might be able to recover beam energy



Opportunities for Missing Momentum Experiments

Nikita Blinov

November 5, 2020

DND 2020, TRIUMF



Missing Energy/Momentum

Detect DM indirectly by observing recoiling SM particle. Background



Every Mediator/DM particle produced is detected!

Variations on a Theme



Qualitatively similar targets in a wide variety of other models

Beam Requirements

1) Need to track each incident beam particle

low current

2) High statistics on a ~year time scale (>10¹⁴ EOT)

single/few electrons @>30 MHz repetition rate

Candidate beams:

- S30XL@SLAC SLAC-R-1147; must be parasitic to freeelectron laser program
- CEBAF@JLAB primarily a nuclear physics facility
- eSPS@CERN CERN-SPSC-2018-023 hypothetical

Signal Kinematics



Light Dark Matter eXperiment -DMX

 Detector design developed by the LDMX collaboration, using technology from CMS, Mu2e and HPS experiments

LDMX Collaboration (1808.05219) '18

 Background studies using realistic detector simulation show the design achieves the necessary background rejection for 10¹⁴ EOT LDMX Collaboration (1912.05535) '19



Backgrounds



LDMX Collaboration '19

LDMX Projections



Phase 1: $\sim 10^{14}$ EOT, 4 GeV e Beam Phase 2: $\sim 10^{16}$ EOT, 8 GeV e Beam

LDMX+Belle II can decisively test thermal DM below a GeV!

Nikita Blinov performed quick ARIEL beam study for LDMXtype setup!

ARIEL Beam

Much of previous discussion translates to ~50 MeV electron beam





Possible Reach

50 MeV electron beam, 10^{16} EOT on $0.1X_0$ Tungsten



- Challenge 1: nominal ARIEL current probably too high (pileup) Challenge 2: lower energy, more wide-angle/lost emissions (background)? Viable Parameter Space
 - Challenge 3: Large range of accessible parameter space in tension with cosmology

Nontrivial reach but some area in tension with cosmology \Rightarrow



Prompts from the Organizing Team What ideas could (or could not) be turned into **actual experiments** at TRIUMF/CENPA?

What homework do we need to do to figure this out?

Should we form a **working group** to answer such questions? Who are key people to target for this working group?

Prompts from the Organizing Team

experiments at TRIUMF/CENPA? also interesting. What about other beams, e.g. p?

Should we form a working group to answer such group?

- What ideas could (or could not) be turned into actual
- The DarkLight concept looks really exciting. Beam dump What **homework** do we need to do to figure this out?
- questions? Who are key people to target for this working

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- What ideas could (or could not) be turned into actual
- The DarkLight concept looks really exciting. Beam dump What **homework** do we need to do to figure this out? Further studies and simulations crucial to firm up reach questions? Who are key people to target for this working

Prompts from the Organizing Team

What ideas could (or could not) be turned into actual experiments at TRIUMF/CENPA? The DarkLight concept looks really exciting. Beam dump also interesting. What about other beams, e.g. p? What **homework** do we need to do to figure this out? Further studies and simulations crucial to firm up reach Should we form a working group to answer such questions? Who are key people to target for this working group? Absolutely! DarkLight collaborators & theorists (and others?)